



Kinetic Inductance Detectors as light detectors for neutrino and Dark Matter searches

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Large-mass arrays of bolometers proved to be good detectors for neutrinoless double beta decay (0vDBD) and Dark Matter searches. CUORE and LUCIFER are bolometric 0vDBD experiments that will start to take data in 2015 at Laboratori Nazionali del Gran Sasso in Italy. The sensitivity of CUORE to 0vDBD could be increased by removing the dominant source of background, due to α particles, exploiting the small amount of Cherenkov light emitted by the β (signal) and not by α s. LUCIFER could be extended to search also for Dark Matter interactions, provided that an efficient rejection of β/γ background in the energy region of interest (below 10 keV) can be obtained using the simultaneous detection of heat and scintillation signals. The interest in sensitive cryogenic light detectors gave birth to the CALDER project, that aims at the development of detectors with an active area of 5×5 cm² (the face of typical bolometric crystals), operating in a wide temperature range around 10 mK, and with an energy resolution at the baseline below 20 eV RMS. We have chosen to develop phonon-mediated detectors with Kinetic Inductance Detector (KID) sensors. We have tested the first prototypes, that were able to trigger the first pulses due to cosmic rays.

Technology and Instrumentation in Particle Physics 2014 2-6 June, 2014 Amsterdam, the Netherlands

1. Introduction

In the near future, bolometric experiments searching for rare events could play a key role in our understanding of neutrino physics and Dark Matter nature. The experimental sensitivity, that is mainly limited by the background due to natural radioactivity, could be enhanced by the capability to identify the nature of the interacting particle.

Bolometers are large crystals coupled with temperature sensors and operated as calorimeters at about 10 mK. They are particularly suitable for 0vDBD searches because they can reach high efficiency (larger than 75%), excellent energy resolution (0.1% from a few keV to a few MeV) and low intrinsic background. The lack of an active background rejection tool can be compensated using a light detector, that identifies the nature of the interacting particle exploiting the different light emission of different nature.

Two good examples to illustrate the potentiality of cryogenic light detection are CUORE [1] and LUCIFER [2], that will start taking data in 2015 at the Laboratori Nazionali del Gran Sasso (LNGS), Italy.

CUORE searches for the ¹³⁰Te 0vDBD using an array of 988 TeO₂ cubic crystals ($5 \times 5 \times 5$ cm³). The CUORE sensitivity to 0vDBD is currently limited by the background produced by α decays originating on the surface of the inert materials facing the detector. It was observed [3] that β particles, differently from α s, have sufficient energy to emit Cherenkov light in TeO₂. The signal is however of only about 100 eV [4] and therefore Light Detectors (LDs) with energy resolution below 20 eV are needed to achieve a S/N ratio of 4-5 and a satisfactory α background rejection.

LUCIFER will have a much smaller background compared to CUORE (zero background limit in the region of interest). This result will be obtained using scintillating crystals as bolometers and isotpes with high Q-values (e.g. ZnSe crystals enriched in ⁸²Se or ZnMoO₄ crystal enriched in ¹⁰⁰Mo). The LDs needed to obtain a good particle discrimination for this experiment has to be less performing (energy resolution of about 100 eV). They were already realized by the collaboration and consist of 5 cm diameter germanium slabs operated as bolometers [6]. On the other hand LU-CIFER would receive an important benefit from a LD with better resolution. In fact, with a energy resolution of 20 eV RMS it would have enough sensitivity to discriminate nuclear recoils from β/γ background in the region of spectrum where the signal from WIMP Dark Matter interaction is expected (0/20 keV) [5].

2. A new way for light detectors: CALDER

CALDER (Cryogenic wide-Area Light Detectors with Excellent Resolution) is a R&D started in 2014. The aim of the project is to develop large area, high energy resolution cryogenic LD, that are extremely needed to enhance rare events experiments as shown in previous section. The current state of the art in this field is the CRESST LD [7]. This detector, realised with Transition Edge Sensors, shows an excellent resolution (lower than 30 eV RMS), but suffers from reliability and reproducibility problems. Moreover the multiplexing readout system for a TES array is expensive and quite complex to realize.

We chose to develop our LD, using a novel technologies, called Kinetic Inductance Detectors (KIDs) [8]. KIDs are superconducting microresonator devices with very high Q factors $(10^4 -$



Figure 1: Left. A sketch of our resonator design. Right. The first 9-pixel prototype in his holder.

10⁶). The main advantages of this technology are represented by the excellent reliability, simple signal readout and easy scalability thanks to the intrinsic frequency multiplexing. The main issue regarding the adoption of KIDs is represented by the need to cover a large surface ($\approx 25 \text{ cm}^2$) with few pixels (4-10), because the largest possible size of a KID without an important loss of sensitivity and bandwidth is about 5 mm × 5 mm.

We are implementing an indirect phonon-mediated detection [9]: photons are absorbed in the substrate of the KID, generating a population of athermal phonons able to break Cooper pairs in the resonator. Using this approach, it is possible to reach high efficiency (up to 50 %) with relatively small filling ratio (about 10 %) if the main loss mechanisms (thermalization and escape of the phonons through the holder) are properly minimized. Our starting point is the the work of Moore [10], who developed a similar detector to be used for Dark Matter searches. With respect to this work we are not interested in resolving the position of interaction. This will allow us to enhance the merit factor of the detectors, realizing slower devices with higher sensitivity.

3. Status of the Project

We have recently completed the test of the first prototype of LD, produced at IFN of CNR (Rome). A 300 μ m thick, 2×2 cm² wide silicon substrate is sampled by 9 KID sensors, made with 40 nm thick aluminum deposited by electron gun. The lithography is performed using an EBL with a lift-off approach. The single pixel (left part of fig 1) has an active area of 2.3 mm² (total filling ratio of 5 %) and resonates at 2.5 GHz with a quality factor of 3 x 10⁴. This first array design was realized after studies to optimize detector parameter as the volume of the absorber substrate, the thickness of the resonator and the filling ratio. The detector's holder (see right panel of fig. 1) was designed to minimize the loss of energy due to the escape of the phonons through the supports.

During the first measures we were not able to test completely the device, since our electronics readout and analysis tools were still not well developed. We were able instead to study the shape of the resonances (see the left part of fig. 2). Using the readout system [11] developed for NIKA [12], we were also able to trigger CRs and light signals, sent through an optical fiber (respectively the red and black line in the right panel of fig. 2). The sensitivity of the device was strongly deteriorated by the broken feedline and we expect to improve significantly our sensitivity with the 2nd generation



Figure 2: Left. The 9 resonances measured in our first prototype **Right**. Two pulses caused by a Cosmic ray (red) and by a light signal (black).

of devices, that is currently in production. The new sample will include a simplification of the geometry of the array (from 9 to 4 pixels) in order to minimize the effect of the crosstalk, that is well evident in fig. 2. We will also try to increase the merit factor of the resonator up to 10^5 . The 2nd generation of devices will be tested in the next weeks, profiting of the fully operational readout electronics and of the model of response of the resonator to pulses, developed during the first tests. At the same time we are now beginning the study of materials different from aluminum (e.g. non stoichiometric TiN or multilayer Ti TiN), since their lower critical temperature and higher kinetic inductance will contribute to enhance the energy resolution of our LDs.

Acknowledgments

The CALDER experiment is funded by the European Research Council under the European Unions Seventh Framework Programme (FP7) / ERC grant agreement n. 335359.

References

- [1] C. Arnaboldi et al., Nucl. Instrum. Methods A518, 775-798 (2004)
- [2] J.W. Beeman et al., Advances in High Energy Physics 2013 (2013)
- [3] T. Tabarelli de Fatis, Eur. Phys. J. C65 (2010) 359.
- [4] N. Casali, M. Vignati et al. arxiv:1403.5528
- [5] J.W. Beeman et al., JINST 8, P05021 (2013)
- [6] J.W. Beeman et al., JINST 8, P07021 (2013)
- [7] G. Angloher et al., Eur. Phys. J. C72, 1971 (2012)
- [8] P. Day et al., Nature 425, 817-821 (2003)
- [9] L. Swenson et al., Appl. Phys. Lett. 96, 263511 (2010)
- [10] D.C. Moore et al., Appl. Phys. Lett. 100, 262601 (2012)
- [11] O. Bourrion et al., JINST 6, P06012 (2011)
- [12] A. Monfardini et al., ApJS 194, 24 (2011)