

Searches for Chiral Effects and Prospects for Isobaric Collisions at STAR/RHIC

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Searches for the chiral effects (the Chiral Magnetic Effect, the Chiral Vortical Effect and the Chiral Magnetic Wave) in heavy-ion collisions have been under theoretical and experimental investigations in recent years. A three-point correlator γ , has been used to measure electric charge separation across the reaction plane. Significant background sources arising from charge/momentum conservation coupled with the elliptic flow have also been identified. Isobaric collisions are proposed to potentially disentangle the flow related background and magnetic field induced charge separation signals in heavy-ion collisions. In these proceedings, we show recent STAR results on the γ correlator using charged hadrons and identified particles in 200 GeV Au+Au collisions. Moreover, the CME signal significance vs. background level is projected for the upcoming isobaric collisions (Zr+Zr and Ru+Ru) in 2018 at RHIC.

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1. Introduction

Quark chirality imbalance and a strong magnetic field can both be present in heavy-ion collisions, and their coupling can result in the chiral magnetic effect (CME) [1]. Searches for chirality effects using event-by-event analyses have been investigated theoretically and experimentally [2]. For Au+Au collisions at STAR, we use a three-point correlator γ (see Section 2.1) to measure the fluctuations in the charge separation magnitudes, and use difference between the opposite-sign (OS) and the same-sign (SS) charged-particle pairs to suppress common backgrounds. Recent studies have indicated that a considerable amount of flow-related background remains in the γ correlator [3, 4]. To quantify the range of the residual background, we use the κ_{112} parameter (see Section 2.1) to estimate the expected range of known background levels from AMPT simulations (v2.21, string melting, hadronic scattering turned on, charge not strictly conserved [5]).

The isobaric collisions, such as ${}^{96}_{44}$ Ru + ${}^{96}_{44}$ Ru and ${}^{96}_{40}$ Zr+ ${}^{96}_{40}$ Zr, are proposed to disentangle the flow-related background and the real CME signals [6]. In these two collision systems, the background (elliptic flow) will be similar, while the "driving-force" of the CME, the initial magnetic field, will differ significantly.

In Section 2.1 we describe the observables and analysis details. In Section 2.2, we present correlation measurements for charged hadrons and identified particles (π , K, p) from Au+Au collisions. For the upcoming isobaric collision program at STAR, we make a projection of the CME signal (see Section 3) based on the theoretical calculation of the initial magnetic field and eccentricity in Ru+Ru and Zr+Zr collisions.

2. Particle correlation in Au+Au collisions

2.1 Observables and data

The three-point γ correlator [6] is defined as follows,

$$\begin{split} \gamma &\equiv \cos(\phi_a + \phi_b - 2\Psi_{\rm RP})\rangle \tag{2.1} \\ &= \langle \cos(\phi_a + \phi_b - 2\Psi_2)\rangle / \langle \cos(2\Psi_2 - 2\Psi_{\rm RP})\rangle \\ &= \langle \cos(\phi_a + \phi_b - 2\phi_c)\rangle / v_{2,c}, \end{split}$$

which is sensitive to charge separation across the reaction plane in heavy-ion collisions. The ϕ_a and ϕ_b are the azimuthal angles of a correlated particle pair, and Ψ_2 and ϕ_c are used to approximate the reaction plane, Ψ_{RP} . The $v_{2,c}$ is the elliptic flow, v_2 , of the third particle *c*. A finite $\gamma_{OS} - \gamma_{SS}$, which quantifies the event-by-event fluctuation of charge distribution asymmetry, will indicate a possible charge separation across the reaction plane. Here "OS" denotes the particle pair with opposite charge sign and "SS" the same charge sign. For convenience, the notation $\Delta\gamma$ and $\gamma_{OS} - \gamma_{SS}$ will be used interchangeably in the rest of the proceedings.

The flow-related background is known to contribute to the γ correlation. To quantitatively estimate the flow background, a background model for γ and the two-particle correlator $\delta \equiv \langle \cos(\phi_a - \phi_b) \rangle$ has been proposed and a *H* correlator has been introduced [28]:

$$\Delta \delta \equiv \langle \cos(\phi_a - \phi_b) \rangle_{\text{OS-SS}} = \Delta F + \Delta H \tag{2.2}$$

$$\Delta \gamma \equiv \langle \cos(\phi_a + \phi_b - 2\Psi_{\rm RP}) \rangle_{\rm OS-SS} = \kappa v_2 \Delta F - \Delta H, \qquad (2.3)$$

where ΔF and ΔH are the amounts of charge separation resulting from the flow-related background and magnetic field, respectively. " Δ " refers to difference between "OS" and "SS" correlations. κ quantifies the coupling between v_2 and δ , which contaminates the γ measurements [7]:

$$\kappa = \frac{\Delta \gamma + \Delta H}{\nu_2 (\Delta \delta - \Delta H)}.$$
(2.4)

We introduce the κ_{112} parameter (this is the κ_K parameter in the talk slides) by setting ΔH to be zero for a maximum background scenario $\kappa_{112} \equiv \kappa|_{\Delta H=0} = \frac{\Delta \gamma}{v_2 \Delta \delta}$. The κ_{112} parameter can be compared with κ from background models such as AMPT simulations. Any excess of measured κ_{112} for particle pairs of interest over the expected range from background models could indicate the presence of physics beyond known background, *e.g.*, the CME.

For identified particle correlation measurements, ~200 million minimum-bias (MB) events of Au+Au collisions at $\sqrt{s_{NN}} = 200$ GeV from RHIC Run11 were used. Tracks reconstructed in the Time-Projection Chamber (TPC) [29, 30] with transverse momentum p_T from 0.15 to 2.0 GeV/c were used for the estimation of the second-order event plane. Particle species (π , K, p) were identified by applying cuts on the energy loss (dE/dx) in the TPC. Moreover, the particle velocity β measured via the Time-Of-Flight (TOF) [31] detector helps further select the desired particles.

2.2 Identified particle correlation in Au+Au collisions

Figure 1a shows the κ_{112} measurements for charged hadrons in 7.7, 19.6, 39, and 200 GeV Au+Au collisions [3]. We are using the AMPT ($|\eta| < 1.5$) $h^{\pm}h^{\pm}$ correlation result for Au+Au collisions at 200 GeV as a background reference without the CME and it shows a weak centrality dependence between values of 1 and 2. STAR data for energies from 19.6 to 200 GeV show very similar magnitudes of κ_{112} , well above the AMPT estimation of background levels for 200 GeV in mid-peripheral and mid-central collisions. Results at 7.7 GeV fall into the estimated background range. STAR data [8] showed that in peripheral collisions an additional short or intermediate-range background correlation also contributes to the γ , which may finally lead to the rise in κ_{112} in peripheral collisions. With additional data from STAR BES II (Beam Energy Scan II) program, the errors will be reduced as estimated in Figure 1b.

For correlations involving identified particles, Fig. 2 shows $\Delta \gamma$ measurements of particle pairs $(\pi \pi, pK, \pi K, pp, p\pi)$ as functions of centrality. The *pp* correlation magnitude is significantly larger than the other pairs. However, κ_{112} (see Fig. 2b) for these particle pairs indicates a very different behavior. Only $\pi \pi$ correlation has κ_{112} values larger than the AMPT background level (~1-2, estimated from charged hadrons) in mid-peripheral and mid-central collisions, while the other pairs are falling within the range of 1 to 2 except the *pp* pairs. κ_{112} for *pp* is found to be lower than 1 in those centrality bins. These results suggest that $\pi \pi$ correlation may dominate the $h^{\pm}h^{\pm}$ charge separation signal. The other pairs will require more simulation events (the current statistics are limited by speed of AMPT simulations) to determine the background. The charge separation and background level for *pp* pairs will require further investigation.

3. Isobaric collision projection

In spite of the observed sizable difference between same-sign and opposite-sign γ correlations in Au+Au 200 GeV collisions, the interpretation of the results still remain controversial. Studies [9,





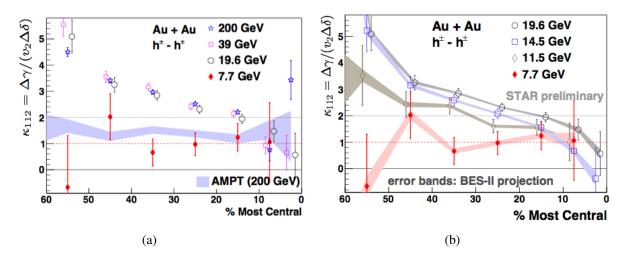


Figure 1: (a) κ_{112} for $h^{\pm}h^{\pm}$ in Au+Au collisions at 7.7, 19.6, 39, 200 GeV, and AMPT Au+Au 200 GeV simulations. (b) κ_{112} error projection (error bands) for the Beam Energy Scan phase II. The data points and error bars are from Beam Energy Scan Phase I.

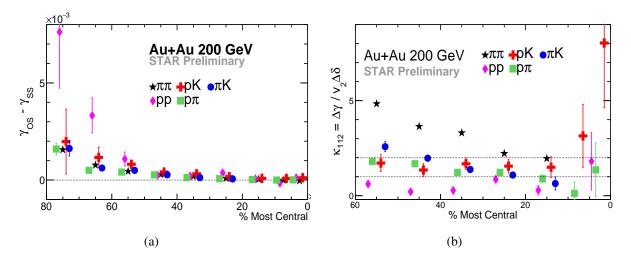


Figure 2: (a) $\gamma_{OS} - \gamma_{SS}$ and (b) κ_{112} for $\pi\pi$, *pK*, πK , *pp*, and *p* π as functions of centrality in Au+Au collisions at 200 GeV.

10, 11, 12, 13] show that some conventional physics mechanisms such as local charge conservation and transverse momentum conservation could couple with the elliptic flow to contribute to $\Delta\gamma$ signals.

One possible strategy to disentangle flow-related backgrounds and the real CME signal is to see how the signal varies in two collision systems where the background (elliptic flow) is similar, but the signal "driving forces" (initial magnetic fields) are significantly different. Isobaric collisions, for example, ${}^{96}_{44}$ Ru $+ {}^{96}_{44}$ Ru and ${}^{96}_{40}$ Zr, $+ {}^{96}_{40}$ Zr, have been proposed [14] to potentially achieve this goal. The Glauber model simulation demonstrates the multiplicity distributions in these two systems are almost identical [18] except for 5-10% central collisions, so the background contribution due to the multiplicity can be ignored for the CME analysis, in which only the 20-60% central

collisions will be our focus.

In the theoretical calculation [18], the spatial distribution of nucleons in either ${}^{96}_{44}$ Ru or ${}^{96}_{40}$ Zr in the rest frame can be written in the Woods-Saxon form:

$$\rho(r,\theta) \propto \frac{1}{1 + \exp[(r - R_0 - \beta_2 R_0 Y_2^0(\theta))/a]},$$
(3.1)

where R_0 and *a* represent the "radius" of the nucleus and the surface diffuseness parameter, respectively. The β_2 is the deformity of the nucleus. The e-A experiments [19, 20] and comprehensive model deduction [21] present two different sets of nucleus structure parameters for Ru and Zr as shown in Table. 1. In the rest of these proceedings, the set of parameters from e-A experiment will be referred to as case 1, and the one from the comprehensive model deduction will be referred to as case 2. The β_2 of Ru is larger than Zr (Ru is more deformed) in case 1, whereas it is the opposite in case 2. This systematic uncertainty impacts the multiplicity distribution negligibly, but may exert noticeable influence on the CME signal (via the magnetic field) and the v_2 driven backgrounds (via ε_2 , the initial spatial eccentricity of the participant zone).

Sources	Nucleus	R_0	а	β_2
Case 1	⁹⁶ ₄₄ Ru	5.085	0.46	0.158
(e-A Experiment)	$^{96}_{40}Zr$	5.02	0.46	0.08
Case 2	$^{96}_{44}Ru$	5.085	0.46	0.053
(Comprehensive Model Deduction)	$^{96}_{40}Zr$	5.02	0.46	0.217

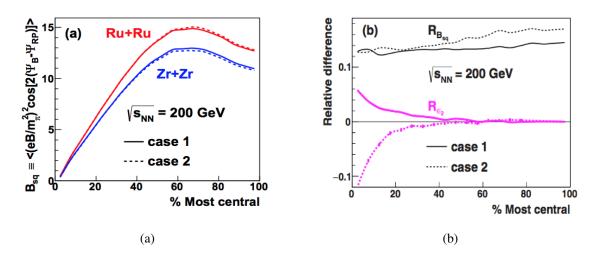


Table 1: Woods-Saxon Parameters.

Figure 3: (a) Initial magnetic field squared with correction from azimuthal fluctuation (B_{sq}) in Ru+Ru and Zr+Zr collisions at 200 GeV. (b) The relative difference of B_{sq} and ε_2 (initial eccentricity). Solid and dotted lines represent case 1 and case 2 parameter sets.

3.1 Initial magnetic field and eccentricity

The different charges carried by these two kinds of nuclei result in different initial magnetic fields (mostly generated by spectator protons) [22, 23], which finally produce distinctive magneticfield driven charge separation signals in these two collision systems. Figure 3 shows the theoretical calculation of the initial magnetic field squared with correction from azimuthal fluctuation of the magnetic field orientation $(B_{sq} = \langle (eB/2m_{\pi}^2)^2 \cos[2(\psi_{\rm B} - \psi_{\rm RP})] \rangle)$ as a function of centrality in Ru+Ru and Zr+Zr collisions at $\sqrt{s_{NN}} = 200$ GeV, as well as their relative difference. The B_{sq} quantifies the capability of the magnetic field to drive charge separation signal in the correlator γ [26, 27]. Most of the systematic uncertainty can be canceled by their relative difference ($R_{B_{eq}}$, in the following text, we use $R_F \equiv 2(F^{\text{Ru+Ru}} - F^{\text{Zr+Zr}})/(F^{\text{Ru+Ru}} + F^{\text{Zr+Zr}})$, where F can be B_{sq} , ε_2 or S to denote "relative difference" of this quantity between Ru+Ru and Zr+Zr collision systems). We can observe $R_{B_{sd}}$ increases slowly from 13% to 15% (18%) as the collision becomes more peripheral. The nucleus deformity uncertainty has more influence in peripheral collisions. In addition, the relative difference of the initial eccentricity ε_2 is also plotted on Fig. 3b, which is consistent with zero in peripheral collisions. In central collisions, however, these parameter sets for the two cases deviate significantly. Since the relative difference in v_2 should be very similar to R_{ε_2} , we can deduce that in these two collision systems, their v_2 values will stay almost the same for 20%-60% events. Also, the measurement of the relative difference of elliptic flow in central isobaric collisions may provide a definite answer to the question: which nucleus is more deformed?

3.2 Charge separation signal projection

Given the initial magnetic field and eccentricities, the magnetic field driven charge separation contribution to the charge separation ($S = \Delta \gamma \times N_{part}$) can be estimated. The signal in these two collision systems can be written [18]:

$$S^{\text{Ru+Ru}} = \bar{S}[(1-bg)(1+\frac{R_{B_{sq}}}{2}) + bg(1+\frac{R_{\varepsilon_2}}{2})]$$
(3.2)

$$S^{\text{Zr+Zr}} = \bar{S}[(1-bg)(1-\frac{R_{B_{sq}}}{2}) + bg(1-\frac{R_{\varepsilon_2}}{2})]$$
(3.3)

where $bg \in [0, 1]$ denotes the background contribution from elliptic flow and $\bar{S} = (S^{Ru+Ru} + S^{Zr+Zr})/2$. Then the relative difference of S can be expressed by:

$$R_S = (1 - bg)R_{B_{sq}} + bgR_{\varepsilon_2} \tag{3.4}$$

Figure 4a (Figure 4b) shows the projected signal S, with 1.2 billion minimum-bias events and 67% (80%) background level (percentage of $\Delta\gamma$ caused by the v_2) as a function of centrality in Ru+Ru and Zr+Zr collisions at 200 GeV. The relative difference is shown in Fig. 4c (Fig. 4d), where R_S and R_{ε_2} display different trends. This feature promises a chance to quantify the CME signals even with a high level of flow-related background. As a function of background level [18], the CME signal and its significance are plotted in Fig. 5. Given 1.2 billion minimum-bias events and 80% background, the significance of the CME signal reaches 5σ . With the same background level, but improved event plane resolution from a new STAR detector, the Event Plane Detector(EPD), the significance of the CME signal can be further raised to 7.8 σ .

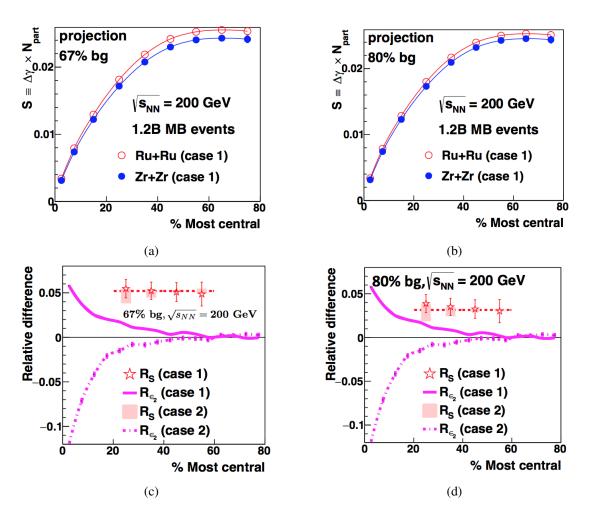


Figure 4: (a-b) The CME signal projection with 67%(a) / 80%(b) background level and 1.2 billion events in Ru+Ru and Zr+Zr collisions at 200 GeV. The nucleus structure parameters are from case1. (c-d) The Relative differences of the predicted signal and eccentricity between Ru+Ru and Zr+Zr 200 GeV collisions in both cases, with 67%(c) and 80%(d) background level.

4. Summary and outlook

We have presented STAR measurements of γ correlations and κ_{112} parameters for identified particle pairs ($\pi\pi$, pK, πK , pp, $p\pi$) in Au+Au collisions at $\sqrt{s_{NN}} = 200$ GeV. The κ_{112} values for particle pairs pK, πK , $p\pi$ are between 1 and 2. The κ_{112} parameter for $\pi\pi$ correlations has a similar magnitude to that for charged hadrons in Au+Au collisions at 200 GeV, which is larger than the background estimated from AMPT. The pp correlations yield a κ_{112} smaller than 1.

Theoretical calculation of the initial magnetic field and eccentricity of Ru+Ru and Zr+Zr is reported. The results show that even with background level as high as 80%, it is still possible to obtain a 5σ significance of the magnetic field induced charge separation signal from 1.2 billion minimum-bias events at STAR. With improved event plane resolution of the EPD, the statistical error may be further reduced. Meanwhile, isobaric collision program can also discern which of the two nuclei (Ru and Zr) is more deformed.

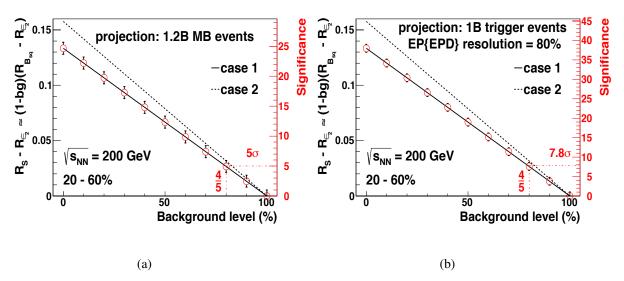


Figure 5: Projected signal magnitude and significance without (a) and with (b) EPD (Event-Plane Detector) as a function of background level.

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