

Gamow-Teller strength distributions at finite temperature and electron capture in stellar environment

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The thermal evolutions of the GT_+ distributions are presented for the sample nuclei ⁵⁴Fe and ⁷⁶Ge. They are obtained within the proton-neutron QRPA extended to finite temperature by the thermofield dynamics formalism. The strength distributions are used to calculate stellar electron capture rates and cross sections.

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1. Introduction

It is well known that reliable electron capture rates are essential for simulations of corecollapse supernovae and their calculation requires a detailed knowledge of the Gamow-Teller strength distributions in nuclei [1]. Furthermore, the finite temperature in the star requires the implicit consideration of capture on excited nuclear states, for which the GT_+ distribution can be very different from the one for the ground states.

To date, the most reliable calculations of electron capture rates have been performed for iron group nuclei using a large-scale shell model (LSSM) diagonalization approach [2]. For more massive and neutron-rich nuclei electron capture rates have been derived within the framework of a hybrid model combining the Shell Model Monte-Carlo (SMMC) approach and the Random Phase Approximation (RPA) [4]. Nevertheless, both approaches have their own shortcomings: The LSSM partially employs the Brink hypothesis when treating GT transitions from nuclear excited states. The hybrid model is free of this disadvantage, but it does not include explicitly pairing correlations when calculating the strength distributions. An alternative method for description of Gamow-Teller strength distributions and electron capture rates at finite temperatures is presented in this talk.

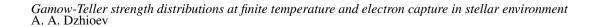
2. Approach

In Ref. [5], we have introduced a theory based on the proton-neutron QRPA extended to finite temperature (TQRPA) by the thermofield dynamics formalism (TFD) [6]. This technique does not rely on Brink's hypothesis and allows to calculate the strength and the energies of allowed and forbidden transitions as functions of the nuclear temperature. Moreover, the Ikeda sum rule for Fermi and GT transitions is fulfilled within the approach. In our study we use a phenomenological nuclear Hamiltonian of the quasiparticle-phonon nuclear model (QPM) [7]. It consists of a spherically symmetric Woods-Saxon mean field potential for protons and neutrons, BCS pairing interactions and separable multipole and spin-multipole particle-hole interactions. All parameters in the QPM Hamiltonian are fitted to reproduce experimental data at zero temperature (see Ref. [5] for more details).

3. The results

The temperature evolution of the GT_+ strength distribution for ⁵⁴Fe is displayed in Fig. 1(lhs). With increasing temperature two effects occur in our model that influence the GT_+ strength distribution: (i) At low temperatures, due to pairing, GT_+ transitions involve the breaking of a proton Cooper pair associated with some energy cost. This extra energy vanishes at temperatures higher then critical one ($T_{cr} \approx 0.8$ MeV). (ii) GT_+ transitions, which are Pauli blocked at low temperatures due to closed neutron subshells (e.g., $1f_{7/2}$ orbital), become thermally unblocked as temperature increases. Similarly, protons which are thermally excited to higher orbitals can undergo GT_+ transitions. Because of thermally unblocked transitions, some GT_+ strength appears well below the zero-temperature threshold, including negative energies.

Due to the vanishing of the pairing correlations and appearance of negative- and low-energy transitions, the centroid of the GT_+ strength distribution in ⁵⁴Fe is shifted to lower excitation energies at high temperatures. Our calculations indicate that as temperature increases to 0.8 MeV,



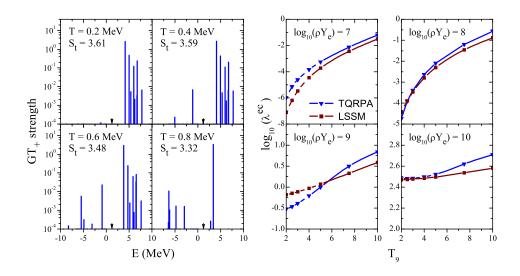
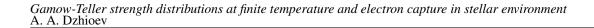


Figure 1: (Left panel) Temperature evolution of GT_+ strength distributions for ⁵⁴Fe versus parent excitation energy. S_t is the total GT_+ strength. The arrows indicate the zero-temperature threshold $Q = M_f - M_i =$ 1.21 MeV, where $M_{i,f}$ are the masses of the parent and daughter nuclei. (Right panel) Electron capture rates for ⁵⁴Fe calculated using the TQRPA and LSSM approaches as a function of temperature ($T_9 = 10^9$ K) for selected values of density ρY_e (in g cm⁻³).

the GT_+ centroid shifts by 1.5 MeV. Thus, the present approach does not support Brink's hypothesis. Similar results have been obtained in SMMC calculations of the GT centroids at finite temperatures [8]. A gradual decrease of the total GT_+ strength S_t takes place when the temperature increases from zero to 0.8 MeV (see S_t values in Fig. 1(lhs)).

In Fig. 1(rhs), we compare electron capture rates calculated in the TQRPA with the ones calculated using the LSSM approach [3]. There is a certain disagreement between them. At low temperatures (*T*) and densities (ρ) the disagreement is due to difference in the strength and the energy of negative-energy transitions which dominate the rates at low (*T*, ρ). At low *T* and intermediate ρ , the near threshold part of the GT₊ strength dominates the capture rates. The TQRPA accounts for the Landau damping only and thus, underestimates fragmentation of the GT₊ strength as compared to the LSSM, leading to too small strength at the threshold. Accordingly our rates appear to be smaller than the LSSM ones. Accounting for coupling of one-phonon states to complex configurations in our model should make predictions of two approaches closer. At high temperatures the rates are always dominated by strong transitions involving the GT₊ resonance. Since the present approach predicts that with increasing temperatures the GT₊ resonance moves to lower excitation energies, our high temperature rates always slightly surpass the shell-model ones.

The strength distribution of GT_+ transitions from the even germanium isotope ⁷⁶Ge is displayed in Fig. 2(lhs). The distribution has been folded with a Breit-Wigner function of 1 MeV width. With increasing temperature, the peaks in the GT_+ distribution shift to lower excitation energies and the total strength decreases in the vicinity of the critical temperature ($T_{cr} \sim 0.9$ MeV). The shift is about 8 MeV and, hence, cannot be explained solely by disappearing of the extra energy needed to break a proton pair. To explain both the effects we consider single-particle transitions which mainly contribute to the total GT_+ strength in ⁷⁶Ge. These are the $1g_{9/2}^p \rightarrow 1g_{7/2}^n$ particle-



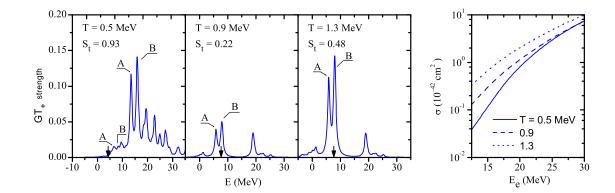


Figure 2: (Left panel) Strength distribution (folded) of GT_+ transitions in ⁷⁶Ge at various temperatures. The arrows indicate the zero-temperature threshold Q = 7.52 MeV. A and B label the transitions: $A \equiv 1f_{7/2}^p \rightarrow 1f_{5/2}^n$, $B \equiv 1g_{9/2}^p \rightarrow 1g_{7/2}^n$. (Right panel) Electron capture cross sections for ⁷⁶Ge calculated within the TQRPA approach for various temperatures. E_e is the electron energy.

particle and $1f_{7/2}^p \to 1f_{5/2}^n$ hole-hole transitions. Both transitions are blocked at zero temperature in an independent particle model. At relatively low temperatures, configuration mixing induced by pairing correlations in the ground state is the main unblocking mechanism. The position of the GT₊ peaks is given by $\varepsilon_{1g_{9/2}}^p + \varepsilon_{1g_{7/2}}^n + \delta_{np}$ and $\varepsilon_{1f_{7/2}}^p + \varepsilon_{1f_{5/2}}^n + \delta_{np}$, where $\varepsilon_j^{p(n)}$ is the proton (neutron) quasiparticle energy. The quantity $\delta_{np} = \Delta \mu_{np} + \Delta M_{np}$ takes into account the difference between the proton and the neutron chemical potentials and the proton-neutron mass splitting. As a result, at low temperatures the GT₊ peaks resides at energies around 15 MeV. When pairing correlations disappear at temperatures near the critical one, the corresponding peaks completely vanish. The total GT₊ strength is noticeably reduced. At higher temperatures, the above mentioned single transitions become unblocked due to thermal excitations [9]. But now their energies are determined by the difference of quasiparticle energies, $\varepsilon_{1g_{7/2}}^n - \varepsilon_{1g_{9/2}}^p + \delta_{np}$ and $\varepsilon_{1f_{7/2}}^p - \varepsilon_{1f_{5/2}}^n + \delta_{np}$. As a result the GT₊ peaks appear near the zero-temperature threshold. Obviously, the significant shift of the GT₊ peaks to lower excitation energies favors electron capture.

No shift to lower excitation energies was observed within the hybrid model calculations [4]. In the hybrid model, occupation numbers at finite temperature are calculated within the SMMC approach, accounting for all many-body *npn*h correlations induced by a pairing+quadrupole residual interaction. These occupation numbers have been then used to define a thermal ground state that is the basis of an RPA approach to calculate the electron capture cross sections, considering only 1p1h excitations on the top this ground state. Therefore the hybrid model does not include explicitly 2p2h pairing correlations when calculating strength distributions. As a consequence the energy of GT_+ transitions is nearly independent on the temperature within the hybrid model and GT_+ peaks are always located near the zero-temperature threshold.

To reveal the importance of the thermal unblocking for GT_+ transitions in neutron-rich nuclei we perform electron capture cross section calculations (see Fig. 2(rhs)) taking into account the contribution of GT_+ and first-forbidden transitions in our calculations. The strong temperature sensitivity of the cross section at low electron energies reflects the temperature dependence of the TQRPA GT_+ strength distribution in our model. It is amplified by the strong phase space energy dependence, leading to a much stronger temperature dependence of the cross section in the TQRPA model than in the hybrid model [4]. For higher electron energies, the first-forbidden transitions become increasingly important [9]. As the strength of the first-forbidden transitions is less sensitive to temperature [5], the capture cross sections at $E_e \sim 30$ MeV depend only weakly on temperature.

4. Conclusion and outlook.

We have presented the approach that allows calculations of stellar weak-interaction processes at finite temperature in a thermodynamically consistent way. It accounts for correlations described by the TQRPA. Our approach is conceptually superior to the hybrid approach of the SMMC+RPA which has previously been used to estimate electron capture rates for neutron-rich nuclei. Whereas much of the essential physics is already recovered, the detailed comparison to the shell model results implies that the approach should be further improved. The improvement would be inclusion of correlations beyond TQRPA by coupling one-phonon states to more complex (e.g., two-phonon) configurations [10, 11]. For charge-exchange excitations in cold nuclei this problem has been considered within the QPM [12] and by approaches that solve the (second) RPA equations in the space of two-particle/two-hole excitations [13]. Another direction is the combination of our TFD based approach at finite temperature with self-consistent QRPA calculations based on more realistic effective interaction [14].

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