

Machine Learning Approach for Air Shower Recognition in EUSO-SPB Data

**Michal Vrabel^a, Ján Genčí^a, Pavol Bobik^b, Francesca Bisconti^{*c},
for the JEM-EUSO Collaboration[†]**

^a Faculty of Electrical Engineering and Informatics, Technical University of Košice, Slovakia

^b Institute of Experimental Physics, Slovak Academy of Sciences, Slovakia

^c INFN, Section of Turin, Italy

E-mail: michal.vrabel@tuke.sk, jan.genci@tuke.sk, bobik@saske.sk,
fbiscont@to.infn.it

The main goal of The Extreme Universe Space Observatory on a Super Pressure Balloon (EUSO-SPB1) was to observe from above extensive air showers caused by ultra-high energy cosmic rays. EUSO-SPB1 uses a fluorescence detector that observes the atmosphere in a nadir observation mode from a near space altitude. During the 12-day flight, an onboard first level trigger detected more than 175 000 candidate events. This paper presents an approach to recognize air showers in this dataset. The approach uses a feature extraction method to create a simpler representation of an event and then it uses established machine learning techniques to classify data into at least two classes - shower and noise. The machine learning models are trained on a set of air shower simulations put on top of the background observed during the flight and a set of events from the flight. We present the efficiency of the method on datasets of simulated events. The flight data events are also used in unsupervised learning methods to identify groups of events with similar features. The presented methods allow us to shorten the candidate events list and, thanks to the groups of similar events identified by the unsupervised methods, the classification of the triggered events is made simpler.

*36th International Cosmic Ray Conference -ICRC2019-
July 24th - August 1st, 2019
Madison, WI, U.S.A.*

*Speaker.

†for collaboration list see PoS(ICRC2019)1177

1. Introduction

EUSO-SPB1 experiment is a path-finder mission within the Joined Experiment Missions for the Extreme Universe Space Observatory (JEM-EUSO). EUSO-SPB1 uses a fluorescence detector that observes the atmosphere in a nadir observation mode from a near space altitude. A full description of the mission is given in the paper of Wiencke [1]. The detector consists of 2 Fresnel lenses which focus the image on a pixel-matrix of photomultipliers capable of single photoelectron counting in temporal frames of $2.5 \mu\text{s}$, named Gate Time Units (GTU). The detector's focal surface consists of a single Photo Detector Module (PDM) with 48×48 pixels. The module is organized in Elementary Cells (EC) with 16×16 pixels. The EC itself is a block of 2×2 Multi-Anode Photomultipliers (MA-PMT). The hardware-implemented First Level Trigger (FLT) [2] operating on a level of EC provides an initial filtration of the data. The trigger is optimized towards storage and bandwidth and it is not designed for a detailed pattern recognition. Therefore, after the 12-day flight it produced more than 175 000 candidate events. The events have been analyzed in several aspects [3].

Due to shortened duration of the mission the expected number of observed air shower events is one or fewer [4]. The goal of this work is to develop an approach that would reduce the candidate events list into a smaller subset that is eventually easier to review manually. A similar work has also been done independently of this work by Diaz et al. [5].

The approach described in this paper uses a feature extraction method to create a simpler representation of an event and then it uses established machine learning techniques to classify data into two classes - "*air shower*" and "*noise*". Below we present the efficiency of the method on datasets of simulated events. We also discuss the usage of the flight data events in unsupervised learning methods to identify clusters of events with similar features.

2. Model training and testing dataset

The dataset used for training and testing of machine learning models consisted of simulated EUSO-SPB1 data with air shower events (positive samples) and various noise samples from real EUSO-SPB1 acquisition data (negative samples). The dataset had two main categories of the negative samples: background noise frame sequences and a small subset of labeled triggered events from the acquisition data, which were known not to be the air shower events.

The samples handled by the models were created by a procedure which, if necessary, transforms input data into EUSO-SPB-like packets. An acquisition data packet consists of 128 frames, where at frame 40 the hardware trigger activation happens, thus there are background noise frames available from every acquisition packet. Simulated packets were created by selecting a sequence of first 32 frames from an acquisition data packet and repeating this sequence for the whole 128 GTU of the packet. The repetition was not expected to cause large biases, but other approaches should also be explored in the future. All data were processed by a software implementation of the FLT, and the trigger data were used for temporal segmentation of a packet. This aimed to extract only a part of the packet, in this work called "*the event frame sequence*", which was expected to contain "interesting" data - ideally persistent presence of increased photo-electron counts. Note that photo-

electron counts of the flight acquisition data considered in this analysis have been normalized by the application of the flat frame map created prior to the EUSO-SPB1 flight [6].

2.1 Positive samples

The simulation of the air shower events was done by the simulation module of the software framework ESAF [7]. The simulated primary particle energies ranged from $10^{17.6}$ eV to 10^{19} eV in constant steps on a logarithmic scale. Each energy step had a 0.05 higher exponent value than the previous one. The simulated balloon altitudes ranged from 18 to 33 km. Events were simulated uniformly arriving from all directions. However, note that in a resulting training dataset, the distributions were not uniform due to applied selections based on visibility of a simulated air shower. The simulated events had to contain recognizable air shower tracks, and the simulated event frame sequences had to contain at least 2 frames, where a maximum signal pixel value was greater than a maximum background noise pixel value. Applying both rules on the available data processed by the feature extraction algorithm, the total number of events available for analysis was 34521.

2.2 Negative samples

The main portion of the negative samples in the dataset were the event frame sequences triggered by the software FLT at least 20 frames before or after frame 42. These event frame sequences are called "*unlabeled noise events*". In the whole dataset, the number of such event frame sequences was 56229. Another significant part of the dataset were the event frame sequences triggered by the software FLT less than 20 frames before or after frame 42 and classified by a manual review. Such event frame sequences are referred to as "*labeled noise events*". The main difference of these events from the *unlabeled noise events* was that the *labeled flight events* were typically activations of the trigger due to high intensity peaks in the photo-electron counts. For the purpose of the air shower detection, the only important information was that these events are not air showers. However, several classes were identified on an intuitive basis. The most typically assigned classes of such events were the following: "*a single pixel*", "*the top-left elementary cell*", "*a suddenly increased background intensity*", "*a blob*", "*a large blob*", "*a bright blob*", "*the noise*", and "*a single GTU track*".

3. Feature extraction based approach

The feature extraction approach classified the event frame sequences in two general steps: feature extraction and classifier training. Following the temporal segmentation of the event frame sequence using data from the software FLT, this approach applied a feature extraction procedure which produced a set of features describing various properties of the sequence. Afterwards, the results were stored in a database table prepared for the analysis. The air showers or other types of events could then be searched by SQL queries specifying feature thresholds. However, such manual classification experiments did not yield satisfactory results. Therefore, it was decided to use a machine learning approach to perform a binary classification. The machine learning algorithm would, based on the training dataset, in a sense formulate such conditions.

3.1 Feature extraction procedure

The feature extraction procedure uses X-Y, GTU-X, GTU-Y projections of an event and processes them by a Hough transform with the normal parametrization [8]. A sequence of frames can be viewed as a three dimensional matrix. The projection is created by selecting a maximum value in a dimension not visible in that projection - it is the time in X-Y projection (a maximum value of a pixel within a sequence of frames), a column of PDM pixels in GTU-X, and a row of PDM pixels in GTU-Y projection. An inclined line with an inclination different from 0° and 90° indicates movement of a bright spot, 0° line indicates no movement, and 90° line indicates pixel value increase appearing only in a single frame (e.g. a track appearing only in a single frame).

The transform creates a parameter space (Hough space) which for each line associates line parameters (angle of the line's normal and the line's distance from the origin) with the sum of values of pixels crossing this line in the image. Important parts of the procedure are also the data preparation before the transform and the method of determining an optimal line orientation from the parameter space. Several variants of X-Y, GTU-X, GTU-Y projections are created by selecting only triggered pixels or pixels above a threshold. Several threshold levels are considered: pixels with values above mean nonzero pixel value + $\{0, 1, 2\} \times \sigma$, and pixels with values above threshold selected by the Yen method [9]. A Hough space is constructed for each of those projections. Then the space is thresholded in several levels (75 %, 85 %, 90 %, 95 % of a peak value), and it is analyzed afterwards.

Some examples of the extracted features include the following: the number of frames in an event frame sequence; the maximum and minimum pixel values; the properties of pixel clusters in the projections; Hough space properties - orientation and location of the most significant line, number of clusters in the parameter space, maximum and minimum cluster dimensions in the space (width, height, area), dimensions of a cluster with the maximum value, dimensions of a cluster with the maximum total sum of pixel values.

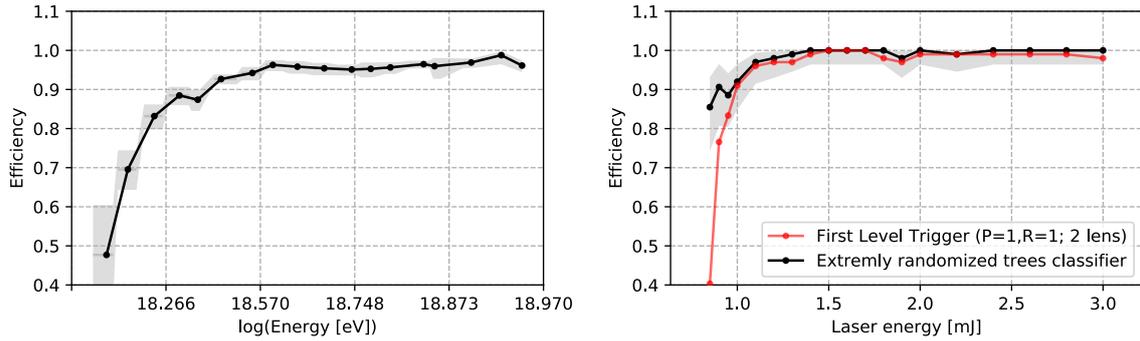
3.2 Classifier training

The positive and negative datasets were joined, balanced, shuffled and split into training and testing sets in 60:40 ratio. The training set samples were also weighted such that the total weights are equal between the labeled noise samples and the unlabeled noise samples. Each sample was initially represented by a set of more than 1000 features. Recursive feature elimination [10] (RFE) was used to select an optimal set of features to train the extremely randomized trees classifier [11].

Most of the selected features describe the width of the cluster in a Hough space. Unexpectedly, sets of 150-250 features in any classification experiment do not include orientation of the line in the time-wise projections of the air shower. Using the dataset described above, the most important feature, out of 153, is the width of a cluster in a Hough space from the X-Y projection of the event frame sequence. The projection was thresholded by selecting pixels with intensity above the mean value, and the Hough space was thresholded by selecting pixels above 75 % of maximum value in this space. The classifier trained on this feature alone achieved 79 % overall accuracy.

3.3 The classifier accuracy on the test data

Figure 1a shows efficiency (ratio correctly classified to all simulated event frame sequences



(a) Efficiency as a function of the energy on the test set. (b) Efficiency as a function of the laser energy.

Figure 1: Efficiency of the classifier.

in the test set) of the classifier described above as a function of energy of a simulated air shower event. The grey areas on the plot show binomial proportion confidence intervals (Clopper-Pearson intervals) at 95 % confidence level. The areas have a box-like shape because the width indicates range of the attribute values grouped under a single point. Figure 1a shows that the efficiency increases with energy of the primary particle - the plot can be approximated by the exponential function $\epsilon_{cls}(E) = -0.927 \exp(-1.123 \times 10^{-18} E) + 0.96$, where E is the primary particle energy in eV. Furthermore, the efficiency decreases with increasing distance of a shower maximum from the center of the FoV, and the efficiency slightly increases with the zenith angle. However, note that the plot shows results on the *visible tracks* dataset, if those rules were not applied, inclusion of the high-zenith angle events that are passing through an edge of the PDM might skew the results.

Overall non-balanced cross-validated accuracy of the presented classifier is approximately 95 %, overall sensitivity (efficiency) is 93 %, overall specificity is 98 %, but specificity on the labeled noise data of the test set is approximately 92 %. Based on a comparison with several previous classification experiments with a different labeled noise sample count (not shown here), the sensitivity seems to decrease as more air-shower-like noise samples are included in a training set.

4. Air shower recognition in the dataset of laser shots

An independent evaluation of the classifier performance was done by classifying the data from EUSO-SPB calibration and testing in 2016 [6, 12]. The goal to this exercise was to make sure that the classifier had not missed valid air shower events passed by the FLT. Figure 1b represents the same trigger data as the figure 3 in the paper of Bayer et. al [12]. However, note that these are long tracks going through the middle of the focal surface. Unlike the flight data, this dataset was not normalized by the flat field map, and this might have contributed to the high number of the event frame sequences identified by the software implementation of FLT - approximately 4.18 times more than the actual number of packets in the acquisition run. The analysis have also displayed a weakness of the current version of the classifier - an actual number of event frame sequences classified as an air shower was about 1.2 times higher than number of event frame sequences which have overlap with frame 40.

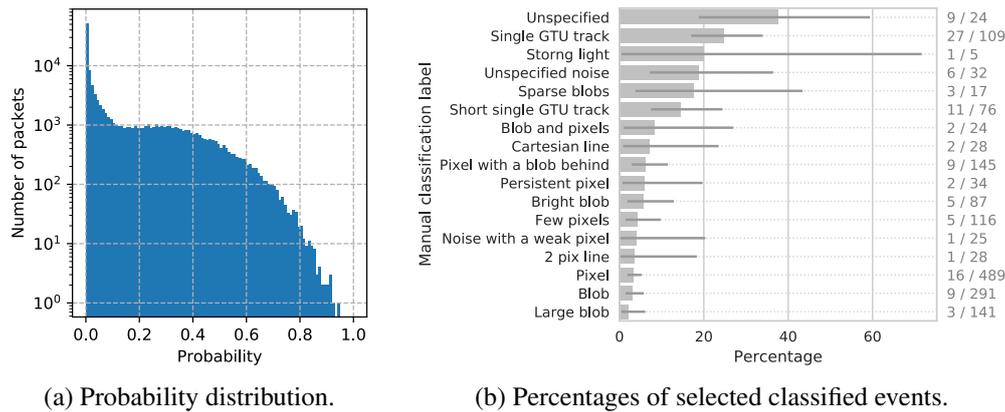


Figure 2: Classification of the EUSO-SPB flight data - samples classified as an air shower.

5. Air shower recognition in the EUSO-SPB flight data

The event frame sequences selected for the classification were required to contain at least 6 active elementary cells and the software FLT activation had to happen between frames 36 and 45. If there were several independently identified frame sequences in this range, the sequence closer to the frame 40 was selected. This cut yielded 118 141 event frame sequences (80 % of packets passed through the software trigger). After classifying the data using the extremely randomized trees classifier, 5450 events (4.6 %) had higher probability than 50 % of being an air shower. The number of these entries seems to depend on the labeled noise entries count in the training set - in this case it was 2619. A model trained without such data selected almost 50 %, and comparable models, trained on a dataset with 1800 of such entries, selected around 10 % of the classified entries.

Figure 2a shows the distribution of probabilities in the classified flight data: With the increasing probability, the number of packets decreases. There are only 87 packets with 80 % probability and only 7 packets with 90 % probability. Just for a comparison, all of the laser shots considered in the figure 1b have been classified as an air shower with higher than 80 % probability. A manual review of a subset of the samples, classified as an air shower in an order from entries with the highest probability, did not find any actual air shower.

Figure 2b shows the fraction of incorrectly classified to all labeled noise event frame sequences in the flight data. The plot indicates how successfully the classifier rejected false positive events in the labeled flight data subset. Error bars show binomial proportion confidence interval at 95 % confidence level. The interval was calculated via the Clopper-Pearson method considering all class samples count and the selected count. These numbers are shown on the right side of the plot. Note that in case of the flight data classification, some entries were in the classifier's training set.

In this case, almost 40 % of events classified "unspecified" have been selected. Out of an abundance of caution, these are types of events that have not been included in the training set. Although, these are not really considered to be the strong air shower candidates, the events were considered to share some similarity with an air shower pattern. Another typical false positives are "single GTU tracks". In all realized classification experiments so far, this class has been one of most typical false positives. An increase in a number of the events in the training set so far decreased the number of false positive classifications.

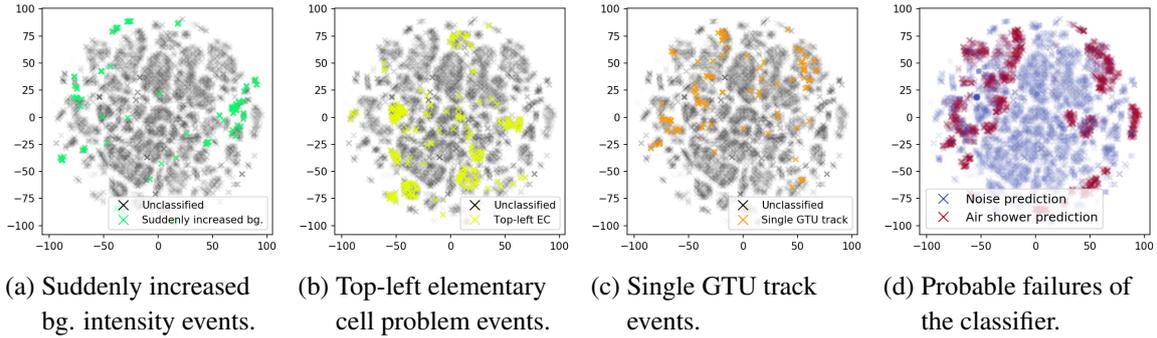


Figure 3: T-SNE of the flight data using the feature set optimized for the air shower classification.

6. An attempt to understand the data via T-SNE

T-distributed stochastic neighbor embedding (T-SNE) [13] is a non-linear dimensionality reduction technique. The technique iteratively optimizes point positions in a low-dimensional space (embedding) to ensure that they are similar if these points are similar in the original high-dimensional space. In this application, the points represent event frame sequences in the dataset. The technique provides a tool to evaluate whether the selected features capture the differences between the air showers and the noise. Although not shown here, the visualization of the T-SNE embedding on the testing and training data showed separation between the classes for most of the samples. On the other hand, the application of the technique on the unlabeled flight data with the small set of the labeled noise events helps us to better understand the dataset contents. The approach also allows us to visualize areas where the classifier tends to make false positive detections.

Figure 3d illustrates the application of the technique on the EUSO-SPB1 flight data. Note, that the cluster sizes are not well representing the number of samples in a cluster, and the actual absolute coordinates of the points are not very important. The first three figures show locations of manually labeled entries in the embedding. Types of events, such as an increase of pixel intensity over the whole PDM (fig. 3a) or a malfunction of top-left elementary cell (fig. 3b) are well identified and generally in the separate clusters. Single GTU track events (fig. 3c) are not as well separated but there is an observable clustering. The last figure (fig 3d) shows locations where the classifier marked the data as an air shower. Those should be investigated in detail, and eventually put into the training set to better optimize the classifier.

7. Conclusion and outlook

The paper describes one approach used in search for air shower events in the dataset of EUSO-SPB1. The approach relies on the features extracted by a dedicated feature extraction program. The classifier has achieved 95 % cross-validated balanced accuracy in detection of the air shower events in the testing dataset. Its comparison with the FLT efficiency on laser shot data have shown that the air shower events triggered by the FLT should not be rejected by the classifier. None of the currently applied approaches have found any air shower event in the EUSO-SPB1 flight data.

An alternative approach is to train, for instance, a neural network directly on the raw data. Such approach has also been investigated and it will be presented in future publications. Further

development of the air shower recognition methods will continue. Eventually, the methods will be important for the upcoming EUSO-SPB2 mission presently planned for the flight in 2022.

8. Acknowledgment

This work was partially supported by NASA grants NNX13AH54G, NNX13AH52G, French Space Agency (CNES), Italian Space Agency through the ASI INFN agreement n. 2017-8-H.0, Italian Ministry of Foreign Affairs and International Cooperation, the Basic Science Interdisciplinary Research Projects of RIKEN and JSPS KAKENHI Grant (22340063, 23340081, and 24244042), Deutsches Zentrum für Luft und Raumfahrt, 'Helmholtz Alliance for Astroparticle Physics HAP' funded by the Initiative and Networking Fund of Helmholtz Association (Germany), and by Slovak Academy of Sciences MVTS JEM-EUSO as well as VEGA grant agency project 2/0132/17. We acknowledge NASA Balloon Program Office, Columbia Scientific Balloon Facility, Telescope Array Collaboration, and Wanaka airport. The present research used resources of the National Energy Research Scientific Computing Center (NERSC), a US Department of Energy Office of Science User Facility operated under Contract No. DE-AC02-05CH11231.

References

- [1] L. Wiencke and A. Olinto, *EUSO-SPB1 Mission and Science*, in *Proc. 35th ICRC – PoS(ICRC2017)*, p. 1097, aug, 2017, [DOI](#).
- [2] M. Battisti et al., *Performance results of the trigger logic implemented in EUSO-SPB*, *Nucl. Instr. and Meth. A* (2018) .
- [3] J. Eser, *Results of the EUSO-SPB1 flight*, in *This conference proceedings*, 2019.
- [4] K. Shinozaki et al., *Estimation of the exposure for the air shower detection mode of EUSO-SPB1*, in *This conference proceedings*, 2019.
- [5] A. Diaz, *EUSO-SPB1: Flight data classification and Air shower search results*, in *This conference proceedings*, 2019.
- [6] J. B. Eser et al., *Preflight calibration and testing of EUSO-SPB in the lab and the desert*, in *Proc. 35th ICRC – PoS(ICRC2017)*, p. 457, aug, 2017, [DOI](#).
- [7] C. Berat et al., *Full simulation of space-based extensive air showers detectors with ESAF*, *Astropart. Phys.* **33** (2010) 221.
- [8] R. O. Duda and P. E. Hart, *Use of the Hough Transformation to Detect Lines and Curves in Pictures*, *Commun. ACM* **15** (1972) 11.
- [9] J.-C. Yen et al., *A new criterion for automatic multilevel thresholding*, *IEEE Trans. Image. Process.* **4** (1995) 370.
- [10] I. Guyon et al., *Gene selection for cancer classification using support vector machines*, *Machine Learning* **46** (2002) 389.
- [11] P. Geurts et al., *Extremely randomized trees*, *Machine Learning* **63** (2006) 3.
- [12] J. Bayer et al., *The trigger logic of EUSO-SPB and its performance*, in *Proc. 35th ICRC – PoS(ICRC2017)*, p. 443, aug, 2017, [DOI](#).
- [13] L. v. d. Maaten and G. Hinton, *Visualizing data using t-sne*, *JMRL* **9** (2008) 2579.