

## Belle II Early Bottomonia Physics Program

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The Belle-II experiment will begin ramping up data collection at the SuperKEKB accelerator facility in Tsukuba, Japan, within the next 1-2 years. Projecting peak luminosities approximately 40 times higher than KEKB, Belle-II is expected to accumulate  $50 \text{ ab}^{-1}$  by 2024, or a roughly 50-fold increase over Belle-I, enabling higher statistics investigation of a wide range of physics, including bottomonium spectroscopy. Additional scientific reach is afforded by the improved charged particle tracking and refined particle ID capabilities of Belle-II over Belle-I.

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## 1. Introduction

The Belle detector[1] has accumulated the world's largest data sample of electron-positron annihilations in the  $\sim 10$  GeV center-of-mass range. Although the bulk of the Belle data sample is collected at the  $\Upsilon(4S)$  center-of-mass energy in order to study weak decays of B-mesons, and, in particular, make precision measurements of CP-violation in the bottom sector, there has also been considerable data accumulated on the narrow  $\Upsilon(1S)$  and  $\Upsilon(2S)$  resonances, corresponding to  $5.7 fb^{-1}$  and  $24.7 fb^{-1}$  of luminosity on these lowest two  $J(PC)=1(- -)$  resonances, respectively, and  $1.8 fb^{-1}$  and  $1.7 fb^{-1}$  on the continua just below the  $\Upsilon(1S)$  and  $\Upsilon(2S)$  resonances, respectively. These data samples correspond to approximately  $(101.5 \pm 1.6) \times 10^6$  resonant  $\Upsilon(1S)$  and  $(157.8 \pm 3.6) \times 10^6$  resonant  $\Upsilon(2S)$  decays, offering enormous statistical improvements over the previous CLEO-III data sample, by factors of approximately 5 and 15, respectively.

## 2. Belle-II accelerator and beam

The KEKB accelerator will, relative to KEK, have higher beam currents, mostly by compressing the beam to a width of O(several hundred atomic layers). This will result in an enormous increase in luminosity; currently a 40-fold increase in  $\mathcal{L}$  ( $2.11 \times 10^{34}/cm^2s \rightarrow 80 \times 10^{34}/cm^2s$ ) is anticipated, such that after reaching the design luminosity in  $\sim 2022$ , we expect 50/ab delivered in the space of two years. The accelerator will also have slightly more energy reach than KEK, reaching  $ECM^{max}=11.24$  GeV ( $\Lambda_b\bar{\Lambda}_b$  threshold)

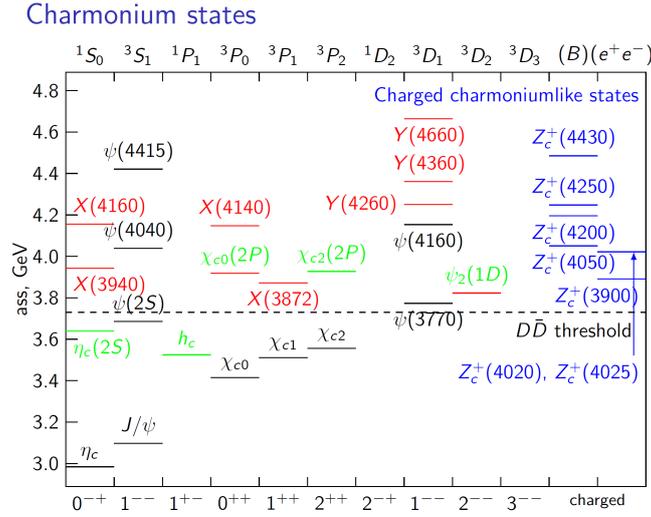
## 3. Physics Interest

### 3.1 The XYZ affair

Among the most interesting results to emerge from both the Babar and Belle experiments within the last 10-15 years has been the discovery of a rich, and unexpected spectroscopy of ‘‘charmonium-like’’ states, with masses in the range 3-4.5 GeV[2, 3, 4, 5]. These states, labeled ‘‘X’’, ‘‘Y’’, or ‘‘Z’’ are observed typically in one of four ways (charge conjugates are implicit): a) direct production in B-decay, e.g.,  $B \rightarrow XK$ , analogous to  $B \rightarrow J/\psi K$ , b) direct two-photon production, which selects parity and charge conjugation (++) states, c) correlated production, in association with a charmonium state:  $e^+e^- \rightarrow J/\psi(D\bar{D})$ , where the  $D\bar{D}$  invariant mass distribution near threshold often shows a significant excess above background, or d) direct observation in  $e^+e^-$  production, in association with an initial state radiation (ISR) photon, e.g.,  $e^+e^- \gamma_{ISR} \rightarrow Y(4660)$ ;  $Y \rightarrow \Lambda_c\bar{\Lambda}_c$ . Processes a), c) and d) also permit an inference of (PC) quantum numbers based on either the associated produced particle, or the decay products ( $D\bar{D}$  vs.  $D\bar{D}^*$ , e.g.). The spectroscopy of these charmonium-like states, which may have a parallel hierarchy of bottomonium-like states (the  $Z_b$  states are now well-established[6]) is shown in Figure 1.

### Running Plan Proposal

Since the existing  $\Upsilon(4S)$  and  $\Upsilon(5S)$  Belle-I samples are already satisfactorily large, the biggest science gains are likely to be obtained by running at the  $\Upsilon(3S)$  resonance, for which Belle-I has a sample of only  $3fb^{-1}$ ) and the  $\Upsilon(6S)$  resonance ( $M=11020$  MeV; Belle-I sample of  $6fb^{-1}$ ). Given



**Figure 1:** Spectroscopy of particles in the charmonium mass regime. Black represents ‘conventional’ known charmonium states; green are ‘conventional’ states which have not yet conclusively been identified, but may have already been observed (e.g., the X(3915) may be the  $\chi_{c0}(2P)$  and the X(3823) may be the  $\psi_2(1D)$ ). X, Y, and Z states are as shown.

that BaBar has  $30\text{fb}^{-1}$   $\Upsilon(3S)$  “in the can”, we would like to quickly obtain  $\sim 200\text{fb}^{-1}$  samples on both these resonances.

In addition to WXZ, ‘conventional’  $\Upsilon(6S)$  spectroscopic studies follow patterns observed already in  $\Upsilon(4S)$  and  $\Upsilon(5S)$  transitions, including likely final states:

- $\pi\pi\Upsilon(n^3D_J)$ ,  $\pi\pi h_b(3P)$  &  $\pi\pi\Upsilon(2D)$  (discovery)
- $\eta\Upsilon(mS)$  &  $\eta\Upsilon(n^3D_J)$ ;  $\eta\Upsilon(2D)$  (discovery)
- $KK\Upsilon(mS)$
- $\omega\chi_b(1P)$

Also at  $\text{ECM}=11020$  MeV, we can access the physics of excited  $B^*$  and  $B_s^*$  states, make precision measurements of the color factor  $R = \frac{\sigma(e^+e^- \rightarrow q\bar{q})}{\sigma(e^+e^- \rightarrow \mu^+\mu^-)}$ , and also measure the  $B_s$  decay constant,  $f_{B_s}$  from the annihilation channel:  $B_s \rightarrow \mu^+\mu^-$ . The latter represents a stringent test of lattice gauge calculations, as it depends critically on the wave-function overlap between the bottom and strange quarks.

### $\Upsilon(3S)$ physics

Running on the  $\Upsilon(3S)$ , below open  $B\bar{B}$  threshold, will hopefully not only resolve the D-states (which are also accessible in four-photon transitions in which there is one unit of  $\Delta J$  at each step:  $S \rightarrow P \rightarrow D \rightarrow P' \rightarrow S'$ ), but is, of course, one of the preferred modes for finding either vector or scalar exotica:  $\Upsilon(3S) \rightarrow \pi\pi\Upsilon(1S)$ ;  $\Upsilon(1S) \rightarrow$  “nothing” (V), or  $\Upsilon(1S) \rightarrow \gamma$  “nothing” (Axial Vector coupling). Of course, dark photon ( $A'$ ) exotica, e.g., may be probed at any center-of-mass energy:  $e^+e^- \rightarrow \gamma A'$ ;  $A' \rightarrow$  “invisible”, which is possible via mixing with SM  $\gamma$ . Depending on the nature of

the mixing, we may also fortuitously observe  $A' \rightarrow e^+e^-$ . One caution, however - high efficiency for exotica sensitivity often requires **loose** trigger configurations, as they may result in either single-photon and/or low-momentum charged track final states.

We can guesstimate some expected  $\Upsilon(3S)$  final state yields:

- $\gamma\eta_b(1S,2S)$ : if  $\mathcal{B}\epsilon=2\times 10^{-4} \Rightarrow 800/\text{fb}^{-1}$
- $\pi^0 h_b(1P); h_b \rightarrow \gamma\eta_b(1S)$ :  $\Rightarrow 400/\text{fb}^{-1}$
- $\gamma\chi_{b0}; \chi_{b0} \rightarrow \eta\eta_b(1S)(1S)$ : assume latter  $\mathcal{B} \sim 10^{-3}$ ;  $\Rightarrow 10/\text{fb}^{-1}$

## References

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