

# Soft corrections to inclusive DIS at four loops and beyond

#### Goutam Das\*

Theory Group, Deutsches Elektronen-Synchrotron (DESY), D-22607 Hamburg, Germany E-mail: goutam.das@desy.de

#### **Sven-Olaf Moch**

II. Institute for Theoretical Physics, Hamburg University, D-22761 Hamburg, Germany E-mail: sven-olaf.moch@desy.de

## **Andreas Vogt**

Department of Mathematical Sciences, University of Liverpool, Liverpool L69 3BX, United Kingdom E-mail: andreas.vogt@liverpool.ac.uk

We study the threshold corrections to the structure functions in deep-inelastic scattering (DIS) at the fifth logarithmic (N<sup>4</sup>LL) order of the soft-gluon exponentiation in massless perturbative QCD. Using recent results for the splitting functions and the quark form factor, we derive the fourth-order contribution to the coefficient  $f^q$  of the form factor and from it the N<sup>4</sup>LL part of the exponentiation coefficient  $B^{DIS}$  in the limit of a large number of colours. An approximation scheme is shown that leads to sufficiently accurate N<sup>4</sup>LL results for full QCD. The N<sup>4</sup>LL corrections are small and lead to a further stabilization of the perturbative expansion for the soft-gluon exponent.

XXVII International Workshop on Deep-Inelastic Scattering and Related Subjects - DIS2019 8-12 April, 2019, Torino, Italy

<sup>\*</sup>Speaker.

#### 1. Introduction

The Wilson coefficients (coefficient functions) for the structure functions of inclusive DIS have been a subject of research since the early days of QCD. These quantities are not only relevant for determining the parton distribution functions (PDFs) and the strong coupling constant  $\alpha_s$  using structure function data, see, e.g., [1], but also to other processes and less inclusive observables in DIS, see, e.g., [2]. The main Wilson coefficients for DIS are presently known up to the third order in  $\alpha_s$  in massless perturbative QCD [3]. Their perturbative expansion is well-behaved except close to the kinematic endpoints x = 0 and x = 1 of the Bjorken variable. The dominant terms  $\ln^{\ell}(1-x)/(1-x)_+$  in the latter (threshold) limit are resummed by the soft-gluon exponentiation, see, e.g., [4–6], which is best formulated in Mellin *N*-space [4]. So far this resummation has been performed up to the next-to-next-to-next-to-leading logarithmic (N<sup>3</sup>LL) accuracy [7].

The threshold resummation coefficients are closely related to the large-x limit of the quark-quark splitting functions for the PDFs and to the quark form factor [8,9] which are both fully known to order  $\alpha_s^3$  [10–13]. Recently the computations of these quantities have been extended to order  $\alpha_s^4$  in the (L $n_c$ ) limit of a large number of colours [14, 15]. Together with approximate results for the  $n_f$ -independent [16] and exact expression for the  $n_f$ -dependent contributions to the cusp anomalous dimension in full QCD [17] these results facilitate the effective extension of the threshold resummation for the DIS Wilson coefficients to the next (N<sup>4</sup>LL) logarithmic order. In the following we recall the theoretical framework, present the N<sup>4</sup>LL resummation coefficient and briefly address the numerical implications of this result for the resummation of DIS in QCD.

### 2. Theoretical framework and new fourth-order coefficients

The all-order large-N behaviour of the DIS Wilson coefficients for  $F_1$ ,  $F_2$  and  $F_3$  can be written as

$$C^{N}(Q^{2}) = g_{0}(Q^{2}) \cdot \exp[G^{N}(Q^{2})] + \mathcal{O}(N^{-1}\ln^{n}N), \qquad (2.1)$$

where the resummation exponent  $G^N$  of the dominant  $N^0 \ln^n N$  contributions is given by [18]

$$G^{N} = \int_{0}^{1} dz \, \frac{z^{N-1} - 1}{1 - z} \left[ \int_{\mu_{f}^{2}}^{(1-z)Q^{2}} \frac{dq^{2}}{q^{2}} A^{q} \left( \alpha_{s}(q^{2}) \right) + B^{\text{DIS}} \left( \alpha_{s}((1-z)q^{2}) \right) \right]. \tag{2.2}$$

Here  $A^q$  is the (light-like) quark cusp anomalous dimension and  $B^{DIS}$  is the resummation coefficient for DIS. Both have perturbative series,  $A^q = \sum_i a_s^i A_i^q$  etc, in terms of the strong coupling which we normalize as  $a_s \equiv \alpha_s/4\pi$ . Performing the integrations one can organize the exponent as

$$G^{N} = \ln N g^{(1)}(\lambda) + g^{(2)}(\lambda) + a_{s}g^{(3)}(\lambda) + a_{s}^{2}g^{(4)}(\lambda) + a_{s}^{3}g^{(5)}(\lambda) + \dots,$$
 (2.3)

where  $\lambda = \beta_0 a_s \ln N$  or  $\lambda = \beta_0 a_s \ln \widetilde{N}$  with  $\ln \widetilde{N} = \ln N + \gamma_e$ . The first n+1 terms in (2.3) are required for the resummation at N<sup>n</sup>LL accuracy. The N<sup>2</sup>LL and N<sup>3</sup>LL contributions to  $G^N$  have been derived in [7, 19]; explicit expressions can be found in (3.3) – (3.6) of [7]. The lengthy new function  $g^{(5)}(\lambda)$  entering at N<sup>4</sup>LL will be presented in [20]. The N-independent prefactor  $g_0$  is presently known to order  $\alpha_s^3$  from the all-N calculation in [3], see (4.6) – (4.8) of [7],

$$g_0 = 1 + a_s g_{01} + a_s^2 g_{02} + a_s^3 g_{03} + \mathcal{O}(a_s^4). \tag{2.4}$$

The resummation to N<sup>4</sup>LL requires the terms up to  $A_5^q$  and  $B_4^{DIS}$  in their corresponding expansions. The impact of the former quantity, for which a first estimate has been obtained in [21], is very small.  $B^{DIS}$  can be calculated from knowledge of the quark form factor or the DIS Wilson coefficients. The form factor satisfies a differential equation which follows from the renormalization group and gauge invariance. Its solution can be found in terms of the cusp anomalous dimension  $A^q$  and the function  $G^q$  containing the quantity  $f^q$  related to a universal eikonal anomalous dimension and the coefficient  $B^q$  of  $\delta(1-x)$  in the quark-quark splitting function. The four-loop coefficient of  $G^q$  (which appears in the  $1/\varepsilon$  coefficient in the solution of the form factor) can be written as

$$G_4^{q} = 2B_4^{q} + f_4^{q} + \beta_2 f_{01}^{q} + \beta_1 f_{02}^{q} + \beta_0 f_{03}^{q} + \mathcal{O}(\varepsilon), \qquad (2.5)$$

where the quantities  $f_{0n}^{q}$  are (combinations of) known lower-order coefficients of  $G^{q}$ , see [12] and (20) of [9]. Hence  $f_{4}^{q}$  can be determined in the large- $n_{c}$  limit from the results of [14, 15]. We find

$$\begin{split} f_4^{\mathsf{q}} \Big|_{\mathsf{L}n_c} &= C_F n_c^3 \left( \frac{9364079}{6561} - \frac{1186735}{729} \zeta_2 - \frac{837988}{243} \zeta_3 + \frac{115801}{27} \zeta_4 + \frac{11896}{9} \zeta_2 \zeta_3 + 3952 \zeta_5 \right. \\ &\quad \left. - \frac{4796}{9} \zeta_3^2 - \frac{129547}{54} \zeta_6 - 416 \zeta_2 \zeta_5 - 720 \zeta_3 \zeta_4 - 1700 \zeta_7 \right) + C_F n_c^2 n_f \left( -\frac{247315}{432} \right. \\ &\quad \left. + \frac{412232}{729} \zeta_2 + \frac{102205}{243} \zeta_3 - \frac{7589}{6} \zeta_4 - \frac{824}{9} \zeta_2 \zeta_3 - \frac{740}{9} \zeta_5 + \frac{2816}{9} \zeta_3^2 + \frac{15611}{27} \zeta_6 \right) \\ &\quad \left. + C_F n_c n_f^2 \left( \frac{329069}{17496} - \frac{22447}{729} \zeta_2 + \frac{25300}{243} \zeta_3 + \frac{140}{3} \zeta_4 - \frac{176}{9} \zeta_2 \zeta_3 - \frac{856}{9} \zeta_5 \right) \right. \\ &\quad \left. + C_F n_f^3 \left( -\frac{16160}{6561} - \frac{16}{81} \zeta_2 - \frac{400}{243} \zeta_3 + \frac{128}{27} \zeta_4 \right). \end{split} \tag{2.6}$$

The  $a_s^4$  contribution to resummation  $B^{DIS}$  reads, in terms of the genuine four-loop contributions  $f_4^q$  and  $B_4^q$ , which are exactly known only in the  $Ln_c$  limit for now, and lower-order coefficients,

$$B_{4}^{\text{DIS}} = -f_{4}^{\text{q}} - B_{4}^{\text{q}} - \beta_{2} \left( f_{01}^{\text{q}} + g_{01} - \frac{1}{2} \zeta_{2} A_{1}^{\text{q}} \right) + \beta_{0}^{3} \left( 3\zeta_{2} f_{01}^{\text{q}} + 3\zeta_{2} g_{01} + 2\zeta_{3} f_{1}^{\text{q}} + 2\zeta_{3} B_{1}^{\text{q}} \right)$$

$$+ \frac{3}{2} \zeta_{4} A_{1}^{\text{q}} - \frac{3}{4} \zeta_{2}^{2} A_{1}^{\text{q}} \right) + \beta_{0} \beta_{1} \left( \frac{5}{2} \zeta_{2} f_{1}^{\text{q}} + \frac{5}{2} \zeta_{2} B_{1}^{\text{q}} + \frac{5}{3} \zeta_{3} A_{1}^{\text{q}} \right) + \beta_{0}^{2} \left( 3\zeta_{2} f_{2}^{\text{q}} + 3\zeta_{2} B_{2}^{\text{q}} + 2\zeta_{3} A_{2}^{\text{q}} \right)$$

$$- \beta_{1} \left( f_{02}^{\text{q}} + 2g_{02} - (g_{01})^{2} - \zeta_{2} A_{2}^{\text{q}} \right) - \beta_{0} \left( f_{03}^{\text{q}} + 3g_{03} - 3g_{02}g_{01} - (g_{01})^{3} - \frac{3}{2} \zeta_{2} A_{3}^{\text{q}} \right) , \quad (2.7)$$

where  $g_{0i}$  are to be taken without the  $\gamma_e$  terms in (4.6) – (4.8) of [7]. Its explicit form is given by

$$\begin{split} B_4^{\text{DIS}}\Big|_{\text{L}n_c} &= C_F n_c^3 \left( -\frac{2040092429}{139968} + \frac{23011973}{1944} \zeta_2 + \frac{517537}{36} \zeta_3 - \frac{312481}{36} \zeta_4 - \frac{39838}{9} \zeta_2 \zeta_3 \right. \\ & \left. -\frac{50680}{9} \zeta_5 - 988 \zeta_3^2 + \frac{12467}{6} \zeta_6 + 496 \zeta_2 \zeta_5 + 688 \zeta_3 \zeta_4 + 2260 \zeta_7 \right) \\ & + C_F n_c^2 n_f \left( \frac{83655179}{11664} - \frac{5160215}{972} \zeta_2 - \frac{639191}{162} \zeta_3 + \frac{24856}{9} \zeta_4 + \frac{8624}{9} \zeta_2 \zeta_3 \right. \\ & \left. + 200 \zeta_5 - 32 \zeta_3^2 - \frac{1201}{3} \zeta_6 \right) + C_F n_f^3 \left( \frac{50558}{2187} + \frac{80}{81} \zeta_3 - \frac{1880}{81} \zeta_2 + \frac{40}{9} \zeta_4 \right) \\ & + C_F n_c n_f^2 \left( -\frac{5070943}{5832} + \frac{160903}{243} \zeta_2 + \frac{14618}{81} \zeta_3 - \frac{2110}{9} \zeta_4 - \frac{400}{9} \zeta_2 \zeta_3 + \frac{904}{9} \zeta_5 \right). \end{split}$$

# 3. Numerical implications

The lower-order coefficients  $B_l^{\rm DIS}$  have the same structure as (2.7), i.e., they contain  $-f_l^{\rm q} - B_l^{\rm q}$  and lower-order coefficients. Therefore, by comparing the exact results to an approximation at N<sup>I</sup>LL in which the Ln<sub>c</sub> expression for  $-f_l^{\rm q} - B_l^{\rm q}$  is used together with the exact lower-order coefficients, we can check whether (2.7) with the Ln<sub>c</sub> results for  $f_4^{\rm q}$  and  $B_4^{\rm q}$  can be expected to provide a good approximation for  $B_4^{\rm DIS}$  and hence  $G^N$  at the N<sup>4</sup>LL accuracy of full QCD.

This comparison is carried out in Fig. 1 for l=2 and l=3 (at l=1 there is no difference between the  $Ln_c$  limit and full QCD). The  $Ln_c$  curves are off by less than 0.5% at N<sup>2</sup>LL and 0.25% at N<sup>3</sup>LL for  $G^{DIS}$  in the N-range shown and, at  $x \le 0.9$ , for the convolution of its exponential with a schematic but sufficiently realistic form for a quark PDF. Therefore we can safely expect that the  $Ln_c$  numbers will deviate from (presumably exceed) the exact QCD results by well below 1%.

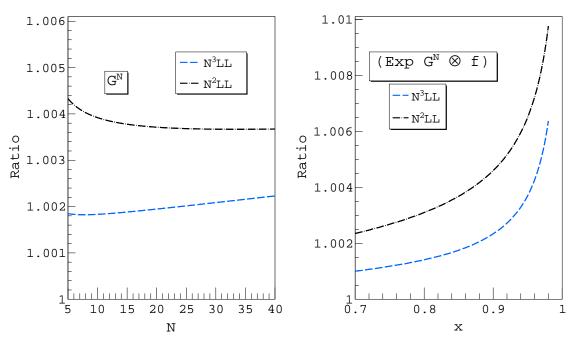


Figure 1: The ratio of the large- $n_c$  approximation, defined as above, and the exact results at N<sup>2</sup>LL and N<sup>3</sup>LL for the DIS resummation exponent  $G^N$  (left) and for the convolution of the exponential with the schematic quark PDF shape  $xf = x^{0.5}(1-x)^3$  (right) for  $\alpha_s = 0.2$  and  $n_f = 3$  flavours.

The cumulative effect, relative to the NLL results, of the exact N<sup>2</sup>LL and N<sup>3</sup>LL contributions and our new N<sup>4</sup>LL corrections, as above determined using the L $n_c$  limit of  $-f_l^q - B_l^q$  in (2.7), is illustrated in Fig. 2. Unlike the N<sup>3</sup>LL contribution, the N<sup>4</sup>LL correction is almost negligible at  $N \le 15$  and  $x \le 0.9$ . Even at N = 40, the functions  $g^{(n)}(\lambda)$  add only 6%, 1.6% and 1% to the NLL result, respectively, for n = 2, n = 3 and n = 4, where the latter L $n_c$  result is presumably a slight overestimate. The corresponding N<sup>2</sup>LL, N<sup>3</sup>LL and N<sup>4</sup>LL percentages for the convolution of  $\exp G^N$  with  $xf = x^{0.5}(1-x)^3$  at x = 0.95 read 9.5%, 1.5% and 0.5%, where we have performed the Mellin inversion using a standard contour, see, e.g., [22], which constitutes a 'minimal prescription' contour [4] in the context of the present exponentiation. It appears that the expansion of  $G^N$  to N<sup>4</sup>LL for the structure functions in inclusive DIS is sufficient for all practical purposes.

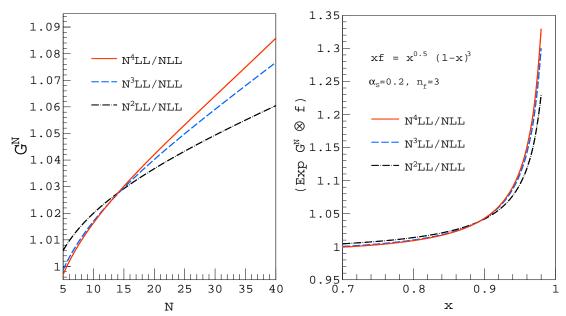


Figure 2: Left: The DIS resummation exponent  $G^N$  in (2.3) up to N<sup>4</sup>LL accuracy, normalized to the NLL result at the standard reference point  $\alpha_s = 0.2$  for  $n_f = 3$ . Right: corresponding x-space results for  $\exp G^N$  after convoluted with a schematic form of a quark PDF of the proton.

# 4. Summary and outlook

We have studied the soft-gluon exponentiation (SGE) of inclusive DIS at the fifth logarithmic (N<sup>4</sup>LL) order. Recent four-loop results on splitting functions and the quark form factor [14, 15] facilitate the exact determination of the form-factor coefficient  $f^q$  and the SGE coefficient  $B^{\text{DIS}}$  at order  $\alpha_s^4$  in the large- $n_c$  (L $n_c$ ) limit. Both coefficients are relevant beyond the context of DIS: Like the lightlike quark and gluon cusp anomalous dimensions  $A^{q,g}$  [10], the quantities  $f^{q,g}$  are maximally non-Abelian and related by a simple Casimir scaling up to three loops. We expect that the generalized Casimir scaling of [16] also applies to  $f^{q,g}$ , hence our result (2.6) fixes also  $f^g$  at large  $n_c$ . The coefficient here called  $B^{\text{DIS}}$  is due to the outgoing unobserved quark; hence it contributes to the SGE for many other processes including, e.g., direct photon production [6].

The L $n_c$  approximation to the N<sup>4</sup>LL resummation exponent  $G^N$  for inclusive DIS, defined as discussed above, is sufficiently accurate to demonstrate that the N<sup>4</sup>LL corrections are small: they contribute well below 1% over a wide range in N and x. As shown in [23], the  $1/N \ln^{\ell} N$  non-SGE contributions are larger; the highest four of these logarithms are currently known to all orders [23,24] – recall the parameter  $\xi_{\text{DIS}_4}$  unspecified in [23] was fixed in [24]. We have considered the case of  $n_f = 3$  light flavours. In electromagnetic and neutral-current DIS, also charm production close to threshold needs to be taken into account beyond the threshold for  $c\bar{c}$  production, see [25].

## References

- [1] A. Accardi et al., Eur. Phys. J. C76 (2016) 471, arXiv:1603.08906.
- [2] P. Bolzoni, F. Maltoni, S. Moch and M. Zaro, Phys. Rev. Lett. 105 (2010) 011801, arXiv:1003.4451,M. Cacciari et al., Phys. Rev. Lett. 115 (2015) 082002, arXiv:1506.02660,

- F.A. Dreyer and A. Karlberg, Phys. Rev. Lett. 117 (2016) 072001, arXiv:1606.00840, J. Currie et al., JHEP 05 (2018) 209, 1803.09973.
- [3] J.A.M. Vermaseren, A. Vogt and S. Moch, Nucl. Phys. B724 (2005) 3, hep-ph/0504242,S. Moch, J.A.M. Vermaseren and A. Vogt, Nucl. Phys. B813 (2009) 220, arXiv:0812.4168.
- [4] S. Catani, M.L. Mangano, P. Nason and L. Trentadue, Nucl. Phys. B478 (1996) 273, hep-ph/9604351.
- [5] H. Contopanagos, E. Laenen and G.F. Sterman, Nucl. Phys. B484 (1997) 303, hep-ph/9604313.
- [6] S. Catani, M.L. Mangano and P. Nason, JHEP 07 (1998) 024, hep-ph/9806484.
- [7] S. Moch, J.A.M. Vermaseren and A. Vogt, Nucl. Phys. B726 (2005) 317, hep-ph/0506288.
- [8] S. Moch and A. Vogt, Phys. Lett. B631 (2005) 48, hep-ph/0508265,E. Laenen and L. Magnea, Phys. Lett. B632 (2006) 270, hep-ph/0508284.
- [9] V. Ravindran, Nucl. Phys. B752 (2006) 173, hep-ph/0603041.
- [10] S. Moch, J.A.M. Vermaseren and A. Vogt, Nucl. Phys. B688 (2004) 101, hep-ph/0403192,
   A. Vogt, S. Moch and J.A.M. Vermaseren, Nucl. Phys. B691 (2004) 129, hep-ph/0404111
- [11] S. Moch, J.A.M. Vermaseren and A. Vogt, JHEP 08 (2005) 049, hep-ph/0507039.
- [12] S. Moch, J.A.M. Vermaseren and A. Vogt, Phys. Lett. B625 (2005) 245, hep-ph/0508055.
- [13] P.A. Baikov et al., Phys. Rev. Lett. 102 (2009) 212002, arXiv:0902.3519,
  R.N. Lee, A.V. Smirnov and V.A. Smirnov, JHEP 04 (2010) 020, arXiv:1001.2887,
  T. Gehrmann et al., JHEP 06 (2010) 094, arXiv:1004.3653,
  T. Gehrmann et al., JHEP 11 (2010) 102, arXiv:1010.4478.
- [14] J. Henn et al., JHEP 03 (2017) 139, arXiv:1612.04389.
- [15] S. Moch et al., JHEP 10 (2017) 041, arXiv:1707.08315.
- [16] S. Moch et al., Phys. Lett. B782 (2018) 627, arXiv:1805.09638.
- [17] J.M. Henn, A.V. Smirnov, V.A. Smirnov and M. Steinhauser, JHEP 05 (2016) 066, arXiv:1604.03126, A. Grozin, PoS LL2016 (2016) 053, arXiv:1605.03886,
  J. Davies et al., Nucl. Phys. B915 (2017) 335, arXiv:1610.07477,
  R.N. Lee, A.V. Smirnov, V.A. Smirnov and M. Steinhauser, Phys. Rev. D96 (2017) 014008, arXiv:1705.06862, A. Grozin, JHEP 06 (2018) 073, arXiv:1805.05050,
  R.N. Lee, A.V. Smirnov, V.A. Smirnov and M. Steinhauser, JHEP 02 (2019) 172, arXiv:1901.02898, J.M. Henn, T. Peraro, M. Stahlhofen and P. Wasser, (2019), arXiv:1901.03693.
- [18] S. Moch, J.A.M. Vermaseren and A. Vogt, Nucl. Phys. B646 (2002) 181, hep-ph/0209100,
  S. Forte and G. Ridolfi, Nucl. Phys. B650 (2003) 229, hep-ph/0209154,
  E. Gardi and R.G. Roberts, Nucl. Phys. B653 (2003) 227, hep-ph/0210429.
- [19] A. Vogt, Phys. Lett. B497 (2001) 228, hep-ph/0010146,S. Catani, D. de Florian, M. Grazzini and P. Nason, JHEP 07 (2003) 028, hep-ph/0306211.
- [20] G. Das, S. Moch and A. Vogt, DESY 19-088, LTH 1205, to appear.
- [21] F. Herzog et al., Phys. Lett. B790 (2019) 436, 1812.11818.
- [22] A. Vogt, Comput. Phys. Commun. 170 (2005) 65, hep-ph/0408244.
- [23] S. Moch and A. Vogt, JHEP 11 (2009) 099, arXiv:0909.2124.
- [24] G.Grunberg, Phys. Lett. B687 (2010) 405, arXiv:0911.4471v5,
   A.A. Almasy, G. Soar and A. Vogt, JHEP 03 (2011) 030, arXiv:1012.3352.
- [25] H. Kawamura, N.A. Lo Presti, S. Moch and A. Vogt, Nucl. Phys. B864 (2012) 399, arXiv:1205.5727