

# Status of RENO-50

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RENO-50, 18 kton ultra-low-radioactivity liquid scintillator detector is proposed to determine the neutrino mass hierarchy and to make highly precise measurement of  $\theta_{12}$ ,  $\Delta m_{21}^2$  and  $\Delta m_{31}^2$ . The detector is located at roughly 50 km away from the Hanbit nuclear power plant in Korea where the neutrino oscillation due to  $\theta_{12}$  takes place at maximum. The detector is expected to detect not only neutrinos from nuclear reactors but also neutrinos from the Sun, Supernova, the Earth, any possible stellar object and J-PARC neutrino beam. The experimental requirements and the strategy to achieve the main physics goals and sensitivities of RENO-50 are described.

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#### 1. Introduction

An underground detector of RENO-50 under proposal will consist of 18,000 tons of ultralow-radioactivity liquid scintillator and 15,000 high quantum efficiency 20" photomultiplier tubes, located at roughly 50 km away from the Hanbit nuclear power plant in Korea where the neutrino oscillation due to  $\theta_{12}$  takes place at maximum (see Figure 1). The detector is expected to detect neutrinos from nuclear reactors, the Sun, Supernova, the Earth, any possible stellar object and a J-PARC neutrino beam as well. It will be a multi-purpose and long-term operational detector, and also a neutrino telescope. The main goal is to determine the neutrino mass hierarchy and to measure the unprecedentedly accurate (<1%) values of  $\theta_{12}$ ,  $\Delta m_{21}^2$ , and  $\Delta m_{31}^2$ . It is also expected to detect ~5,600 events of a neutrino burst from a Supernova in our Galaxy, ~1,000 geo-neutrino events for 6 years, and ~200 events of muon neutrinos from the J-PARC beam every year.

#### 2. R&D for 3 % energy resolution at 1 MeV

The RENO-50 would be able to observe the manifestation of mass hierarchy in the oscillation effect if it establishes an extremely good energy resolution of ~3% at 1 MeV. The energy resolution could be achieved based on maximized light yield of LS larger than 1,000 photoelectrons per MeV, through making (a) an increased photosensitive area using 15,000 20" PMTs, (b) use of high (35%) quantum efficiency PMTs, and (c) an increased attenuation length of LS up to 25m. Figure 2 shows the expected energy resolution with and without improvement in light yield, quantum efficiency and attenuation length.

### 3. Study of precise (<1 %) determination of mixing parameters

The high precision measurements of  $\theta_{12}$ ,  $\Delta m_{21}^2$ , and  $\Delta m_{31}^2$  can make a strong impact on explaining the pattern of neutrino mixing and its origin. It will also provide useful information on the effort of finding a flavor symmetry. Figure 3 shows the result of MC study for



Expect ~10,000 events / year with oscillation





**Figure 2.** Expected energy resolution. Solid line is the case without improvement in light yield, quantum efficiency and attenuation length and dashed line is with improvement of them.

determination of mixing parameters. We expect to reach the following accuracies of measurements from 10 years data assuming fraction of background = 5 %, background error = 8 % and detection efficiency error = 1.0 %.

$$\delta(\Delta m_{21}^2) = -0.2\%$$
,  $\delta \sin^2(2\theta_{12}) = -0.8\%$ ,  $\delta(\Delta m_{21}^2) = -0.1\%$ 

Figure 3 shows the sensitivity of the measurments vs. Year.

## 4. Conclusion

We have studied the requirements and the strategy to reach 3 % energy resolution at 1 MeV and shown that high precision measurements (< 1 %) of  $\theta_{12}$ ,  $\Delta m_{21}^2$ , and  $\Delta m_{31}^2$  are reachable by RENO-50. A R&D funding (US \$ 2M) is given by the Samsung Science & Technology Foundation from the end of 2014, and will continue in the next 3 years. R&D efforts will be made on demonstrating feasibility of 3% energy resolution at 1 MeV, essential for determining the neutrino mass hierarchy.



Figure 3. Sensitivity of the measurement of the mixing parameters vs. year.