

Theory and Phenomenology of Leptonic CP Violation

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The current status of our knowledge of the 3-neutrino mixing parameters and of the CP violation in the lepton sector is summarised. The discrete symmetry approach to understanding the observed pattern of neutrino mixing and the related predictions for neutrino mixing angles and leptonic Dirac CP violation are reviewed.

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1. Introduction: the Three Neutrino Mixing

It was announced on February 24 this year (2017) that The Pontecorvo Prize for 2016 was awarded to Prof. Yifang Wang (from the Daya Bay Collaboration), Prof. Soo-Bong Kim (from RENO Collaboration) Prof. K. Nishikawa (from T2K Collaboration) "For their outstanding contributions to the study of the neutrino oscillation phenomenon and to the measurement of the Theta13 mixing angle in the Daya Bay, RENO and T2K experiments." As is well known, the relatively large value of the "reactor" (or "CHOOZ") angle $\theta_{13} \cong 0.15$ measured in the Daya Bay, RENO and Double Chooz experiments, indications for which were obtained first in the T2K experiment¹, opened up the possibility to search for CP violation effects in neutrino oscillations. Determining the status of CP symmetry in the lepton sector is one of the principal goals of the program of current and future research in neutrino physics. Information on leptonic Dirac CP violation is presently provided by the T2K and NOvA neutrino oscillation experiments using as input the reactor neutrino data on θ_{13} (see, e.g., [1]), and from analyses of the global neutrino oscillation data (see, e.g., [2, 3]). In the future it is expected to be provided principally by the planned DUNE [4] and T2HK [5] experiments.

Other goals of primal importance of the program of research in neutrino physics, which extends beyond 2030, include [1]:

- i) determination of the status of lepton charge conservation and the nature - Dirac or Majorana - of massive neutrinos (which is one of the most challenging and pressing problems in present day elementary particle physics);
- ii) determination of the spectrum neutrino masses possess, or neutrino mass ordering;
- iii) determination of the absolute neutrino mass scale, or $\min(m_j)$.

A successful realisation of this program² is of fundamental importance for making progress in understanding the origin of neutrino masses and mixing and its possible relation to new beyond the Standard Model (BSM) physics.

All compelling neutrino oscillation data is compatible with 3-neutrino mixing in vacuum, which we are going to consider in what follows:

$$\mathcal{L}_{CC} = -\frac{g}{\sqrt{2}} \sum_{l=e,\mu,\tau} \bar{l}_L(x) \gamma_\alpha \nu_{lL}(x) W^{\alpha\dagger}(x) + h.c., \quad \nu_{lL}(x) = \sum_{j=1}^3 U_{lj} \nu_{jL}(x), \quad (1.1)$$

where $\nu_{lL}(x)$ are the flavour neutrino fields, $\nu_{jL}(x)$ is the left-handed (LH) component of the field of the neutrino ν_j having a mass m_j , and U is a unitary matrix - the Pontecorvo, Maki, Nakagawa, Sakata (PMNS) neutrino mixing matrix [6, 7, 8], $U \equiv U_{PMNS}$.

In the case of 3 light neutrinos, the 3×3 unitary neutrino mixing matrix U can be parametrised, as is well known, by 3 angles and, depending on whether the massive neutrinos ν_j are Dirac or Majorana particles, by one Dirac, or one Dirac and two Majorana, CP violation (CPV) phases [9]:

$$U = VP, \quad P = \text{diag}(1, e^{i\frac{\alpha_{21}}{2}}, e^{i\frac{\alpha_{31}}{2}}), \quad (1.2)$$

¹For a review of the Daya Bay, RENO, Double Chooz and T2K data on θ_{13} see, e.g., [1].

²See, e.g., [1] for a rather detailed list of current and planned experiments that are foreseen to contribute to the comprehensive long-term program of research in neutrino physics.

Parameter	Best fit value	2σ range	3σ range
$\sin^2 \theta_{12}/10^{-1}$	2.97	2.65 – 3.34	2.50 – 3.54
$\sin^2 \theta_{13}/10^{-2}$ (NO)	2.15	1.99 – 2.31	1.90 – 2.40
$\sin^2 \theta_{13}/10^{-2}$ (IO)	2.16	1.98 – 2.33	1.90 – 2.42
$\sin^2 \theta_{23}/10^{-1}$ (NO)	4.25	3.95 – 4.70	3.81 – 6.15
$\sin^2 \theta_{23}/10^{-1}$ (IO)	5.89	$3.99 - 4.83 \oplus 5.33 - 6.21$	3.84 – 6.36
δ/π (NO)	1.38	1.00 – 1.90	$0 - 0.17 \oplus 0.76 - 2$
δ/π (IO)	1.31	0.92 – 1.88	$0 - 0.15 \oplus 0.69 - 2$
$\Delta m_{21}^2/10^{-5} \text{ eV}^2$	7.37	7.07 – 7.73	6.93 – 7.96
$\Delta m_{31}^2/10^{-3} \text{ eV}^2$ (NO)	2.56	2.49 – 2.64	2.45 – 2.69
$\Delta m_{23}^2/10^{-3} \text{ eV}^2$ (IO)	2.54	2.47 – 2.62	2.42 – 2.66

Table 1: The best fit values, 2σ and 3σ ranges of the neutrino oscillation parameters obtained in the global analysis of the neutrino oscillation data performed in [2] (The Table is taken from ref. [11]).

where $\alpha_{21,31}$ are the two Majorana CPV phases and in the “standard” parametrisation [1] the matrix V is given by:

$$V = \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{pmatrix}. \quad (1.3)$$

In eq. (1.3), $c_{ij} = \cos \theta_{ij}$, $s_{ij} = \sin \theta_{ij}$, the angles $\theta_{ij} = [0, \pi/2)$, and $\delta = [0, 2\pi]$ is the Dirac CPV phase. It follows from the current data that the three massive neutrinos $\nu_{1,2,3}$ should have masses not exceeding approximately 0.5 eV, $m_{1,2,3} \lesssim 0.5 \text{ eV}$. On the basis of the existing neutrino data it is impossible to determine whether the massive neutrinos ν_j are Dirac or Majorana fermions.

In the case of 3-neutrino mixing, oscillations involving all flavour neutrinos ν_l (antineutrinos $\bar{\nu}_l$), $\nu_l \leftrightarrow \nu_{l'}$ ($\bar{\nu}_l \leftrightarrow \bar{\nu}_{l'}$), $l, l' = e, \mu, \tau$, are possible. The 3-neutrino oscillation probabilities $P(\nu_l \rightarrow \nu_{l'})$ and $P(\bar{\nu}_l \rightarrow \bar{\nu}_{l'})$ are functions of the neutrino energy, E , the source-detector distance L , of the elements of U and, for relativistic neutrinos used in all neutrino experiments performed so far, of the two independent neutrino mass squared differences $\Delta m_{21}^2 \neq 0$ and $\Delta m_{31}^2 \neq 0$, ($\Delta m_{jk}^2 \equiv m_j^2 - m_k^2$) present in the case of 3-neutrino mixing (see, e.g., ref. [10]).

The existing data, accumulated over many years of studies of neutrino oscillations, allow us to determine Δm_{21}^2 , θ_{12} , and $|\Delta m_{31(32)}^2|$, θ_{23} and θ_{13} , with a relatively high precision [2, 3]. Since 2013 there are also persistent hints that the Dirac CPV phase δ has a value close to $3\pi/2$ (see [12]). The best fit values (b.f.v.) and the 2σ and 3σ allowed ranges of Δm_{21}^2 , s_{12}^2 , $|\Delta m_{31(32)}^2|$, s_{23}^2 , s_{13}^2 and δ , found in the latest analysis of global neutrino oscillation data performed in [2] are given in Table 1. Similar results were obtained in ref. [3]. In both analyses [2, 3] the authors find, in particular, that $\sin^2 \theta_{23} = 0.5$ lies outside the 2σ range allowed by the current data, but is within the 3σ allowed interval. Both groups also find that the best fit value of the Dirac CPV

phases δ is close to $3\pi/2$: in [2], for example, the authors find $\delta = 1.38\pi$ (1.31π) for $\Delta m_{31(32)}^2 > 0$ ($\Delta m_{31(32)}^2 < 0$). The absolute χ^2 minimum takes place for $\Delta m_{31(32)}^2 > 0$, the local minimum in the case of $\Delta m_{31(32)}^2 < 0$ being approximately by 0.7σ higher. According to ref. [2], the CP conserving value $\delta = 0$, or 2π , is disfavored at 2.4σ (3.2σ) for $\Delta m_{31(32)}^2 > 0$ ($\Delta m_{31(32)}^2 < 0$); the CP conserving value $\delta = \pi$ in the case of $\Delta m_{31(32)}^2 > 0$ ($\Delta m_{31(32)}^2 < 0$) is statistically approximately 2.0σ (2.5σ) away from the best fit value $\delta \cong 1.38\pi$ (1.31π). In what concerns the CP violating value $\delta = \pi/2$, it is strongly disfavored at 3.4σ (3.9σ) for $\Delta m_{31(32)}^2 > 0$ ($\Delta m_{31(32)}^2 < 0$)³. At 3σ , δ/π is found to lie in the case of $\Delta m_{31(32)}^2 > 0$ ($\Delta m_{31(32)}^2 < 0$) in the following intervals [2]: $(0.00 - 0.17(0.16)) \oplus (0.76(0.69) - 2.00)$. The results on δ obtained in [3] differ somewhat from, but are compatible at 1σ C.L. with, those found in [2].

It follows also from the results quoted in Table 1 that $\Delta m_{21}^2/|\Delta m_{31(32)}^2| \cong 0.03$. We have $|\Delta m_{31}^2| = |\Delta m_{32}^2 - \Delta m_{21}^2| \cong |\Delta m_{32}^2|$. The angle θ_{12} is definitely smaller than $\pi/4$: the value of $\theta_{12} = \pi/4$, i.e., maximal solar neutrino mixing, is ruled out at high confidence level by the data: $\cos 2\theta_{12} \geq 0.29$ at 99.73% C.L. The quoted results imply also that the value of θ_{23} can deviate by approximately ± 0.1 from $\pi/4$, $\theta_{12} \cong \pi/5.4$ and that $\theta_{13} \cong \pi/20$. Thus, the pattern of neutrino mixing differs drastically from the pattern of quark mixing.

Apart from the hint that the Dirac phase $\delta \sim 3\pi/2$, no other experimental information on the Dirac and Majorana CPV phases in the neutrino mixing matrix is available at present. Thus, the status of CP symmetry in the lepton sector is essentially unknown. With $\theta_{13} \cong 0.15 \neq 0$, the Dirac phase δ can generate CP violating effects in neutrino oscillations [9, 13], i.e, a difference between the probabilities of the $\nu_l \rightarrow \nu_{l'}$ and $\bar{\nu}_l \rightarrow \bar{\nu}_{l'}$ oscillations, $l \neq l' = e, \mu, \tau$. The magnitude of CP violation in $\nu_l \rightarrow \nu_{l'}$ and $\bar{\nu}_l \rightarrow \bar{\nu}_{l'}$ oscillations, $l \neq l' = e, \mu, \tau$, is determined by [14] the rephasing invariant J_{CP} , associated with the Dirac CPV phase in U :

$$J_{CP} = \text{Im} (U_{\mu 3} U_{e 3}^* U_{e 2} U_{\mu 2}^*) . \quad (1.4)$$

It is analogous to the rephasing invariant associated with the Dirac CPV phase in the CKM quark mixing matrix [15]. In the standard parametrisation of the neutrino mixing matrix (1.3), J_{CP} has the form:

$$J_{CP} \equiv \text{Im} (U_{\mu 3} U_{e 3}^* U_{e 2} U_{\mu 2}^*) = \frac{1}{8} \cos \theta_{13} \sin 2\theta_{12} \sin 2\theta_{23} \sin 2\theta_{13} \sin \delta . \quad (1.5)$$

Thus, given the fact that $\sin 2\theta_{12}$, $\sin 2\theta_{23}$ and $\sin 2\theta_{13}$ have been determined experimentally with a relatively high precision, the size of CP violation effects in neutrino oscillations depends essentially only on the magnitude of the currently not well determined value of the Dirac phase δ . The current data implies $0.026(0.027)|\sin \delta| \lesssim |J_{CP}| \lesssim 0.035|\sin \delta|$, where we have used the 3σ ranges of $\sin^2 \theta_{12}$, $\sin^2 \theta_{23}$ and $\sin^2 \theta_{13}$ given in Table 1. For the current best fit values of $\sin^2 \theta_{12}$, $\sin^2 \theta_{23}$, $\sin^2 \theta_{13}$ and δ we find in the case of $\Delta m_{31(2)}^2 > 0$ ($\Delta m_{31(2)}^2 < 0$): $J_{CP} \cong 0.032 \sin \delta \cong -0.030$ ($J_{CP} \cong 0.032 \sin \delta \cong -0.027$). Thus, if the indication that δ has a value close to $3\pi/2$ is confirmed by future more precise data, i) the J_{CP} factor in the lepton sector would be approximately by 3 orders of magnitude larger in absolute value than the corresponding J_{CP} factor in the quark sector, and ii) the CP violation effects in neutrino oscillations would be relatively large and observable.

³The quoted confidence levels for $\delta = 0, \pi$ and $\pi/2$ are all with respect to the absolute χ^2 minimum.

If the neutrinos with definite masses ν_i , $i = 1, 2, 3$, are Majorana particles, the 3-neutrino mixing matrix contains two additional Majorana CPV phases [9]. However, the flavour neutrino oscillation probabilities $P(\nu_l \rightarrow \nu_{l'})$ and $P(\bar{\nu}_l \rightarrow \bar{\nu}_{l'})$, $l, l' = e, \mu, \tau$, do not depend on the Majorana phases [9, 16]. The Majorana phases can play important role, e.g. in $|\Delta L| = 2$ processes like neutrinoless double beta $((\beta\beta)_{0\nu^-})$ decay $(A, Z) \rightarrow (A, Z + 2) + e^- + e^-$, L being the total lepton charge, in which the Majorana nature of massive neutrinos ν_i manifests itself (see, e.g. refs. [10, 17]).

Our interest in the CPV phases present in the neutrino mixing matrix is stimulated also by the intriguing possibility that the Dirac phase and/or the Majorana phases in U_{PMNS} can provide the CP violation necessary for the generation of the observed baryon asymmetry of the Universe (BAU) [18] (for specific models in which this possibility is realised see, e.g., [19]).

Understanding the origin of the patterns of neutrino mixing and of neutrino mass squared differences, revealed by the data obtained in the neutrino oscillation experiments is one of the most challenging problems in neutrino physics. It is part of the more general fundamental problem in particle physics of understanding the origins of flavour, i.e., of the patterns of quark, charged lepton and neutrino masses, and of the quark and lepton mixing.

In this article we will review aspects of the SYMMETRY approach to understanding the form of neutrino mixing, which is based on non-Abelian discrete flavour symmetries and is widely explored at present (see, e.g., [20, 21, 22] and references therein). One of the most striking features of this approach is that it leads to specific correlations between the values of at least some of the mixing angles of the neutrino mixing matrix U_{PMNS} and, either to specific fixed values of CPV phases present in U_{PMNS} , which are “trivial” (e.g., $\delta = 0$ or π , $\alpha_{21} = \alpha_{31} = 0$), (see, e.g., [22]), or to a correlation between the values of the neutrino mixing angles and of the cosine of the Dirac CPV phase δ of U_{PMNS} [23, 24, 25, 26]⁴, i.e., to a “sum rule” for $\cos \delta$. As a consequence of this correlation one obtains predictions for the value of δ and, correspondingly, for the J_{CP} factor and for the CP violating effects in neutrino oscillations. These predictions depend, in particular, on the underlying discrete symmetry used to derive the observed pattern of neutrino mixing and on the type of breaking of the symmetry, necessary to reproduce the measured values of the neutrino mixing angles. We will review also the predictions for δ and the J_{CP} factor in the cases of widely discussed underlying symmetry patterns of the PMNS matrix and the prospects of testing these predictions in future planned neutrino oscillation experiments.

2. Discrete Symmetry Approach to Neutrino Mixing (The Quest for Nature's Message)

We believe, and we are not alone in holding this view, that with the observed pattern of neutrino mixing Nature is “sending” us a Message. The Message is encoded in the values of the neutrino mixing angles, leptonic CPV phases in the PMNS matrix and neutrino masses. We do not know at present what is the content of Nature's Message. However, on the basis of the current ideas about the possible origins of the observed pattern of neutrino mixing, the Nature's Message can have two

⁴In the case of Majorana massive neutrinos one can obtain (under specific conditions) also correlations between the values of the two Majorana CPV phases present in U_{PMNS} and of the three neutrino mixing angles and of the Dirac CPV phase [23].

completely different contents, each of which can be characterised by one word: ANARCHY or SYMMETRY. In the ANARCHY approach [27] to understanding the pattern of neutrino mixing it is assumed that Nature “threw dice” when Nature was “choosing” the values of the neutrino masses, mixing angles and leptonic CPV phases. The main prediction of the ANARCHY explanation of the pattern of neutrino mixing is the absence of whatever correlations between the values of the neutrino mixing angles, between the values of the neutrino mixing angles and the CPV phases and between the values of the neutrino masses, all of them being random quantities. In contrast, one of the most characteristic prediction of the SYMMETRY approach to neutrino mixing is the existence of correlations between the values of at least some of the observables (angles, CPV phases) of the neutrino mixing matrix.

Within the SYMMETRY approach, the observed pattern of neutrino mixing can be naturally understood on the basis of specific class of symmetries - the class of non-Abelian discrete (finite) flavour symmetries (see, e.g., [20, 21, 22]). Thus, the specific form of the neutrino mixing can have its origin in the existence of new fundamental symmetry in the lepton sector. The most distinctive feature of the approach to neutrino mixing based on non-Abelian discrete flavour symmetries is the predictions i) of the values of some of the neutrino mixing angles and leptonic CPV phases, and/or ii) of existence of correlations between the values of at least some the neutrino mixing angles and/or between the values of the neutrino mixing angles and the Dirac CPV phase in the PMNS matrix, etc. (see, e.g., [22, 23, 24, 25])⁵. Most importantly, these predictions and predicted correlations, and thus the discrete symmetry approach itself, can be tested experimentally (see, e.g., [23] and [11, 24, 33, 34, 35]).

2.1 Symmetry Forms of Neutrino Mixing

The observed pattern of neutrino mixing in the reference 3-neutrino mixing scheme we are going to consider in what follows is characterised, as we have seen, by two large mixing angles θ_{12} and θ_{23} , and one small mixing angle θ_{13} : $\theta_{12} \cong 33^\circ$, $\theta_{23} \cong 45^\circ \pm 6^\circ$ and $\theta_{13} \cong 8.4^\circ$. These values can naturally be explained by extending the Standard Theory (ST) with a flavour symmetry corresponding to a non-Abelian discrete (finite) group G_f . This symmetry is supposed to exist at some high-energy scale and to be broken at lower energies to residual symmetries of the charged lepton and neutrino sectors, described respectively by subgroups G_e and G_ν of G_f . Flavour symmetry groups G_f that have been used in this approach to neutrino mixing and lepton flavour include S_4 , A_4 , T' , A_5 , D_n (with $n = 10, 12$), $\Delta(27)$, the series $\Delta(6n^2)$, to name several⁶ (see, e.g., ref. [21] for definitions of these groups and discussion of their properties⁷). The numbers of elements, of generators and of irreducible representations of the groups S_4 , A_4 , T' , A_5 , D_{10} and D_{12} are given in Table 2. The choice of the non-Abelian discrete groups S_4 , A_4 , T' , A_5 , etc. is related, in par-

⁵Combining the discrete symmetry approach with the idea of generalised CP invariance [28, 29] – a generalisation of the standard CP invariance requirement – allows to obtain predictions also for the Majorana CPV phases in the PMNS matrix in the case of massive Majorana neutrinos (see, e.g., [11, 30, 31, 32] and references quoted therein).

⁶Some of the groups T' , A_5 , etc. can be and have been used also for a unified description of the quark and lepton flavours, see, e.g., refs. [36] and references quoted therein.

⁷ S_4 is the group of permutations of 4 objects and the symmetry group of the cube. A_4 is the group of even permutations of 4 objects and the symmetry group of the regular tetrahedron. T' is the double covering group of A_4 . A_5 is the icosahedron symmetry group of even permutations of five objects, etc. All these groups are subgroups of $SU(3)$ and this will be assumed to hold for G_f considered by us.

Group	Number of elements	Generators	Irreducible representations
S_4	24	S, T, U	$\mathbf{1}, \mathbf{1}', \mathbf{2}, \mathbf{3}, \mathbf{3}'$
A_4	12	S, T	$\mathbf{1}, \mathbf{1}', \mathbf{1}'', \mathbf{3}$
T'	24	S, T, R	$\mathbf{1}, \mathbf{1}', \mathbf{1}'', \mathbf{2}, \mathbf{2}', \mathbf{2}'', \mathbf{3}$
A_5	60	S, T	$\mathbf{1}, \mathbf{3}, \mathbf{3}', \mathbf{4}, \mathbf{5}$
D_{10}	20	A, B	$\mathbf{1}_1, \mathbf{1}_2, \mathbf{1}_3, \mathbf{1}_4, \mathbf{2}_1, \mathbf{2}_2, \mathbf{2}_3, \mathbf{2}_4$
D_{12}	24	A, B	$\mathbf{1}_1, \mathbf{1}_2, \mathbf{1}_3, \mathbf{1}_4, \mathbf{2}_1, \mathbf{2}_2, \mathbf{2}_3, \mathbf{2}_4, \mathbf{2}_5$

Table 2: Number of elements, generators and irreducible representations of some discrete groups.

particular, to the fact that they describe symmetries with respect to rotations on fixed large mixing angles and, correspondingly, lead to values of the neutrino mixing angles θ_{12} and θ_{23} , which can differ from the measured values at most by sub-leading perturbative corrections, with θ_{13} typically (but not universally) predicted to be zero. For instance, the groups A_4 , S_4 and T' are commonly utilised to generate tri-bimaximal (TBM) mixing [37]; the group S_4 can also be used to generate bimaximal (BM) mixing⁸ [39]; A_5 can be utilised to generate golden ratio type A (GRA) [40] mixing; and the groups D_{10} and D_{12} can lead to golden ratio type B (GRB) [41] and hexagonal (HG) [42] mixing. For all these symmetry forms the neutrino mixing matrix U_ν° has the form: $U_\nu^\circ = R_{23}(\theta_{23}^\nu)R_{13}(\theta_{13}^\nu)R_{12}(\theta_{12}^\nu)$ with $\theta_{23}^\nu = -\pi/4$ and $\theta_{13}^\nu = 0$:

$$U_\nu^\circ = R_{23}(\theta_{23}^\nu = -\pi/4)R_{12}(\theta_{12}^\nu) = \begin{pmatrix} \cos \theta_{12}^\nu & \sin \theta_{12}^\nu & 0 \\ -\frac{\sin \theta_{12}^\nu}{\sqrt{2}} & \frac{\cos \theta_{12}^\nu}{\sqrt{2}} & -\frac{1}{\sqrt{2}} \\ -\frac{\sin \theta_{12}^\nu}{\sqrt{2}} & \frac{\cos \theta_{12}^\nu}{\sqrt{2}} & \frac{1}{\sqrt{2}} \end{pmatrix}. \quad (2.1)$$

The value of the angle θ_{12}^ν , and thus of $\sin^2 \theta_{12}^\nu$, depends on the symmetry form of U_ν° . For the TBM, BM, GRA, GRB and HG forms we have: i) $\sin^2 \theta_{12}^\nu = 1/3$ (TBM), ii) $\sin^2 \theta_{12}^\nu = 1/2$ (BM), iii) $\sin^2 \theta_{12}^\nu = (2+r)^{-1} \cong 0.276$ (GRA), r being the golden ratio, $r = (1 + \sqrt{5})/2$, iv) $\sin^2 \theta_{12}^\nu = (3-r)/4 \cong 0.345$ (GRB), and v) $\sin^2 \theta_{12}^\nu = 1/4$ (HG).

In the approach under discussion it is standardly assumed that the LH neutrino fields, $\nu_{iL}(x)$, and the LH components of the charged lepton fields (in the basis in which charged lepton mass term is not diagonal) $\tilde{l}_L(x)$, which form an $SU(2)_L$ doublet in the Standard Theory, are assigned to the same r -dimensional irreducible unitary representation $\rho_r(g_f)$ of the Group G_f , g_f being an element of G_f . In the cases of $G_f = A_4, S_4, T'$ and A_5 , which possess 3-dimensional irreducible representations, $\rho(g_f)$ is standardly taken to be a 3-dimensional irreducible unitary representation $\mathbf{3}$, $\rho_r(g_f) = \rho_3(g_f)$. This is equivalent to the assumption of unification of the three lepton families at some high energy scale. We are going to consider this choice in what follows.

⁸Bimaximal mixing can also be a consequence of the conservation of the lepton charge $L' = L_e - L_\mu - L_\tau$ (LC) [38], supplemented by $\mu - \tau$ symmetry.

At low energies the flavour symmetry G_f has necessarily to be broken so that the three lepton flavours can be distinguished and the electron, muon and tauon as well as the three neutrinos with definite mass ν_1 , ν_2 and ν_3 , can get different masses. Thus, G_f is broken to different residual symmetries G_e and G_ν , $G_e \neq G_\nu$, of the charged lepton and neutrino mass terms, respectively⁹. Possible discrete symmetries G_e of the charged lepton mass term (leaving $M_e M_e^\dagger$ invariant, M_e being the charged lepton mass matrix in left-right (L-R) convention) are: i) $G_e = Z_n$, with integer $n \geq 2$, or ii) $Z_m \times Z_k$, with integer $m, k \geq 2$. The maximal symmetry G_ν of the Majorana mass term of the LH flavour neutrino fields $\nu_{iL}(x)$ (leaving M_ν and $M_\nu^\dagger M_\nu$ invariant, M_ν being the mass matrix in R-L convention) is the $Z_2 \times Z_2$ (sometimes referred to as the Klein four group) symmetry. G_ν can obviously be just Z_2 . The subgroup G_e , in particular, can be trivial.

For fixed G_f and irreducible representation $\rho_r(g_f)$, non-trivial residual symmetries constrain the forms of the 3×3 unitary matrices U_e and U_ν , which diagonalise the charged lepton and neutrino mass matrices, and the product of which represents the PMNS matrix:

$$U_{\text{PMNS}} = U_e^\dagger U_\nu. \quad (2.2)$$

Thus, the residual symmetries constrain also the form of U_{PMNS} .

The TBM form of U_ν° (see eq. (2.1)) can originate from $G_f = S_4$ symmetry with residual symmetry $G_\nu = Z_2 \times Z_2$; it can be obtained also from a $G_f = A_4$ symmetry with $G_\nu = Z_2$ and imposing additional ‘‘accidental’’ $\mu - \tau$ (i.e., Z_2) symmetry of the neutrino Majorana mass matrix M_ν (see, e.g., [22] and references quoted therein). The group $G_f = S_4$ can also be used to generate the BM form of U_ν° (e.g., by choosing $G_\nu = Z_2$ combined with an accidental $\mu - \tau$ symmetry) [43, 25]. In all these cases $U_\nu = U_\nu^\circ P^\circ$, $P^\circ = P^\circ(\xi_{21}, \xi_{31})$ being a diagonal matrix containing two phases ξ_{21} and ξ_{31} which contribute to the Majorana phases α_{21} and α_{31} .

In the symmetry approach to neutrino mixing typically the matrix U_ν has an underlying symmetry form, for example, TBM, BM, GRA, GRB, HG. For all these five forms $U_\nu = U_\nu^\circ P^\circ$, $\theta_{13}^\nu = 0$ and needs to be corrected; if the measured θ_{23} is established to differ significantly from $\pi/4$, $|\theta_{23}^\nu| = \pi/4$ should be corrected. The sub-leading perturbative corrections, needed to bring the ‘‘symmetry’’ values of the three neutrino mixing angles in U_ν° to the measured values of θ_{12} , θ_{23} and θ_{13} in U_{PMNS} can most naturally be provided by the unitary matrix U_e (see, e.g., [26]). In certain classes of models, however, U_e coincides with the unit 3×3 matrix and the requisite corrections are incorporated in a factor contained in the matrix U_ν (see, e.g., [22]).

As we have indicated, one of the main characteristics of the discussed approach to neutrino mixing based on discrete flavour symmetries is that it leads to certain specific predictions for the values of, and/or correlations between, the low-energy neutrino mixing parameters, which can be tested experimentally. These predictions depend on the chosen G_f , $\rho(g_f)$, G_e and G_ν . We give a few examples [22, 23, 24, 25, 29, 30, 31, 32, 34, 43].

- I.** In a large class of models one gets $\sin^2 \theta_{23} = 0.5$.
- II.** In different class of models one finds that the values of $\sin^2 \theta_{23}$ and $\sin^2 \theta_{13}$ are correlated: $\sin^2 \theta_{23} = 0.5(1 \mp \sin^2 \theta_{13} + O(\sin^4 \theta_{13}))$.
- III.** In certain models $\sin^2 \theta_{23}$ is predicted to have specific values which differ significantly from

⁹Given a discrete (finite) G_f , there are more than one (but still a finite number of) possible residual symmetries G_e and G_ν , see, e.g., [21, 22].

those in cases **I** and **II** [24]: $\sin^2 \theta_{23} = 0.455$; or 0.463 ; or 0.537 ; or 0.545 , the uncertainties in these predictions being insignificant.

IV. Certain class of models predict a correlation between the values of $\sin^2 \theta_{12}$ and $\sin^2 \theta_{13}$: $\sin^2 \theta_{12} = 1/(3 \cos^2 \theta_{13}) = (1 + \sin^2 \theta_{13} + O(\sin^4 \theta_{13}))/3 \cong 0.340$, where we have used the b.f.v. of $\sin^2 \theta_{13}$.

V. In another class of models one still finds a correlation between the values of $\sin^2 \theta_{12}$ and $\sin^2 \theta_{13}$, which, however, differs from that in Case **IV**: $\sin^2 \theta_{12} = (1 - 3 \sin^2 \theta_{13})/(3 \cos^2 \theta_{13}) = (1 - 2 \sin^2 \theta_{13} + O(\sin^4 \theta_{13}))/3 \cong 0.319$, where we have used again the b.f.v. of $\sin^2 \theta_{13}$.

VI. In large classes of models in which the elements of the PMNS matrix are predicted to be functions of just one real continuous free parameter (“one-parameter models”), the Dirac and the Majorana CPV phases have “trivial” CP conserving values 0 or π . In certain one-parameter schemes, however, the Dirac phases $\delta = \pi/2$ or $3\pi/2$.

VII. In models in which the elements of the PMNS matrix are functions of two (angle, or one angle and one phase) or three (two angle and one phase) parameters, the Dirac phase δ satisfies a sum rule by which $\cos \delta$ is expressed in terms of the three neutrino mixing angles θ_{12} , θ_{23} , θ_{13} and one (or more) fixed (known) parameters θ^V which depend of the discrete symmetry G_f employed and on the residual symmetries G_e and G_ν [23, 24, 25, 26]. In this cases the J_{CP} factor which determines the magnitude of CP violation effects in neutrino oscillations, is also completely determined by the values of the three neutrino mixing angles and the symmetry parameter(s) θ_ν .

The predictions listed above, and therefore the respective models can be and will be tested in the currently running (T2K, NOvA) and future planned (JUNO, T2HK, DUNE) experiments.

2.2 Predictions for the Dirac CPV Phase

We will consider next for concreteness the approach followed in [23, 33, 26, 24] in which the requisite corrections to the underlying symmetry form of the neutrino mixing matrix are provided by the matrix U_e corresponding to G_f i) either broken to $G_e = Z_2$, or ii) completely broken, by the charged lepton mass term. In this case the PMNS matrix has the following general form [44]:

$$U = U_e^\dagger U_\nu = (\tilde{U}_e)^\dagger \Psi U_\nu^\circ P^\circ, \quad (2.3)$$

Here \tilde{U}_e is a 3×3 unitary matrix and Ψ is a diagonal phase matrix. The matrix \tilde{U}_e was chosen in [23, 33, 26] to have the following two forms:

$$\mathbf{A} : \tilde{U}_e = R_{23}^{-1}(\theta_{23}^e) R_{12}^{-1}(\theta_{12}^e); \quad \mathbf{B} : \tilde{U}_e = R_{12}^{-1}(\theta_{12}^e). \quad (2.4)$$

where θ_{12}^e and θ_{23}^e are free real angle parameters. These two forms appear in a large class of theoretical models of flavour and studies, in which the generation of charged lepton masses is an integral part (see, e.g., [30, 45]). The phase matrix Ψ in cases **A** and **B** is given by [23, 26]:

$$\mathbf{A} : \Psi = \text{diag}(1, e^{-i\psi}, e^{-i\omega}); \quad \mathbf{B} : \Psi = \text{diag}(1, e^{-i\psi}, 1). \quad (2.5)$$

The phases ω and/or ψ serve as a source for the Dirac CPV phase δ of the PMNS matrix and contribute to the Majorana CPV phases of the PMNS matrix α_{21} and α_{31} [23]. The diagonal phase matrix P° in eq. (2.3) contains two phases, ξ_{21} and ξ_{31} , which also contribute to the Majorana phases α_{21} and α_{31} , respectively.

Consider first the case of the five underlying symmetry forms of U_ν° - TBM, BM, GRA, GRB and HG - corrected by the matrix U_e , with the PMNS matrix given in eq. (2.3) and the matrices \tilde{U}_e

and Ψ as given in eqs. (2.4) and (2.5). In the considered setting the Dirac phase δ of the PMNS matrix satisfies the following sum rule [23]:

$$\cos \delta = \frac{\tan \theta_{23}}{\sin 2\theta_{12} \sin \theta_{13}} \left[\cos 2\theta_{12}^{\nu} + (\sin^2 \theta_{12} - \cos^2 \theta_{12}^{\nu}) (1 - \cot^2 \theta_{23} \sin^2 \theta_{13}) \right]. \quad (2.6)$$

Within the approach employed this sum rule is exact and is valid for any value of the angle θ_{23}^{ν} [24] (and not only for $\theta_{23}^{\nu} = -\pi/4$ of the five discussed symmetry forms of U_{ν}°). As we see, via the sum rule $\cos \delta$ is expressed in terms of the three neutrino mixing angles θ_{12} , θ_{23} , θ_{13} and one fixed (known) parameter θ^{ν} which depends on the underlying symmetry form (TBM, BM, GRA, GRB, HG) of the PMNS matrix. The difference between the cases **A** and **B** of forms of \tilde{U}_e in eq. (2.4) is, in particular, in the correlation between the values of $\sin^2 \theta_{23}$ and $\sin^2 \theta_{13}$ they lead to. In case **A** of \tilde{U}_e the values of $\sin^2 \theta_{23}$ and $\sin^2 \theta_{13}$ are not correlated and $\sin^2 \theta_{23}$ can differ significantly from 0.5 [23]. For the form **B** of \tilde{U}_e we have [23]:

$$\sin^2 \theta_{23} = \frac{1}{2} \frac{1 - 2 \sin^2 \theta_{13}}{1 - \sin^2 \theta_{13}} \cong \frac{1}{2} (1 - \sin^2 \theta_{13}). \quad (2.7)$$

Thus, in contrast to the case **A**, in case **B** the value of $\sin^2 \theta_{23}$ is correlated with the value of $\sin^2 \theta_{13}$ and as a consequence $\sin^2 \theta_{23}$ can deviate from 0.5 insignificantly - only by $0.5 \sin^2 \theta_{13}$.

Given the values of $\sin \theta_{23}$, $\sin \theta_{12}$, $\sin \theta_{13}$ and θ_{12}^{ν} , $\cos \delta$ is determined uniquely by the sum rule (2.6). This allows us to determine also $|\sin \delta|$ uniquely, but not $\text{sgn}(\sin \delta)$, which leads to a two-fold ambiguity in the predicted value of δ .

The fact that the value of the Dirac CPV phase δ is determined (up to an ambiguity of the sign of $\sin \delta$) by the values of the three mixing angles θ_{12} , θ_{23} and θ_{13} of the PMNS matrix and the value of θ_{12}^{ν} of the matrix U_{ν}° , eq. (2.1), is the most striking prediction of the models considered. This result implies that in the schemes under discussion, the rephasing invariant J_{CP} associated with the Dirac phase δ , eq. (1.5), is also a function of the three angles θ_{12} , θ_{23} and θ_{13} of the PMNS matrix and of θ_{12}^{ν} :

$$J_{\text{CP}} = J_{\text{CP}}(\theta_{12}, \theta_{23}, \theta_{13}, \delta(\theta_{12}, \theta_{23}, \theta_{13}, \theta_{12}^{\nu})) = J_{\text{CP}}(\theta_{12}, \theta_{23}, \theta_{13}, \theta_{12}^{\nu}). \quad (2.8)$$

This allows to obtain predictions for the possible values of J_{CP} for the different symmetry forms of U_{ν}° (specified by the value of θ_{12}^{ν}) using the current data on θ_{12} , θ_{23} and θ_{13} .

In [23], by using the sum rule in eq. (2.6), predictions for $\cos \delta$, δ and the J_{CP} factor were obtained in the TBM, BM, GRA, GRB and HG cases for the b.f.v. of $\sin^2 \theta_{12}$, $\sin^2 \theta_{23}$ and $\sin^2 \theta_{13}$. It was found that the predictions of $\cos \delta$ vary significantly with the symmetry form of U_{ν}° . For the b.f.v. of $\sin^2 \theta_{12} = 0.308$, $\sin^2 \theta_{13} = 0.0234$ and $\sin^2 \theta_{23} = 0.437$ found for NO spectrum in [12], for instance, one gets [23] $\cos \delta = (-0.0906)$, (-1.16) , 0.275 , (-0.169) and 0.445 , for the TBM, BM (LC), GRA, GRB and HG forms, respectively. For the TBM, GRA, GRB and HG forms these values correspond to $\delta = \pm 95.2^{\circ}$, $\pm 74.0^{\circ}$, $\pm 99.7^{\circ}$, $\pm 63.6^{\circ}$. For the b.f.v. given in Table 1 and obtained in the recent global analysis [2] one finds in the cases of the TBM, BM (LC), GRA, GRB and HG forms the values given in Table 3. Due to the different NO and IO b.f.v. of $\sin^2 \theta_{23}$, the predicted values of $\cos \delta$ and δ for IO spectrum differ (in certain cases significantly) from those for the NO spectrum.

Table 3: Predicted values of $\cos \delta$ and δ for the five symmetry forms, TBM, BM, GRA, GRB and HG, and \tilde{U}_e given by the form **A** in eq. (2.4), obtained using eq. (2.6) and the best fit values of $\sin^2 \theta_{12}$, $\sin^2 \theta_{23}$ and $\sin^2 \theta_{13}$ for NO and IO neutrino mass spectra from ref. [2].

Scheme	$\cos \delta$ (NO)	δ (NO)	$\cos \delta$ (IO)	δ (IO)
TBM	-0.16	$\pm 99^\circ$	-0.27	$\pm 106^\circ$
BM (LC)	-1.26	δ —unphysical	-1.78	δ —unphysical
GRA	0.21	$\pm 78^\circ$	0.24	$\pm 76^\circ$
GRB	-0.24	$\pm 105^\circ$	-0.38	$\pm 112^\circ$
HG	0.39	$\pm 67^\circ$	0.48	$\pm 62^\circ$

Two comments are in order. First, according to the results found in [2] and quoted in Table 1, the predicted values of δ lying in the first quadrant are strongly disfavored (if not ruled out) by the current data. Second, the unphysical value of $\cos \delta$ in the BM (LC) case is a reflection of the fact that the scheme under discussion with BM (LC) form of the matrix U_ν° does not provide a good description of the current data on θ_{12} , θ_{23} and θ_{13} [26]. Physical values of $\cos \delta$ can be obtained in the case of the NO spectrum, e.g., for the b.f.v. of $\sin^2 \theta_{13}$ if the value of $\sin^2 \theta_{12}$ ($\sin^2 \theta_{23}$) is larger (smaller) than the current best fit value¹⁰ [23, 33]. However, with the current b.f.v. of $\sin^2 \theta_{23}$ in the case of IO spectrum, the BM (LC) form is strongly disfavored.

The results quoted above imply [23] that a measurement of $\cos \delta$ can allow to distinguish between at least some of the different symmetry forms of U_ν° , provided θ_{12} , θ_{13} and θ_{23} are known, and $\cos \delta$ is measured, with sufficiently high precision¹¹. Even determining the sign of $\cos \delta$ will be sufficient to eliminate some of the possible symmetry forms of U_ν° .

These conclusions were confirmed by the statistical analyses performed in ref. [33] where predictions of the sum rule (2.6) for i) δ , $\cos \delta$ and the rephasing invariant J_{CP} using the “data” (best fit values and χ^2 -distributions) on $\sin^2 \theta_{12}$, $\sin^2 \theta_{13}$, $\sin^2 \theta_{23}$ and δ from [12], and ii) for $\cos \delta$, using prospective uncertainties on $\sin^2 \theta_{12}$, $\sin^2 \theta_{13}$ and $\sin^2 \theta_{23}$, were derived for the TBM, BM (LC), GRA, GRB and HG symmetry forms of the matrix U_ν° . Both analyses were performed for the case of NO neutrino mass spectrum. The results for the IO spectrum are similar. The aim of the first analysis, the results of which for J_{CP} are shown in Fig. 1, was to derive the allowed ranges for δ and J_{CP} , predicted on the basis of the current data on the neutrino mixing parameters for each of the symmetry forms of U_ν° considered (see [33] for details of the analysis). We have found [33], in particular, that the CP-conserving value of $J_{\text{CP}} = 0$ is excluded in the cases of the TBM, GRA, GRB and HG neutrino mixing symmetry forms, respectively, at approximately 5σ , 4σ , 4σ and 3σ C.L. with respect to the C.L. of the corresponding best fit values which all lie in the interval $J_{\text{CP}} = (-0.034) - (-0.031)$. The best fit value for the BM (LC) form is much smaller and close to zero: $J_{\text{CP}} = (-5 \times 10^{-3})$. For the TBM, GRA, GRB and HG forms at 3σ we have $0.020 \leq |J_{\text{CP}}| \leq 0.039$. Thus, for these four forms the CP violating effects in neutrino oscillations are predicted to be relatively large and observable in the T2HK and DUNE experiments [4, 5], and

¹⁰For, e.g., $\sin^2 \theta_{12} = 0.34$ allowed at 2σ by the current data, we have $\cos \delta = -0.943$. Similarly, for $\sin^2 \theta_{12} = 0.32$, $\sin^2 \theta_{23} = 0.41$ and $\sin \theta_{13} = 0.158$ we have [23]: $\cos \delta = -0.978$.

¹¹Detailed results on the dependence of the predictions for $\cos \delta$ on $\sin^2 \theta_{12}$, $\sin^2 \theta_{23}$ and $\sin^2 \theta_{13}$ when the latter are varied in their respective 3σ experimentally allowed ranges can be found in [33].

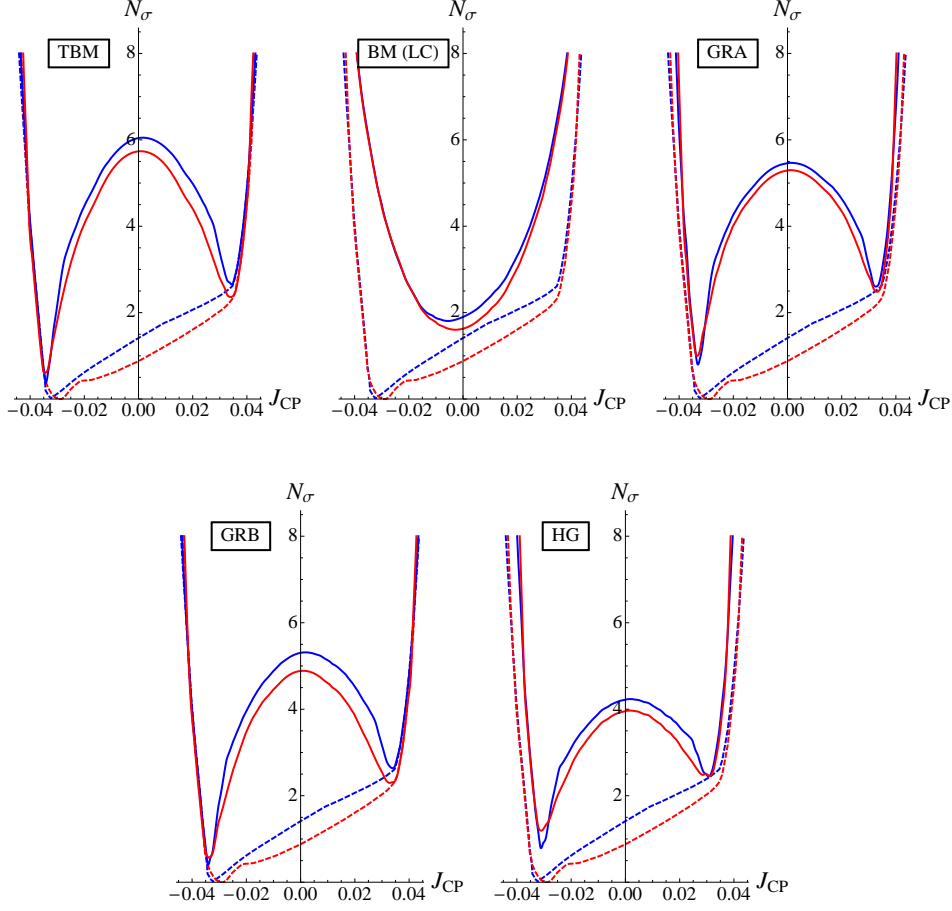


Figure 1: $N_\sigma \equiv \sqrt{\chi^2}$ as a function of J_{CP} . The dashed lines represent the results of the global fit [12], while the solid lines represent the results we obtain for the TBM, BM (LC), GRA (upper left, central, right panels), GRB and HG (lower left and right panels) neutrino mixing symmetry forms. The blue (red) lines are for NO (IO) neutrino mass spectrum. (From ref. [33].)

perhaps even in T2K experiment [46]. These conclusions hold if one used in the analysis the results on the neutrino mixing parameters and δ , obtained in the most recent global analysis [2].

In Fig. 2 (left panel) we present the results of the statistical analysis of the predictions for $\cos \delta$, namely the likelihood function versus $\cos \delta$ within the Gaussian approximation (see [33] for details) performed using the current b.f.v. of the mixing angles for NO neutrino mass spectrum given in ref. [12] and the prospective rather small 1σ uncertainties i) of 0.7% on $\sin^2 \theta_{12}$, planned to be reached in JUNO experiment [47], ii) of 3% on $\sin^2 \theta_{13}$, foreseen to be obtained in the Daya Bay experiment [48], and iii) of 5% on $\sin^2 \theta_{23}$, expected to be reached in the currently running and future planned long baseline neutrino oscillation experiments. In the proposed upgrading of the currently taking data T2K experiment [46], for example, θ_{23} is estimated to be determined with a 1σ error of 1.7° , 0.5° and 0.7° if the best fit value of $\sin^2 \theta_{23} = 0.50$, 0.43 and 0.60 , respectively. This implies that for these three values of $\sin^2 \theta_{23}$ the absolute (relative) 1σ error would be 0.0297 (5.94%), 0.0086 (2%) and 0.0120 (2%). This error on $\sin^2 \theta_{23}$ is expected to be further reduced in the future planned T2HK [5] and DUNE [4] experiments.

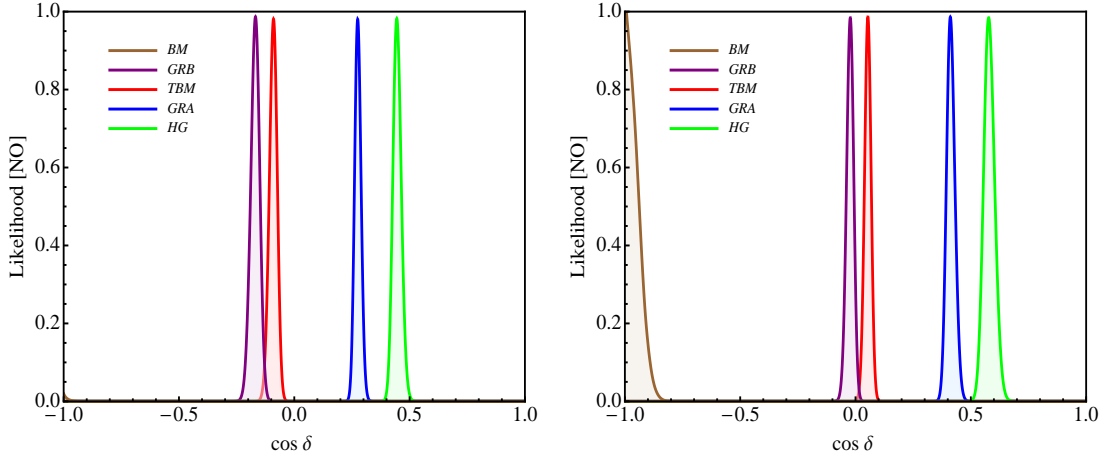


Figure 2: The likelihood function versus $\cos \delta$ for NO neutrino mass spectrum after marginalising over $\sin^2 \theta_{13}$ and $\sin^2 \theta_{23}$, for the TBM, BM (LC), GRA, GRB and HG symmetry forms of the mixing matrix U_ν° . The figure is obtained by using the prospective 1σ uncertainties in the determination of $\sin^2 \theta_{12}$, $\sin^2 \theta_{13}$ and $\sin^2 \theta_{23}$ within the Gaussian approximation. In the left (right) panel $\sin^2 \theta_{12}$ is set to its b.f.v. of [12] 0.308 (is set to 0.332), the NO best fit values of the other angles are taken from [12]. (From ref. [33].)

As we have already remarked, the BM (LC) case is very sensitive to the b.f.v. of $\sin^2 \theta_{12}$ and $\sin^2 \theta_{23}$ and is disfavored at more than 2σ for the b.f.v. found in [12] for the NO spectrum. This case might turn out to be compatible with the data for larger (smaller) measured values of $\sin^2 \theta_{12}$ ($\sin^2 \theta_{23}$). This is illustrated in Fig. 2 (right panel).

The measurement of $\sin^2 \theta_{12}$, $\sin^2 \theta_{13}$ and $\sin^2 \theta_{23}$ with the quoted precision will open up the possibility to distinguish between the BM (LC), TBM/GRB, GRA and HG forms of U_ν° . Distinguishing between the TBM and GRB forms seems to require unrealistically high precision measurement of $\cos \delta$ ¹². Assuming that $|\cos \delta| < 0.93$, which means for 76% of values of δ , the error on δ , $\Delta\delta$, for an error on $\cos \delta$, $\Delta(\cos \delta) = 0.10(0.08)$, does not exceed $\Delta\delta \lesssim \Delta(\cos \delta)/\sqrt{1 - 0.93^2} = 16^\circ(12^\circ)$. This accuracy is planned to be reached in the future neutrino experiments like T2HK (ESSvSB) [5, 49]. Therefore a measurement of $\cos \delta$ in the quoted range will allow one to distinguish between the TBM/GRB, BM (LC) and GRA/HG forms at approximately 3σ C.L. if the precision achieved on $\sin^2 \theta_{12}$, $\sin^2 \theta_{13}$ and $\sin^2 \theta_{23}$ is the same as in Figs. 2. We are performing a more detailed study of the possibility to distinguish between BM (LC), TBM/GRB, GRA and HG forms of U_ν° using the prospective data from DUNE and T2HK experiments [50].

In [24] we extended the analyses performed in [23, 33] by obtaining sum rules for $\cos \delta$ for U_{PMNS} having the general form given in eq. (2.3) and the following forms of \tilde{U}_e and U_ν° ¹³:

- C. $U_\nu^\circ = R_{23}(\theta_{23}^\nu)R_{12}(\theta_{12}^\nu)$ with $\theta_{23}^\nu = -\pi/4$ and θ_{12}^ν as dictated by TBM, BM, GRA, GRB or HG mixing, and i) $\tilde{U}_e = R_{13}^{-1}(\theta_{13}^e)$ ($\Psi = \text{diag}(1, 1, e^{-i\omega})$), ii) $\tilde{U}_e = R_{23}^{-1}(\theta_{23}^e)R_{13}^{-1}(\theta_{13}^e)$ ($\Psi =$

¹²Self-consistent models or theories of (lepton) flavour which lead to the GRB form of U_ν° might still be possible to distinguish from those leading to the TBM form using the specific predictions of the two types of models for the neutrino mixing angles. The same observation applies to models which lead to the GRA and HG forms of U_ν° .

¹³We performed in [24] a systematic analysis of the forms of \tilde{U}_e and U_ν° , for which sum rules for $\cos \delta$ of the type of eq. (2.6) could be derived, but did not exist in the literature.

$\text{diag}(1, e^{-i\psi}, e^{-i\omega})$), and iii) $\tilde{U}_e = R_{13}^{-1}(\theta_{13}^e)R_{12}^{-1}(\theta_{12}^e)$ ($\Psi = \text{diag}(1, e^{-i\psi}, e^{-i\omega})$);

- D. $U_\nu^\circ = R_{23}(\theta_{23}^\nu)R_{13}(\theta_{13}^\nu)R_{12}(\theta_{12}^\nu)$ with θ_{23}^ν , θ_{13}^ν and θ_{12}^ν fixed by arguments associated with symmetries, and iv) $\tilde{U}_e = R_{12}^{-1}(\theta_{12}^e)$ ($\Psi = \text{diag}(1, e^{-i\psi}, 1)$), and v) $\tilde{U}_e = R_{13}^{-1}(\theta_{13}^e)$ ($\Psi = \text{diag}(1, 1, e^{-i\omega})$).

The sum rules for $\cos \delta$ were derived first for $\theta_{23}^\nu = -\pi/4$ for the cases listed in point C, and for the specific values of (some of) the angles in U_ν° , characterising the cases listed in point D, as well as for arbitrary fixed values of all angles contained in U_ν° . In certain models with $\sin^2 \theta_{13}^\nu \neq 0$, $\sin^2 \theta_{23}$ is predicted to have specific values which differ significantly from those in case **B** [24]: $\sin^2 \theta_{23} = 0.455$; or 0.463; or 0.537; or 0.545, the uncertainties in these predictions being insignificant. Predictions for correlations between neutrino mixing angle values and/or sum rules for $\cos \delta$, which can be tested experimentally, were further derived in [25] for a large number of models based on $G_f = S_4, A_4, T'$ and A_5 and: i) $G_e = Z_2$ and $G_\nu = Z_n, n > 2$ or $Z_n \times Z_m, n, m \geq 2$; ii) $G_e = Z_n, n > 2$ or $G_e = Z_n \times Z_m, n, m \geq 2$ and $G_\nu = Z_2$; iii) $G_e = Z_2$ and $G_\nu = Z_2$; iv) G_e is fully broken and $G_\nu = Z_n, n > 2$ or $Z_n \times Z_m, n, m \geq 2$; v) $G_e = Z_n, n > 2$ or $Z_n \times Z_m, n, m \geq 2$ and G_ν is fully broken.

3. Outlook

The results obtained in refs. [23, 33, 24, 25, 30, 32, 11] and, e.g., in [29, 31, 34, 43] and in many other studies (quoted in the cited articles) show that a sufficiently precise measurement of the Dirac phase δ of the PMNS neutrino mixing matrix in the current and future neutrino oscillation experiments, combined with planned improvements of the precision on the neutrino mixing angles, can provide unique information about the possible discrete symmetry origin of the observed pattern of neutrino mixing and, correspondingly, about the existence of new fundamental symmetry in the lepton sector. Thus, these experiments will not simply provide a high precision data on the neutrino mixing and Dirac CPV parameters, but will probe at fundamental level the origin of the observed form of neutrino mixing. These future data will show, in particular, whether Nature followed the the discrete symmetry approach for fixing the values of the three neutrino mixing angles and of the Dirac (and Majorana) CP violation phases of the PMNS neutrino mixing matrix. We are looking forward to these data and to the future exciting developments in neutrino physics.

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