

Cosmic Rays from Heavy Dark Matter from the Galactic Center

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The gamma-ray fluxes observed by the High Energy Stereoscopic System (HESS) from the J1745-290 Galactic Center source is well fitted by the secondary photons coming from Dark Matter (DM) annihilation in particle-antiparticle standard model pairs over a diffuse power-law background. The spectral features of the signal are consistent with different channels: light quarks, electro-weak gauge bosons and top-antitop production. The amount of photons and morphology of the signal localized within a region of few parsecs, require compressed DM profiles as those resulting from baryonic contraction, which offer large enhancements in the signal over DM alone simulations. The fits return a heavy WIMP, with a mass above 10 TeV, but well below the unitarity limit for thermal relic annihilation. The fitted background spectral index is compatible with the Fermi-Large Area Telescope (LAT) data from the same region. This possibility can be potentially tested with the observations of other high energy cosmic rays.

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1. Gamma rays from the galactic center and the dark matter hypothesis

We have studied the possibility of explaining the gamma ray data [1] observed by the High Energy Stereoscopic System (HESS) from the central part of our galaxy by being partially produced by Dark Matter (DM) annihilations or decays [2, 3]. The complexity of the region and the amount of delocalized emitting sources justifies the hypothesis of a non-negligible background. DM annihilations or decays into single standard model particle-antiparticle channels provide good fits if the DM signal is complemented with such a background, that it is consistent with Fermi-LAT measurements [4].

The fits return a DM mass between 15 and 110 TeV depending on the studied channel [5]. Leptonic channels are clearly disfavored, but hadronic channels such as the $d\bar{d}$, or electroweak channels such as the W^+W^- and ZZ channels, offer very good consistency with $\chi^2/dof = 0.73$, 0.84 and 0.85 respectively [2]. The morphology of the signal requires compressed dark halos as the ones that take into account baryonic dissipation [6].

A motivated DM candidate which could have high enough masses and account for the right abundance of DM as a thermal relic, is the branon field [7, 8], which corresponds to brane fluctuations in flexible brane-world models. We have proved that a branon mass of $M \simeq 50.6$ TeV provides a good fit to the mentioned HESS data [2].

2. Conclusions and future work

The heavy DM masses required for fitting the HESS data are practically unconstrained by direct detection searches or colliders experiments [9]. On the contrary, this possibility can be tested with the observations of other cosmic-rays [10] from the GC and from other astrophysical objects. For example, high energy neutrinos could confirm this hypothesis at IceCube or ANTARES for angular resolutions below one degree.

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