

# **Recent SUSY results in ATLAS**

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Supersymmetry (SUSY) is considered one of the best motivated extensions of the Standard Model. It postulates a fundamental symmetry between fermions and bosons, and introduces a set of new supersymmetric particles at the electroweak scale. It addresses the hierarchy and naturalness problem, gives a solution to the gauge couplings unification, and offers a cold dark matter candidate. Different aspects of SUSY searches, using strong, electroweak, third generation production, *R*-parity violation models, and long lived particles are being studied at the LHC. An overview of most recent results in SUSY searches using Run 2 ATLAS data, at 13 TeV with  $36.1 \text{ fb}^{-1}$  of integrated luminosity, was presented.

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## 1. Introduction

Supersymmetry (SUSY) [1-6] is a well motivated extension of the Standard Model (SM). It postulates a fundamental symmetry between fermions and bosons, and introduces a new set of SUSY particles at the electroweak scale, where for each fermion (boson) in the SM there is a boson (fermion) SUSY partner. SUSY can offer solutions to a number of phenomena not explained by the SM. First, the radiative corrections to the Higgs boson mass become extremely large when a fermion couples to the Higgs. However, the radiative corrections of boson couplings to the Higgs are of opposite sign, and can cancel out with the fermionic contributions. SUSY naturally imposes this relation, and consequently allows large mass differences in the mass hierarchy of particles. Second, the Grand Unified Theories aim to provide one general description of electroweak and strong interactions. When experimentally measured values of electromagnetic, weak and strong coupling constants are extrapolated to high energies, these three never become equal. However, running of gauge couplings can be modified by introducing new physics between the electroweak and the Planck scale. When SUSY particles are included in the running of couplings, their evolution to the high mass scale brings to unification of gauge couplings. Third, observations of visible stars (or galactic gas) rotation speeds around the galactic center and their radial distance from it (i.e. galaxy rotation curves), show that they do not fall as the visible matter distribution prediction. They can be explained by introducing a new type of weakly interacting and non-relativistic matter, i.e. cold dark matter, and accommodate for its current estimate from astronomical measurements. SUSY offers a weakly interacting and massive particle candidate as a solution to this problem. In proton-proton collisions, SUSY particles are expected to be produced in pairs, and a typical signature has long chains of consecutive decays.



Figure 1: SUSY cross sections at 13 TeV [7] (a) and integrated luminosity [8] collected by the ATLAS experiment (b).

SUSY searches at the LHC [9] target a broad range of final states, where each analysis defines a set of selections with high sensitivity for considered models. They are grouped around production cross sections, as shown in the Figure 1 (a). For *R*-parity conserving models these include models of  $\tilde{q}$  and  $\tilde{g}$  production, third generation production, and electroweak production. Also *R*-parity violating, and models with long lived particles are considered. A typical SUSY analysis defines a set of selections on observables. The control regions (CR) are designed to be dominated by a certain SM background, and have no or very little expected signal events. They are used to estimate the number of events in the signal regions (SR) which are designed to maximize the signalto-background ratio, while the performance of the estimate is validated in the so-called validation regions (VR), which are designed to be similar to the SRs. More details on a typical SUSY analysis in ATLAS [10] can be found in Ref. [11]. In this summary the highlights of SUSY searches are presented, using 13 TeV p - p collision data, collected with the ATLAS experiment in 2015 and 2016 with integrated luminosity of 36.1 fb<sup>-1</sup>, as shown in the Figure 1 (b).

## 2. Squark and gluino production

The production of  $\tilde{q}$  and  $\tilde{g}$  (strong production) has the highest cross section at the LHC, and therefore represents a very important search. High sensitivity to a large number of SUSY models can be achieved targeting a final state with jets, possible leptons, and missing transverse momentum. There are two main approaches in the searches for strong production used in ATLAS. One is the conventional, where a typical discriminating variable is the effective mass ( $M_{eff}$ ), defined as a scalar sum of transverse momentum of all jets, possible leptons, and missing transverse momentum. It is a good measure of all activity in the event, and is expected to have larger values for the SUSY events, compared to the SM background. The second approach is done using Recursive Jigsaw Reconstruction (RJR) variables [12]. These are kinematic variables defined on the eventby-event basis, designed to use approximations of the rest frames of the invisible (SUSY) particles in each event. They have shown to have very good sensitivity for searches with compressed SUSY mass spectrum.

The analysis with a veto on a lepton, multiple jets and missing transverse momentum [13] targets the production of  $\tilde{q}$  and  $\tilde{g}$ . Due to a large production cross section and a very general selection, it has the sensitivity to a highest number of SUSY models. The main target of the analysis are the decays of  $\tilde{q}$  and  $\tilde{g}$  into quarks and  $\tilde{\chi}_1^0$ , the lightest symmetrical particle (LSP), and their one-step decays with intermediate  $W^{\pm}$  or a  $Z^0$  boson. A number of signal regions using the conventional and RJR analyses was designed to maximize the sensitivity to models with different  $\tilde{q}$ ,  $\tilde{g}$  and  $\tilde{\chi}_1^0$  masses. No significant excess was seen in any of the signal regions, as shown in the Figure 2. Interpretation was done in a number of SUSY models. In Figure 3 exclusion of  $\tilde{q}$  and  $\tilde{g}$  pair production models, with direct and one-step decays, can be seen. For each model, a significant improvement in sensitivity of around 500 GeV for  $\tilde{q}$  and  $\tilde{g}$  mass exclusion, compared to the previous analysis, using lower luminosity at 13 TeV p - p collisions, is achieved. The  $\tilde{q}$  mass is excluded up to around 1600 GeV, and  $\tilde{g}$  is excluded up to around 2 TeV.

The analysis with multiple jets and missing transverse momentum targets the strong production of models with long decay chains [14]. A typical targeted model is a  $\tilde{g}$  pair production with a twostep decay, where  $\tilde{g}$  decays into quarks with intermediate  $\tilde{\chi}_1^{\pm}$ , which decays into  $\tilde{\chi}_2^0$  with a  $W^{\pm}$ , and  $\tilde{\chi}_2^0$  decays into and  $Z^0$  and a  $\tilde{\chi}_1^0$  LSP. A number of signal regions was defined, using selections with large jet multiplicity, and *b*-jet multiplicity. No significant excess was observed in any of the signal regions, and exclusion limits were set on the considered models. A large improvement in the analysis sensitivity was achieved, of around 400 GeV for all  $\tilde{g}$  masses, and in the region of



Figure 2: Signal regions for the SUSY analysis with a veto on a lepton, multiple jets and missing transverse momentum using conventional (a) and RJR (b) variables [13].

compressed mass spectrum in the diagonal of the  $\tilde{g}$  vs  $\tilde{\chi}_1^0$  plane, compared to the previous 13 TeV search. Gluino masses up to around 1800 GeV were excluded, as shown in the Figure 4.

A distinctive signature of SUSY is with at least two same-sign leptons, and this analysis is targeted with a search with two same-sign or three leptons [15]. Signal regions were optimized for a number of *R*-parity conserving and *R*-parity violating models, which differ in the selection on a number of leptons, same-sign requirement and *b*-jet multiplicity. No significant excess was observed in any of the signal regions, and exclusion limits were set on a number of SUSY models. In the exclusion of the  $\tilde{g}$  pair production with a two-step decay, a large improvement in sensitivity for



Figure 3: Interpretation of simplified  $\tilde{q}$  (left) and  $\tilde{g}$  (right) production models, with direct (up) and one-step decays (down) using the analysis with a veto on a lepton, multiple jets and missing transverse momentum [13].





Figure 4: Analysis with multiple jets and missing transverse momentum [14]. No significant excess observed in any of the signal regions (a), and exclusion limits for the  $\tilde{g}$  pair production with a two-step decay (b).



the  $\tilde{g}$  masses of about 500 GeV was achieved, and in the region with compressed mass spectrum. Gluino masses up to around 1600 GeV were excluded, as shown in the Figure 5.

(b)  $\tilde{g}\tilde{g}$  two-step decay



In addition to searches with light (*u* and *d*) quarks, an analysis of SUSY models with decays of a  $\tilde{g}$  with third generation quarks (*b* and *t*) [16] needs to be considered. The signal regions were optimized for models with direct and one-step decays of  $\tilde{g}$  with *b* and *t* quarks. No significant excess was observed in any of the signal regions, and exclusion limits were set. Gluino masses up to about 1.8 - 2 TeV were excluded, as shown in the Figure 6.

In the Figure 7 a comparison of strong production searches is shown. Highest exclusion in the  $\tilde{g}$  masses of around 2 TeV is reached by the analysis with no leptons, 2-6 jets and missing transverse momentum, and it is comparable to the reach of the gluino mediated third generation searches, which in addition have a higher coverage toward the higher  $\tilde{\chi}_1^0$  masses.



(b)  $\tilde{g}\tilde{g}$  to bb direct decay

(c)  $\tilde{g}\tilde{g}$  to tt direct decay

Figure 6: Gluino mediated third generation search [16]. No significant excess in any of the signal regions, exclusion limits for direct gluino decays with b and t quarks (b).



Figure 7: Strong production analyses summary [17]. Exclusion curves shown for different  $\tilde{g}$  decays.

## 3. Third generation production

A strong motivation for third generation searches is given by the Higgs mass measurement [18, 19], which seems to be unnaturally light, and the existing  $\tilde{q}$  and  $\tilde{g}$  exclusion limits which reach up to the TeV scale. The dominant contribution to the Higgs mass comes from the divergent term of the top-quark. If SUSY exists, and  $\tilde{t}$  has masses  $\leq 1$  TeV, loop diagrams of the top-quark can cancel out to a large extent, which gives a natural solution to the mass hierarchy problem. In addition, large splitting between  $\tilde{t}_1$  and  $\tilde{t}_2$  can be achieved due to large t Yukawa couplings, so effects of the renormalization group equations are high for the third generation squarks, which brings to low  $\tilde{t}_1$  masses compared to first generation  $\tilde{q}$  masses. In this light, a dedicated search for light  $\tilde{t}$  is well motivated.

In order to obtain good sensitivity for a variety of typical signatures of  $\tilde{t}$  decays, in optimization a simplified model approach was used, where  $\tilde{\chi}_1^0$  LSP was assumed to be bino-like [20], see more details in the Figure 8. For mass differences of  $\Delta m = m_{\tilde{t}_1} - m_{\tilde{\chi}_1^0}$  higher than the stop mass  $m_{\tilde{t}_1}$ , a  $\tilde{t}$  pair production with the direct decays of  $\tilde{t}$  with a t quark was used in optimization. Similarly, one-step decays with b quarks and intermediate  $\tilde{\chi}_1^{\pm}$  decaying into a  $W^{\pm}$  and  $\tilde{\chi}_1^0$  LSP, were considered for the  $\Delta m > m_{W^{\pm}} + m_b$ . Further dedicated selections were optimized for even lower  $\Delta m$ , where the  $W^{\pm}$  decays off-shell, and for  $\tilde{t}$  decaying into a c quark and  $\tilde{\chi}_1^0$  LSP. No significant excess was seen in any of the signal regions, and exclusion limits were set. A large improvement in sensitivity, of about 250 GeV in the  $\tilde{t}$  mass for the low  $\tilde{\chi}_1^0$ , and strongly improved coverage for the lower  $\Delta m$  models, was achieved compared to 13 TeV searches with lower integrated luminosity.

In addition to using simplified models with a bino  $\tilde{\chi}_1^0$  LSP, more realistic mass spectrum



Figure 8:  $\tilde{t}$  pair production analysis strategy using simplified models [20] (a) and exclusion limits [17] (b).

and other LSP mixtures need to be considered for  $\tilde{t}$  pair production. For this, dedicated model specifications in Phenomenological Minimal SUSY Model (pMSSM) [22,23] were used, as shown in the Figure 9. Models motivated by gauge unification at the GUT scale predict a wino  $\tilde{\chi}_1^0$  LSP, and have  $\tilde{\chi}_1^{\pm}$  and  $\tilde{\chi}_2^0$  masses in between third generation squarks, and  $\tilde{\chi}_1^0$  LSP. Models motivated by obtaining natural SUSY, have a Higgsino LSP, and favor light  $\tilde{\chi}_1^{\pm}$ ,  $\tilde{\chi}_2^0$  and  $\tilde{\chi}_1^0$  LSP, with a large mass difference compared to  $\tilde{t}_1$ . Models that predict a dark matter candidate produced in the right amount favor a bino-Higgsino mixture, where third generation squarks have a higher mass compared to all neutralinos and charginos, with  $\tilde{\chi}_1^0$  even slightly lighter. Dedicated analyses were optimized to target these scenarios. No significant excess was seen in any of the searches, and





sparticle masses



Figure 10: Search for  $\tilde{b}$  pair production [21].

exclusion limits were set on these models [20]. These are the first exclusions of pMSSM inspired models in ATLAS, dedicated for  $\tilde{t}$  pair production, with these given motivations.

In addition to  $\tilde{t}$  pair production, a complementary search is done using the  $\tilde{b}$  quark production [21]. Typical final states in the analysis include a  $\tilde{b}$  pair production, with a direct decay into a b quark and  $\tilde{\chi}_1^0$  LSP, or  $\tilde{b}$  one-step decay modes via  $\tilde{\chi}_1^{\pm}$  with a  $W^{\pm}$  in the decay chain. For the direct production, an improvement in the sensitivity of about 100 GeV in the  $\tilde{b}$  mass is achieved, while for the combined direct and one-step decays in the  $\tilde{b}$  pair production an impressive 400 GeV improvement is achieved for the  $\tilde{b}$  mass exclusion, and significant improvement in sensitivity is achieved for the compressed mass spectrum towards the diagonal in the  $\tilde{b}$  vs  $\tilde{\chi}_1^0$  plane, compared to the previous searches. The exclusion limits of  $\tilde{b}$  reach about 900-1000 GeV for direct and combined decays, as shown in Figure 10.

#### 4. Electroweak production

If the masses of squarks and gluinos are significantly large, the production of  $\tilde{\chi}^0$ ,  $\tilde{\chi}_1^{\pm}$  and  $\tilde{\ell}$  can be dominant. As the limits of the strong production are reaching about 2 TeV, the studies of the electroweak production become well motivated. Typical searches consider final states with two or three leptons, with and without jets, in the  $\tilde{\chi}_1^{\pm}$  pair,  $\tilde{\ell}$  pair, and  $\tilde{\chi}_2^0 \tilde{\chi}_1^{\pm}$  production [24]. Significant improvement in the sensitivity to considered models was obtained, compared to previous searches, by using the analyses with multiple bin fits. The analyses with  $\tilde{\chi}_2^0 \tilde{\chi}_1^{\pm}$  production, with  $W^{\pm}$  and  $Z^0$  in the decay chains were using the binned invariant mass of two leptons. Models with  $\tilde{\ell}$  in the decay chain were using the binned information of the  $m_{T2}$  variable, which is a good measure of the missing transverse momentum in-balance in the event. No significant excess was observed in any of the signal regions, and exclusion limits were set, as shown in the Figure 11.

A comparison of recent searches using the electroweak production [17] is shown in Figure 12. A number of final states for the  $\tilde{\ell}$  and  $\tilde{\chi}_2^0 \tilde{\chi}_1^{\pm}$  production is targeted. The highest exclusion limits were set for the  $\tilde{\chi}_2^0 \tilde{\chi}_1^{\pm}$  production, with decays via  $\tilde{\ell}$  or  $\tilde{\nu}$  up to 1.16 TeV for  $\tilde{\chi}_2^0 / \tilde{\chi}_1^{\pm}$  masses.



Figure 11: Search for electroweak production using two or three leptons [24], using multibin fit for in  $m_{ll}$  (left) and  $m_{T2}$  (right) and interpretation (bottom).



Figure 12: Electroweak production analyses summary [17]. Exclusion curves shown for different productions and decays.

## 5. Long lived particles

The searches for SUSY models where sparticles are long lived, require dedicated analyses. The search for the the long lived  $\tilde{\chi}_1^{\pm}$  is targeted using the so-called disappearing-track analysis [17]. The production with long lived  $\tilde{\chi}_1^{\pm}$ ,  $\tilde{\chi}_1^0$  and initial-state-radiation jet, where  $\tilde{\chi}_1^{\pm}$  decays into a pion and  $\tilde{\chi}_1^0$  LSP, and similarly, a  $\tilde{g}$  pair production, where  $\tilde{g}$  can decay into two quarks and  $\tilde{\chi}_1^0$  LSP or into two quarks and long lived  $\tilde{\chi}_1^{\pm}$ , which decays into a pion and  $\tilde{\chi}_1^0$  LSP, were considered. The long lived  $\tilde{\chi}_1^{\pm}$  leaves hits in the inner parts of the detector (Pixel), and decays into a pion and  $\tilde{\chi}_1^0$  LSP, and has no counterparts of the track in the further parts of the tracking detector (SCT), as shown in the Figure 13 (a). The improved sensitivity of the analysis makes use of the implemented additional layer in the Pixel detector, and improves the sensitivity towards shorter life times of the long lived  $\tilde{\chi}_1^{\pm}$ . No significant excess was observed in any of the signal regions, and a comparison to previous searches [17] is shown in the Figure 13 (b).

## 6. *R*-parity violation

In addition to *R*-parity conserving models, the searches for *R*-parity violating models (RPV) need to be considered. Due to *R*-parity violating terms in the SUSY hyperpotential, in these models lepton and baryon number violation are allowed. This brings to a variety of final states not considered in standard searches, and dedicated analyses were set in place. A broad range of new RPV searches were considered compared to analyses with lower luminosity, but no significant excess was observed in any of the dedicated signal regions [26]. Exclusion limits were set, and for gluino pair production, with  $\tilde{g}$  decaying into a *t* and  $\tilde{t}$ , where  $\tilde{t}$  decays into a *b* and *s* quark,  $\tilde{g}$  is excluded up to masses of about 1600 GeV. Similarly, for a  $\tilde{g}$  pair production, with a decay into a



(a) "Disappearing track" analysis strategy.



Figure 13: Long lived  $\tilde{\chi}_1^{\pm}$  analysis strategy (a) [17] and excluion summary (b) [25].

pair of t quarks and intermediate  $\tilde{\chi}_1^{\pm}$  or  $\tilde{\chi}_1^0$ , where both of these sparticles decay into three quarks,  $\tilde{g}$  masses are excluded up to about 2 TeV, as shown in the Figure 14.

## 7. Conclusion

A broad range of improvements were implemented for the SUSY analyses using the 36.1 fb<sup>-1</sup> of ATLAS data. No significant excess in any of the signal regions was observed, and exclusion limits were set. In the  $\tilde{q}$  and  $\tilde{g}$  production, optimization for different regions of model parameter space brings to large improvements in the sensitivity of  $\tilde{g}$  and  $\tilde{q}$  masses, compared to searches with lower luminosity. In the searches for the third generation production, impressive improvements were obtained in the sensitivity of the analyses for standard and compressed mass regions. In addition, a new type of models were considered using the pMSSM parameter space selection, targeting well motivated LSP mixture scenarios. In the electroweak production, higher sensitivity was obtained



Figure 14: *R*-parity violating models interpretation [26],  $\tilde{g}$  pair production with  $\tilde{t}$  decaying into two quarks (left), and  $\tilde{g}$  pair production with  $\tilde{\chi}_1^0$  and  $\tilde{\chi}_1^{\pm}$  decaying to three quarks (right).

using improved searches and higher luminosity. For the long lived particles, limits on the long lived  $\tilde{\chi}_1^{\pm}$  were set, which benefited from the newly added Pixel detector layer. In addition to *R*-parity conserving models, limits were set on a broad range of *R*-parity violating models. In the Figure 15 a summary of recent SUSY searches in ATLAS is shown [17]. Future searches in ATLAS are planned to use more than two times higher luminosity, and a large number of improvements are set in place. On the one side in extending the sensitivity for uncovered and difficult parameterspace of existing searches, and on the other side in improving the searches by using a large number of new well motivated physics scenarios.

# Acknowledgements

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	TLAS SUSY Sear	ches*	- 95%	Ü	Ľ	wer Limits				ATLAS Preliminary
ב	ecernicer zur z Model	$e, \mu, \tau, \gamma$	Jets	$E_{\mathrm{T}}^{\mathrm{miss}}$	JL dt	њ <sup>-1</sup> ]	Mass limit	$\sqrt{s} = 7, 8$	$\frac{1}{3}$ TeV $\sqrt{s} = 13$ TeV	$\sqrt{v}s = t$ , o, 13 lev Reference
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	$qq, q \rightarrow qx_1$			SPI 1		4 =	00-0 OF	1.01 IEV	$m(x_1) < z = 0$ , $m(x^-) = z = 0$ , $m(x^-) = m(x^-) = 0$	10000
Sé	$qq, q \rightarrow qx_1$ (compressed)			201	- 00	4 ~	10 064	11 - 00 0	m(q)-m(x, 1) c>c>(1, x)-m(q)	
эц	$gg, g \rightarrow q\bar{q}\chi_1$		SIA( 0-7	Yes	30.1	202		2.UZ 16V	m(X_1)<200 GeV	1/12.02332
arc	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow qq\chi_1 \rightarrow qqW_1 \overset{*}{\rightarrow} qqW_1$	Ð	siel o-z	Yes	36.1	00		2.01 lev	m(X <sub>1</sub> )<200 GeV, m(X <sup></sup> )=0.5(m(X <sub>1</sub> )+m(g))	1/12.02332
əs	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\bar{q}(\ell\ell)\chi_1^{-1}$	ee, µµ	Z Jets	Yes	14.7	041		1.7 TeV	m(X <sub>1</sub> )<300 GeV,	1611.05791
ç ə	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow qq(\ell\ell/\nu\nu)\tilde{\chi}_{1}^{0}$	3 ε,μ	4 jets		36.1	001		1.87 TeV	$m(\vec{X}_1^0)=0$ GeV	1706.03731
vis	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow qq WZ \tilde{\chi}_1^0$	0	7-11 jets	Yes	36.1	001		1.8 TeV	m( $\tilde{\chi}_1^0$ ) <400 GeV	1708.02794
snj	GMSB (Ĩ NLSP)	$1-2 \tau + 0-1 \ell$	0-2 jets	Yes	3.2	001		2.0 TeV		1607.05979
ouj	GGM (bino NLSP)	2γ		Yes	36.1	00%		2.15 Te	V cπ(NLSP)<0.1 mm	ATLAS-CONF-2017-080
	GGM (higgsino-bino NLSP)	٨	2 jets	Yes	36.1	100		2.05 TeV	$m(\tilde{\chi}_1^0)=1700 \text{ GeV}, c_T(NLSP)<0.1 \text{ mm}, \mu>0$	ATLAS-CONF-2017-080
	Gravitino LSP	0	mono-jet	Yes	20.3	$F^{1/2}$ scale	865 GeV		$m(\tilde{G}) > 1.8 \times 10^{-4} \text{ eV}, m(\tilde{g}) = m(\tilde{g}) = 1.5 \text{ TeV}$	1502.01518
pe uə	$\tilde{a}\tilde{a} \xrightarrow{\tilde{a}} h \bar{h} \tilde{\chi}_{0}^{0}$	c	3.h	Yes	36.1	101		1.92 TeV	m(ž <sup>0</sup> )<600 GeV	1711.01901
ош 6 <sub>р.</sub>		0-1 e. u	3 h	Yes	36.1	2 304		1.97 TeV	m(2,0)<200.GeV	1711.01901
₽ ,E	1	-				5				
	$\tilde{b}_1 \tilde{b}_1, \tilde{b}_1 \rightarrow b \tilde{\chi}_1^0$	0	2b	Yes	36.1	$\tilde{b}_1$	950 GeV		$m(\tilde{\chi}_1^0) < 420 \text{ GeV}$	1708.09266
iou syu	$\tilde{b}_1 \tilde{b}_1, \tilde{b}_1 \rightarrow \tilde{\mathcal{K}}_1^{\pm}$	2 e,μ (SS)	1 b	Yes	36.1	$\tilde{b}_1$	275-700 GeV		$m(\tilde{\chi}_1^0) < 200 \text{ GeV}, m(\tilde{\chi}_1^{\pm}) = m(\tilde{\chi}_1^0) + 100 \text{ GeV}$	1706.03731
inct 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	$\tilde{t}_1 \tilde{t}_1, \tilde{t}_1 \rightarrow b \tilde{\chi}_1^{\pm}$	$0-2 e, \mu$	1-2 b	Yes 4	4.7/13.3	<i>ž</i> <sub>1</sub> 117-170 GeV	200-720 GeV		$m(\tilde{\chi}_{1}^{\pm}) = 2m(\tilde{\chi}_{1}^{0}), m(\tilde{\chi}_{1}^{0})=55 \text{ GeV}$	1209.2102, ATLAS-CONF-2016-077
ipo bs	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow Wb\tilde{\chi}_1^0$ or $k\tilde{t}_1^0$	$0-2e, \mu = 0$	2 jets/1-2 /	Yes 2	20.3/36.	<i>t</i> <sub>1</sub> 90-198 GeV	0.195-1.0 TeV		$m(\tilde{\chi}_1^0)=1$ GeV	1506.08616, 1709.04183, 1711.11520
ud u	$\tilde{n}\tilde{n}$ , $\tilde{n} \rightarrow c\tilde{\chi}_{1}^{0}$	0	mono-jet	Yes	36.1	$\tilde{t}_1$	90-430 GeV		$m(\tilde{t}_1)-m(\tilde{\chi}_1^0)=5 \text{ GeV}$	1711.03301
act ge	f, <i>i</i> ,	$2 e, \mu (Z)$	1b	Yes	20.3	Ĩı	150-600 GeV		m( $\tilde{\chi}_{0}^{0}$ )>150 GeV	1403.5222
qite عر	$\tilde{t}_2\tilde{t}_2, \tilde{t}_2 \rightarrow \tilde{t}_1 + Z$	3 e, μ (Z)	1b	Yes	36.1	ĩ,	290-790 GeV		$m(\tilde{\chi}_1^0)=0$ GeV	1706.03986
	$\tilde{p}\tilde{p}$ , $\tilde{p} \rightarrow \tilde{t}_1 + h$	$1-2 e, \mu$	4b	Yes	36.1	Ĩ2	320-880 GeV		m( $\tilde{\chi}_{0}^{0}$ )=0 GeV	1706.03986
		•								
	$\hat{\ell}_{L,R}\hat{\ell}_{L,R}, \hat{\ell} \rightarrow \ell \hat{X}_1$	2 e,µ	0	Yes	36.1	r 	90-500 GeV		$m(\tilde{\chi}_1^0)=0$	ATLAS-CONF-2017-039
	$\vec{X}_1^{\top}\vec{X}_1^{-}, \vec{X}_1^{\top} \rightarrow \ell \nu(\ell \tilde{\nu})$	2 e,µ	0	Yes	36.1	$\chi_1^{\pm}$	750 GeV		$m(\tilde{X}_1^0)=0, m(\tilde{\ell}, \tilde{v})=0.5(m(\tilde{X}_1^+)+m(\tilde{X}_1^0))$	ATLAS-CONF-2017-039
	$\check{X}_{1}^{\pm}\check{X}_{1}^{\dagger}/\check{X}_{2}^{2},\check{X}_{1}^{\dagger} \rightarrow \check{r}\nu(r\check{v}),\check{X}_{2}^{2} \rightarrow \check{r}r(v\check{v})$	21		Yes	36.1	$\chi_1^{\pm}$	760 GeV	1	$m(\tilde{X}_1^{''})=0, m(\tilde{r}, \tilde{v})=0.5(m(\tilde{X}_1^{''})+m(\tilde{X}_1^{''}))$	1708.07875
ise N	$\begin{array}{c} \chi_1^T \chi_2^2 \longrightarrow \ell_L \nu \ell_L \ell(\tilde{v}\nu), \ell \tilde{v} \ell_L \ell(\tilde{v}\nu) \\ \tilde{v} + \tilde{v} \\ \tilde{v} \\ \tilde{v} \\ \end{array}$	3 <i>e</i> ,µ	0 0	Yes	36.1	$X_{-1}^{-1}X_{2}^{-1}$	1.13 IeV	m(X <sub>1</sub> )=r	$n(\chi_{2}^{\circ}), m(\chi_{1}^{\circ})=0, m(\ell, \bar{v})=0.5(m(\chi_{1}^{\circ})+m(\chi_{1}^{\circ}))$	ATLAS-CONF-2017-039
diro El	$\begin{array}{ccc} X_1 X_2 \longrightarrow W X_1 Z X_1 \\ \vdots \\ $	1'a C-7	SIAC 2-0	Yes	30.1	$x_1, x_2$ $z \neq z_0$ $z = z_0$	280 GeV		$m(X_{\overline{1}}) = m(X_{\overline{2}}), m(X_{\overline{1}}) = 0, \ell$ decoupled	AILAS-CONF-2017-039
)	$\chi_1^-\chi_2^- \rightarrow W\chi_1^-h\chi_1^-, h \rightarrow bb/WW/\pi\pi/\gamma\gamma$	е, н, у	9 2 0	Yes	20.3	$x_1, x_2$ 2/0 G	eV	07	$m(\chi_1)=m(\chi_2), m(\chi_1)=0, \ell$ decoupled	011/0.1061
	$\dot{X}_{2}^{2}\dot{X}_{3}^{3}, \dot{X}_{2,3}^{2} \rightarrow \ell_{\mathrm{R}}\ell$	4 e,μ ~ .	0	Yes :	20.3	$\chi^{\chi_{2,3}}_{2,3}$	635 GeV	$m(\tilde{X}_2) = r$	$n(\vec{X}_3)$ , $m(\vec{X}_1')=0$ , $m(\vec{\ell}, \vec{v})=0.5(m(\vec{X}_2')+m(\vec{X}_1'))$	1405.5086
	GGM (wino NLSP) weak prod., $\vec{x_1} \rightarrow j$	$G = 1e, \mu + \gamma$		Yes	20.3	W 1	15-370 GeV		cr <1 mm	1507.05493
	GGM (bino NLSP) weak prod., $\tilde{\chi}_1^0 \rightarrow \gamma$	Ğ2γ		Yes	36.1	Ŵ	1.06 TeV		cr<1 mm	ATLAS-CONF-2017-080
	Direct $\tilde{Y}_{+}^{+}\tilde{Y}_{-}^{-}$ and long-lived $\tilde{Y}_{+}^{+}$	Disann trk	tet T	\ ∧oc	36.1	$\tilde{Y}^{\pm}$	460 GeV		m(Ž <sup>±</sup> )-m(Ž <sup>0</sup> )-160 MoV =(Ž <sup>±</sup> )-0.2 nc	1712 02118
	Diroct 2+25 and long-lived 2		ž .	200	100	71 74	ADE GeV		ווואן )-ווואן )~יסט אופע, גואן )=טיב ווא יינעידע יינעטע בועדע איז	1506.05332
	Ctable storned & D-badion		1.5 into	200		۲ ۲				
Sé pə,	Stable & Dedron	- i	sial c-i	res	5.12	900 H 19	800 GBV	11-1 01 1	$m(X_1)=100$ GeV, 10 $\mu s < \tau(g) < 1000$ s	
ol5 Vil∙		Ě			υ N N	1 00		1.58 leV	01	1606.05129
111 61	Metastable g H-hadron $20$	dE/dx trk		' :	2.2	00		1.57 TeV	$m(X_1^{\prime})=100 \text{ GeV}, r>10 \text{ ns}$	1604.04520
ed 107	Metastable $\tilde{g}$ R-hadron, $\tilde{g} \rightarrow qq \chi_1$	displ. vtx		Yes	32.8	50 - 0		2.37	<b>TeV</b> $\tau(\tilde{g})=0.17$ ns, $m(\chi_1^{\prime})=100$ GeV	1710.04901
1	GMSB, stable $\tilde{\tau}, \tilde{X}_{1}^{1} \rightarrow \tilde{\tau}(\tilde{e}, \tilde{\mu}) + \tau(e, \mu)$	η 2-Γ			19.1	¥.	537 GeV		10 <tanß<50< th=""><th>1411.6795</th></tanß<50<>	1411.6795
	$GMSB, \breve{X}_1^{\vee} \to \gamma G$ , long-lived $\breve{X}_1^{\vee}$	2 %		Yes	20.3	17 17	440 GeV		$1 < \tau(\tilde{\chi}_1^0) < 3$ ns, SPS8 model	1409.5542
	<i>§ĝ,X</i> 1→eev/eµv/µµv	aispi. ee/eµ/µ		'	20.3	$X_{\overline{1}}$	1.0 IEV		7 <cr(x<sub>1)&lt; 740 mm, m(g)=1.3 TeV</cr(x<sub>	1504.05162
	LFV $pp \rightarrow \tilde{v}_{\tau} + X, \tilde{v}_{\tau} \rightarrow e\mu/e\tau/\mu\tau$	<i>eμ</i> , <i>e</i> τ,μτ			3.2	$\tilde{v}_r$		1.9 TeV	$\lambda'_{311}$ =0.11, $\lambda_{132/133/233}$ =0.07	1607.08079
	Bilinear RPV CMSSM	2 e,μ (SS)	0-3 b	Yes	20.3	$\tilde{q}, \tilde{g}$	-	45 TeV	$m(\tilde{q})=m(\tilde{g}), c\tau_{LSP} < 1 mm$	1404.2500
	$\tilde{\chi}_1^+ \tilde{\chi}_1^-, \tilde{\chi}_1^+ \rightarrow W \tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow eev, e\muv, \mu\muv$	$4 e, \mu$		Yes	13.3	$\tilde{\chi}_1^{\pm}$	1.14 TeV		$m(\tilde{\chi}_{1}^{0}) > 400 \text{GeV}, \lambda_{12k} \neq 0 \ (k = 1, 2)$	ATLAS-CONF-2016-075
Λ	$\tilde{\chi}_1^+ \tilde{\chi}_1^-, \tilde{\chi}_1^+ \rightarrow W \tilde{\chi}_0^0, \tilde{\chi}_1^0 \rightarrow \tau \tau \nu_e, e \tau \nu_\tau$	$3 e, \mu + \tau$		Yes	20.3	$\tilde{X}_1^{\pm}$	450 GeV		$m(\tilde{\chi}_{1}^{0})>0.2\times m(\tilde{\chi}_{1}^{\pm}), \lambda_{133}\neq 0$	1405.5086
d٤	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow qq \tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow qqq$	0	i large-R je	ts -	36.1	001		1.875 TeV	$m(\tilde{\chi}_{1}^{0})=1075 \text{ GeV}$	SUSY-2016-22
ł	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow t t \tilde{\chi}_{1}^{0}, \tilde{\chi}_{1}^{0} \rightarrow q q q$	1 <i>e</i> ,μ 8-	10 jets/0-4	- 9	36.1	5 <u>0</u>		2.1 TeV	$\int m(\tilde{\chi}_1^0) = 1 \text{ TeV}, \lambda_{112} \neq 0$	1704.08493
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow \tilde{t}_1 t, \tilde{t}_1 \rightarrow bs$	1 <i>e</i> ,µ 8-	10 jets/0-4	- 9	36.1	00 0		1.65 TeV	m( $\tilde{t}_1$ )= 1 TeV, $\lambda_{323}$ ≠0	1704.08493
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow bs$	0	2 jets + 2 b		36.7	Ĩ,	100-470 GeV 480-610 GeV			1710.07171
	$\tilde{h}_1\tilde{h}_1, \tilde{h}_1 \rightarrow b\ell$	2 e,μ	2 b	•	36.1	Ĩı	0.4-1.	45 TeV	BR( $\tilde{t}_1 \rightarrow be/\mu$ )>20%	1710.05544
Other	Scalar charm, $\tilde{c} \rightarrow c \tilde{\chi}_1^0$	0	2 c	Yes	20.3	č	510 GeV		m( ${ ilde X}_1^0){<}200{ m GeV}$	1501.01325
'Only	a selection of the available mas	s limits on n	ew states	or					-	
pher.	nomena is shown. Many of the li	mits are bas	ed on	5		10-1	F		Mass scale [TeV]	
SIMP	Alfried models, c.i. reis. for the at	ssumptions i	naae.							

Figure 15: ATLAS SUSY searches summary [17].

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