

# A model independent determination of the $B \rightarrow X_s \gamma$ decay rate

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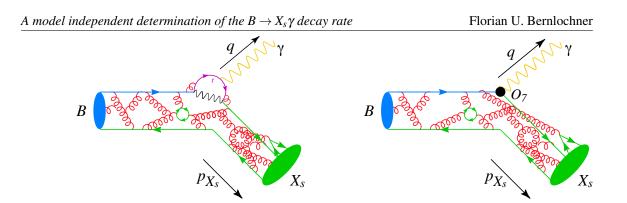
# (The SIMBA Collaboration)

The goal of the SIMBA collaboration is to provide a global fit to the available measurements of inclusive  $B \rightarrow X_s \gamma$  and  $B \rightarrow X_u \ell v$  decays. By performing a global fit one is able to simultaneously determine the relevant normalizations, i.e. the total  $B \rightarrow X_s \gamma$  rate and the CKM-matrix element  $|V_{ub}|$ , together with the required hadronic parameters, most importantly the *b*-quark mass and the *b*-quark distribution function in the *B*-meson, called the shape function. In this talk, the current status on the model-independent determination of the shape function and  $|C_7^{\text{incl}}V_{tb}V_{ts}^*|$ , which parametrizes the total  $B \rightarrow X_s \gamma$  rate, from a global fit to the available  $B \rightarrow X_s \gamma$  measurements from *B*<sub>A</sub>*B*<sub>A</sub>*R* and *Belle* is presented. In particular, the theoretical uncertainties originating from variations of the different factorization scales are evaluated.

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**Figure 1:** On the left-hand side the SM penguin  $B \to X_s \gamma$  decay is shown. In new-physics models, such as e.g. the minimal supersymmetric Standard Model, the Wilson coefficient associated with the effective operator  $\mathcal{O}_7$  shown at the right-hand side is modified.

## 1. Introduction

Rare penguin decays play an important role in the search for new physics in the flavor sector: the presence of new heavy particles which couple to heavy quarks would manifest itself in modifications of the total decay rate. In this presentation, a model-independent analysis is carried out to test the compatibility of four  $B \to X_s \gamma$  measurements with the Standard Model expectation (SM). In Fig. 1 the SM decay and a possible extension via a charged Higgs are shown. In the past, such an analysis used the extrapolation of the measured decay rates to the low  $E_{\gamma}$  region to compare the experimental measured partial branching fraction with the next-to-next-to-leading order SM prediction,  $\mathscr{B}(E_{\gamma} > 1.6 \,\text{GeV}) = (3.55 \pm 0.24 \pm 0.09) \times 10^{-4}$ , from Ref. [1, 2], as adopted e.g. by the analysis of Ref. [3] and more recent updates. The reason for this extrapolation lies in the theoretically poorly known non-perturbative corrections from the b quark distribution function in the *B* meson, called the shape function, which only allows precise prediction for the partial decay rate,  $\mathscr{B}(E_{\gamma} > E_{\gamma}^{\text{cut}})$ , for low values of  $E_{\gamma}^{\text{cut}}$ . The drawback of the extrapolation based analyses lie in the introduction of undesired model dependencies, which are hard to quantify, and also in making poor use of the experimentally most precise regions at high  $E_{\gamma}$  by including the partial branching fraction down to the region of phase-space with low  $E_{\gamma}$ , which is dominated by large background contributions from other B meson decays. The SIMBA collaboration uses an alternative approach, outlined in Ref. [4], by determining the shape function directly from data from a global analysis of the  $B \rightarrow X_s \gamma$  spectra with negligible model-dependence with respect to the present experimental and theoretical uncertainties.

## **2.** Treatment of the shape function and formulae for $B \rightarrow X_s \gamma$ decay rate

#### 2.1 Shape function

The shape function, which enters in the differential  $B \rightarrow X_s \gamma$  decay rate, can be factorized into perturbative and non-perturbative contributions, cf. [4], via

$$S(\boldsymbol{\omega},\boldsymbol{\mu}) = \int \mathrm{d}k \, \widehat{C}_0(\boldsymbol{\omega} - \boldsymbol{k},\boldsymbol{\mu}) \, \widehat{F}(\boldsymbol{k}) \,, \tag{2.1}$$

where  $\widehat{C}_0(\omega - k, \mu)$  is the  $\overline{\text{MS}}$ -renormalized *b*-quark matrix element of the shape function operator calculated in perturbation theory, and  $\widehat{F}(k)$  are the non-perturbative contributions to  $S(\omega, \mu)$ . Constructing the shape function as done in Eq. 2.1 offers several advantages over alternative approaches: the shape function has the correct perturbative tail at large  $\omega$ , and the correct RGE behavior; for small  $\omega$ , the shape function is dominated by the non-perturbative parameter  $\hat{F}(k)$ . This means that the shape of the  $B \to X_s \gamma$  spectrum at large  $E_{\gamma}$  is determined by the non-perturbative parameter  $\hat{F}(k)$ .

The approach outlined in Ref. [4] proposed the determination of  $\widehat{F}(k)$  directly from experimental spectra by employing an expansion in a set of complete orthonormal basis functions,  $f_n$ , as

$$\widehat{F}(k) = \frac{1}{\lambda} \left[ \sum_{n=0}^{\infty} c_n f_n\left(\frac{k}{\lambda}\right) \right]^2 \quad \text{with} \quad \int dk \,\widehat{F}(k) = \sum_{n=0}^{\infty} c_n^2 = 1, \quad (2.2)$$

where  $\lambda \simeq \Lambda_{\text{QCD}}$  is a dimensional parameter of the basis. Since the orthonormal basis in Eq. 2.2 is complete, this description offers a model-independent description of the shape function and the  $B \rightarrow X_s \gamma$  decay rate, and the shape of the spectrum is parametrized by the expansion coefficients  $c_n$ . These coefficients can be determined directly by fitting the available data, taking into account the full experimental uncertainties and correlations. In practice, however, the expansion in Eq. 2.2 needs to be truncated, since the available experimental information only allows for a finite number of N coefficients to be constrained. This truncation introduces a residual model-dependence, which depends on the chosen functional basis and scales as  $1 - \sum_{n=0}^{N} c_n^2$ . The optimal values for  $\lambda$  and N need to be determined from data:  $\lambda$  is chosen such that the fitted series converges quickly, and the number of used basis functions N should be chosen large enough such that the truncation uncertainty is small compared to the experimental uncertainties.

## **2.2** Master formulae for $B \rightarrow X_s \gamma$ decay rate

The  $B \rightarrow X_s \gamma$  photon energy spectrum is given by

$$\frac{\mathrm{d}\Gamma}{\mathrm{d}E_{\gamma}} = \frac{G_F^2 \alpha_{\mathrm{em}}}{2\pi^4} E_{\gamma}^3 \widehat{m}_b^2 |V_{tb} V_{ts}^*|^2 \\
\times \left\{ |C_7^{\mathrm{incl}}|^2 \left[ \int \mathrm{d}k \,\widehat{W}_{77}(k) \widehat{F}(m_B - 2E_{\gamma} - k) + \sum_m \widehat{W}_{77,m} \,\widehat{F}_m(m_B - 2E_{\gamma}) \right] \\
+ \int \mathrm{d}k \sum_{i,j\neq 7} \left[ 2\mathrm{Re}(C_7^{\mathrm{incl}}) C_i \,\widehat{W}_{7i}(k) + C_i C_j \,\widehat{W}_{ij}(k) \right] \widehat{F}(m_B - 2E_{\gamma} - k) \right\},$$
(2.3)

where  $\widehat{W}_{77}(k)$  contains the perturbative corrections to the  $b \to s\gamma$  decay via the electromagnetic dipole operator,  $O_7$ , resummed to next-to-next-to-leading-logarithmic order [5, 4], and including the full NNLO corrections  $[6, 7]^1$ ;  $\widehat{F}(k)$  is the leading shape function as introduced in the previous section, the  $\widehat{F}_m(k)$  denote  $1/\widehat{m}_b$  suppressed subleading shape functions, and  $\widehat{W}_{7i}(k)$  and  $\widehat{W}_{ij}(k)$  contain subleading perturbative corrections. The full set of expressions entering Eq. 2.3 will be given in Ref. [8]. In a fit to  $B \to X_s \gamma$  spectra the subleading shape functions can be absorbed at lowest order in  $\alpha_s$  into the leading shape function, reducing the number of coefficients which need to be

<sup>&</sup>lt;sup>1</sup>At lowest order in perturbation theory:  $\widehat{W}_{77}(k) = \delta(k)$ 

determined from data. The coefficient  $C_7^{\text{incl}}$  multiplying the dominant contribution proportional to  $\widehat{W}_{77}(k)$  in Eq. 2.3 is defined as

$$C_{7}^{\text{incl}} = C_{7}^{\text{eff}}(\mu_{0}) \frac{\overline{m}_{b}(\mu_{0})}{\widehat{m}_{b}} + \sum_{i=1}^{6} r_{i}(\mu_{0}) C_{i}(\mu_{0}) + r_{8}(\mu_{0}) C_{8}^{\text{eff}}(\mu_{0}) \frac{\overline{m}_{b}(\mu_{0})}{\widehat{m}_{b}} + \cdots, \qquad (2.4)$$

where  $C_i^{\text{eff}}(\mu_0)$  are the standard scheme-independent effective Wilson coefficients and  $\overline{m}_b(\mu_0)$  is the  $\overline{\text{MS}}$  *b*-quark mass. The coefficients  $r_{1-6,8}(\mu_0)$  contain all virtual contributions from the operators  $O_{1-6,8}$  that generate the same effective  $b \to s\gamma$  vertex as  $O_7$ . The ellipses denote included terms proportional to  $\ln(\mu_0/\hat{m}_b)$  that are required to cancel the  $\mu_0$  dependence on the right-hand side and vanish at  $\mu_0 = \hat{m}_b$ , such that  $C_7^{\text{incl}}$  is by definition  $\mu_0$ -independent to the order one is working at.

Since the terms in the last line in Eq. 2.3 are small, we consider  $|C_7^{\text{incl}} V_{tb} V_{ts}^*|$  as the parameter that determines the normalization of the  $B \to X_s \gamma$  rate. This normalization is extracted simultaneously with  $\hat{F}(k)$  from a fit to the various measured  $E_{\gamma}$  spectra. The important contributions from  $O_{1-6,8}$  are the virtual corrections contained in  $C_7^{\text{incl}}$ , which have a sizable effect on the normalization of the  $B \to X_s \gamma$  rate. By including them in  $C_7^{\text{incl}}$ , they explicitly do not affect the shape of the spectrum, and so do not enter in our fit. They instead enter in the SM prediction for  $C_7^{\text{incl}}$ , which can be computed independently. Below, the fit result is compared to the NLO SM value,  $C_7^{\text{incl}} = 0.354^{+0.011}_{-0.012}$  [8]. For a more stringent test for new physics, evaluating  $C_7^{\text{incl}}$  in the SM at NNLO along the lines of Refs. [1, 2] would prove very valuable.

## 3. Fit to the available *B*-Factory data

#### 3.1 Experimental data from BABAR and Belle

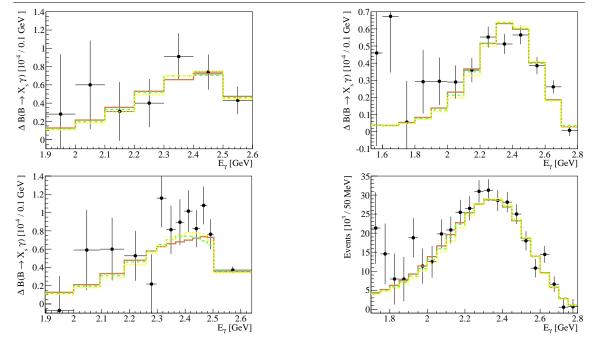
As experimental inputs the *Belle* measurement from Ref. [9], and the three *BABAR* measurements from Refs. [10, 11, 12] are used. The experimental statistical and systematic uncertainties and correlations are fully included in the fit procedure. The *BABAR* spectra of Refs. [10, 11] are measured in the *B* rest frame and are corrected for efficiencies. The experimental resolution in  $E_{\gamma}$  for each spectrum is smaller than its respective bin size, so both spectra can be directly used in the fit. The *Belle* spectrum from Ref. [9] is measured in the  $\Upsilon(4S)$  frame and affected by both efficiency and resolution. The *BABAR* result of Ref. [12] provides a resolution unfolded spectra with smaller bin-by-bin correlations in the *B* and the  $\Upsilon(4S)$  frame. We analyze the unfolded result in the  $\Upsilon(4S)$  frame. All four spectra are shown in Fig. 2.

## 3.2 Fit setup

To fit to the experimentally measured photon energy spectra, the expansion for  $\widehat{F}(k)$  in Eq. 2.2 is used for Eq. 2.3 and integrate over the appropriate range of  $E_{\gamma}$  for each experimental bin and each combination of basis functions  $f_m(x)f_n(x)$ . The theory prediction for the *i*th bin,  $B^i$ , is then given by

$$B^{i} = \widehat{m}_{b}^{2} |C_{7}^{\text{incl}} V_{tb} V_{ts}^{*}|^{2} \sum_{m,n=0}^{N} c_{m} c_{n} B_{mn}^{i} + \cdots, \qquad (3.1)$$

where the ellipses denote the additional included terms arising from the last line in Eq. 2.3. The overall  $\hat{m}_b^2$  is expressed in terms of the moments of  $\hat{F}(k)$ , so it is effectively a function of the  $c_n$ 



**Figure 2:** Used *Belle* and *BaBar*  $B \rightarrow X_s \gamma$  measurements: (top left) Ref. [10], (top right) Refs. [12], (bottom left) Ref. [11], (bottom right) Ref. [9]; The histograms show the result of the fit with a basis of  $\lambda = 0.5$  GeV with two (yellow), three (green), four (blue), and four (orange) coefficients. The default fit result uses four coefficients.

coefficients. A  $\chi^2$  minimization is performed to all available bins with  $c_{0,1,\dots,N}$  and  $|C_7^{\text{incl}}V_{tb}V_{ts}^*|$  as the fit parameters. The constraint  $c_0^2 + \dots + c_N^2 = 1$  is enforced to ensure that  $\widehat{F}(k)$  is properly normalized to unity. This fitting procedure was extensively tested using pseudo-experiments and provides unbiased central values with correct uncertainties.

### 3.3 Fit results

For our default fit a value of  $\lambda = 0.5$  GeV and four basis coefficients  $c_{0,1,2,3}$  was chosen: These values were selected after checking carefully the convergence of various basis expansions, number of coefficients, and with the general focus of avoiding over-tuning. The corresponding fits with two  $(c_{0,1})$ , three  $(c_{0,1,2})$ , four  $(c_{0,1,2,3})$ , and five  $(c_{0,1,2,3,4})$  expansion coefficients with  $\lambda = 0.5$  GeV are shown in Fig. 2. The fit converges after the inclusion of four coefficients and describes the measured spectra well. The  $\chi^2/ndf$  for the default fit with four coefficients is 41.65/48 corresponding to a p-value of 0.87. The fit results for the shape function for 2, 3, 4, and 5 basis coefficients are shown in the left panel of Fig. 3. The corresponding results for  $|C_7^{\text{incl}}V_{tb}V_{ts}^*|$  and  $m_b^{1S}$ , where the latter is computed from the moments of the fitted  $\widehat{F}(k)$ , are shown in the right panel of Fig. 3. The shape function in Fig. 3 verify the convergence of the basis expansion as the number of basis functions is increased. As one expects, the uncertainties returned by the fit increase with more coefficients due to the larger number of degrees of freedom. However, with too few coefficients one would have to add the truncation uncertainty. A reliable value for the final uncertainty is provided by the fitted uncertainty when the central values have converged and the respective last coefficients, here  $c_3$  or  $c_4$ , are compatible with zero. At this point, the truncation uncertainty can be neglected compared to the fit uncertainties. Equivalently, the increase in the fit uncertainties from including the last coefficient that is compatible with zero effectively takes into account the truncation uncertainty.

Using a fixed model function and fitting one or two model parameters would thus underestimate the true model uncertainties in the shape function model.

#### 3.4 Theoretical uncertainties from missing higher order perturbative corrections

Including the theory uncertainties from missing higher order perturbation corrections for the default fit gives

$$|C_7^{\text{incl}} V_{tb} V_{ts}^*| = (14.83 \pm 0.53_{[\text{exp}]} \pm 0.37_{[\text{theo}]}) \times 10^{-3},$$
  
$$m_b^{1S} = 4.77 \pm 0.03_{[\text{exp}]} \pm 0.02_{[\text{theo}]}, \qquad (3.2)$$

with comparable sizes for experimental and theoretical uncertainties. The result for  $|C_7^{\text{incl}}|$  is compatible within one sigma with the NLO SM value, shown as a grey band in Fig. 3, for which a value of  $|V_{tb}V_{ts}^*| = 40.68^{+0.4}_{-0.5} \times 10^{-3}$  was used.

## 4. Summary and outlook

We presented preliminary results from a global fit to  $B \to X_s \gamma$  data, which determines the total  $B \to X_s \gamma$  rate, parametrized by  $|C_7^{\text{incl}}V_{tb}V_{ts}^*|$ , and the *B*-meson shape function within a modelindependent framework. The value of  $|C_7^{\text{incl}}V_{tb}V_{ts}^*|$  extracted from data agrees with the SM prediction within uncertainties. From the moments of the extracted shape function we determine  $m_b^{1S}$ . In the future, information on  $m_b$  from other independent determinations can be included. The shape function extracted from  $B \to X_s \gamma$  is an essential input to the determination of  $|V_{ub}|$  from inclusive  $B \to X_u \ell \nu$  decays.

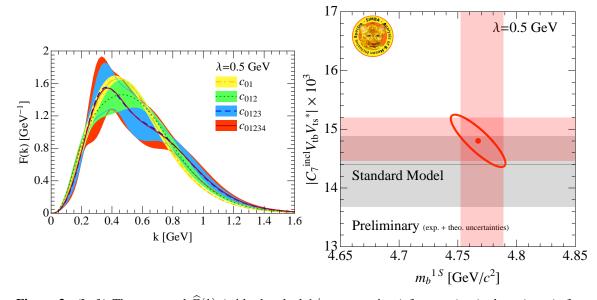
A combined fit to  $B \to X_s \gamma$  and  $B \to X_u \ell v$  measurements within our framework is in progress. It will allow for a simultaneous determination of  $|C_7^{\text{incl}}V_{tb}V_{ts}^*|$  and  $|V_{ub}|$  along with the shape function with reliable uncertainties. In addition to a few branching fractions with fixed cuts, it is important to have measurements of the  $B \to X_u \ell v$  differential spectra (including correlations), e.g. the lepton energy or hadronic invariant mass spectra. As for  $B \to X_s \gamma$ , fitting the differential spectra allows making maximal use of the  $B \to X_u \ell v$  measurement, by using the experimental most precise regions to constrain the nonperturbative inputs and further reduce the associated uncertainties.

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**Figure 3:** (Left) The extracted  $\hat{F}(k)$  (with absorbed  $1/m_b$  corrections) for two  $(c_{0,1})$ , three  $(c_{0,1,2})$ , four  $(c_{0,1,2,3})$ , and five  $(c_{0,1,2,3,4})$  coefficients and basis parameter  $\lambda = 0.5$  GeV. The colored envelopes are given by the uncertainties and correlations of the extracted coefficients  $c_n$ . (Right) Determined values of  $|C_7^{\text{incl}}V_{tb}V_{ts}^*|$  and  $m_b^{1S}$  with  $\Delta \chi^2 = 1$  contour. The grey band shows the SM value of  $|C_7^{\text{incl}}V_{tb}V_{ts}^*|$  and the pink band show the SCET scale variation uncertainties.

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