

## Calibration and first measurements of *MuTe*: a hybrid Muon Telescope for geological structures

Jesús Peña-Rodríguez<sup>\*a</sup>, Adriana Vásquez-Ramírez<sup>a</sup>, José D. Sanabria-Gómez<sup>a</sup>, Luis A. Núñez<sup>a,b</sup>, David Sierra-Porta<sup>a,c</sup>, and Hernán Asorey<sup>d,e</sup>

<sup>a</sup> Escuela de Física, Universidad Industrial de Santander, Bucaramanga-Colombia

<sup>b</sup> Departamento de Física, Universidad de Los Andes, Mérida-Venezuela.

<sup>c</sup> Centro de Modelado Científico, Universidad del Zulia, Maracaibo-Venezuela.

<sup>d</sup> Departamento Física Médica, Centro Atómico Bariloche, Comisión Nacional de Energía Atómica, Bariloche-Argentina;

<sup>e</sup> Instituto de Tecnologías en Detección y Astropartículas (ITEDA), Buenos Aires-Argentina.

E-mail: [jesus.pena@correo.uis.edu.co](mailto:jesus.pena@correo.uis.edu.co),  
[adriana2168921@correo.uis.edu.co](mailto:adriana2168921@correo.uis.edu.co), [jsanabri@uis.edu.co](mailto:jsanabri@uis.edu.co),  
[lnunez@uis.edu.co](mailto:lnunez@uis.edu.co), [david.sierral@correo.uis.edu.co](mailto:david.sierral@correo.uis.edu.co),  
[asoreyh@cab.cnea.gov.ar](mailto:asoreyh@cab.cnea.gov.ar)

In this work, we describe the calibration and first measurements in the commissioning of *MuTe*, a hybrid Muon Telescope with two subdetectors –a scintillator hodoscope and a Water Cherenkov Detector (WCD)– for imaging the inner structures of Colombian volcanoes. The hodoscope estimates the trajectories of particles impinging on the front and rear panel, while the WCD acts as a calorimeter for the through going charged particle.

*MuTe* combines particle identification techniques so as to discriminate noise background from data. It filters the primary noise sources for muography, i.e., the EM-component ( $e^\pm$ ) of Extensive Air Showers (EAS) and scattered/upward-coming muons. The WCD identifies Electrons/positrons events by their deposited energy identifies, while scattered and backward muons are rejected using a pico-second Time-of-Flight(ToF) system.

Muon generated events were found in the deposited energy deposited range of ( $144\text{MeV} < E_d < 400\text{MeV}$ ), represent only about the 40% of the WCD-hodoscope acquired events. The other 60% of data is composed by ( $e^\pm$ ) events under 144 MeV and multiparticle events above 400 MeV. Subsequently, low-momentum muons ( $< 1 \text{ GeV}/c$ ), which are scattered by the volcano surface, measures a ToF  $> 3.3 \text{ ns}$  for traversing one meter length.

*36th International Cosmic Ray Conference -ICRC2019-  
July 24th - August 1st, 2019  
Madison, WI, U.S.A.*

<sup>\*</sup>Speaker.

## 1. Introduction

Muography is a non-invasive technique –spatial resolution in the order of tens of meters– for imaging anthropic and geologic structures implemented several decades ago [1, 2]. Recently it has been boosted with several new successful applications such as: detection of hidden materials in containers [3], archaeological building scanning [4, 5], nuclear plant inspection [6], nuclear waste monitoring, underground cavities [7], overburden of railway tunnels [8] and volcanology ([9] and references therein). For sometime now, research groups in Colombia have been exploring this technique which measures the variation in the atmospheric muon flux crossing geological edifices [10, 11, 12, 13, 14].

The flux variance between trajectories allows us to extract information about the inner density distribution of the scanned object. However, muography is affected by a density sub-estimation due to the recording of false-positive events, generated by three main phenomena: horizontal upward/backward-coming muons, low energy muons ( $< 1\text{GeV}$ ) scattered from the volcano surface [15, 16] and the EM-component ( $e^-$ ,  $e^+$ ) of EAS.

Various methods have been tried in different experiments for the removal of this background using: Time-of-Flight, ToF, systems for upward-coming muons rejection [17, 18], installation of absorbent panels for low energy muons filtering and increasing the number of sensitive panels for decreasing the probability of EAS-generated events [19].

Implementation of extra (absorbent or sensitive) panels affects the detector complexity and cost but, ToF systems arise as a reliable solution, filtering low energy muons with a discrimination resolution  $\sim 10\text{MeV}$ , with time differences of about 10ps.

In this paper we report the calibration method and first measurements of a Colombian hybrid Muon Telescope, MuTe, highlighting a new method for removing the background noise based on a ToF system which filters upward-coming & low energy muons and, also the inclusion of a WCD –based on the Latin American Giant Observatory, LAGO experiences – for the rejection of the EM-component of EAS.

## 2. The instrument

MuTe is a hybrid detector with two subdetectors: a scintillator hodoscope and a WCD. The hodoscope –with an angular resolution of 26.18 mrad for an inter-panel distance of 2.5 m– consists of two panels each of  $30 \times 30$  strips of  $120\text{ cm} \times 4\text{ cm} \times 1\text{ cm}$ , of polystyrene (Dow Styron 663) with an external coating of  $\text{TiO}_2$  and dopants (1% PPO, 0.03% POPOP), an absorption cut off  $\sim 40\text{ nm}$  and an emission maximum at 420 nm.

Each strip has a 1.8 mm hole for a wavelength shifting (WLS) multi-cladding fiber (Saint-Gobain BCF-92 with 1.2 mm diameter) with a core refraction index of 1.42, an absorption peak at 410 nm and an emission peak of 492 nm. Each WLS fiber is coupled, with a silicon photomultiplier (SiPM, Hamamatsu S13360-1350CS). The SiPM has a photosensitive area of  $1.3 \times 1.3\text{ mm}^2$ , 667 pixels, a fill factor of 74%, a breakdown voltage of  $53 \pm 5\text{ V}$ , a gain from  $10^5$  to  $10^6$  and a photon-detection efficiency of 40% at 450nm.

Figure 1 illustrates how each SiPM has a front-end electronics for polarisation and signal conditioning (i.e., pre-amplification for Signal-to-Noise Ratio enhancement). An ASIC MAROC3A

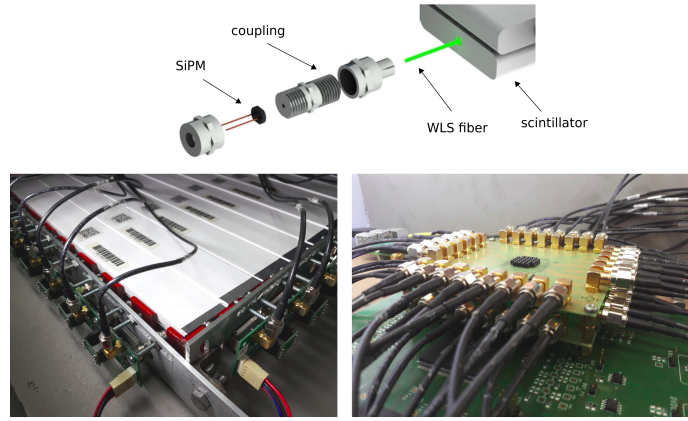


Figure 1: Assembling of the Saint-Gobain WLS fiber, mechanical coupling and the Hamamatsu SiPM in a scintillator bar (up). Structure of the rear scintillator panel, the electronics front-end and the transmission lines (down-left). MAROC3A-shield connector for 60 signals coming from a scintillator panel (down-right).

–whose slow control parameters are set by means of a FPGA Cyclone 3– individually amplifies and jointly discriminates the 60 signals from each panel. The data is managed by a Raspberry Pi 2 and stored in a central hard disk.

The cubic LAGO-WCD (120 cm side) has a Tyvek internal coating and an eight inch Hamamatsu R5912 photomultiplier tube (PMT) as the sensitive element, which provides two signals (anode and the last dynode), digitized by a 10 bits Fast Analog to Digital Converters (FADCs) with a sampling frequency of 40 MHz [20]. This sub-detector measures (see Figure 2) the energy loss ( $-dE/dx$ ) of the passing charged particles generating Cherenkov photons and allows the muon/background discrimination, by identifying the EM-component of cosmic ray showers and low energy muons ( $< 2\text{GeV}$ ) which are the main noise sources in muography.

A Time-to-Digital Converter (TDC) implemented on a Xilinx FPGA Spartan 6 to measure the crossing particle ToF. The TDC has two stages: a counter based on a 100 stages delay line with a resolution of 40ps and a course counter based on a Ring Oscillator whose resolution rises up to 5ns. The range of the TDC system is 90ns. The ToF is critical for filtering false-positive events related to upward-coming muons (i.e., negative ToF) and for estimating the particle momentum.

### 3. Trigger system

The MuTe trigger system has five hierarchical levels (T1, T2, T3, T4, and T5). T1 is formed by a discriminator (d1) in the MAROC3A when the scintillator pulse amplitude exceeds the threshold ( $V_{th0}$ ), which is established by a 10 bits Digital to Analog Converter (DAC).

Level T2 is an offline condition to single out orthogonal strip coincidences, i.e., an event candidate may trigger only one x-bar and y-bar to resolve the pixel positioning. T3 trigger occurs when two events, between the frontal and the rear panel, coincide within a time window of 7ns (shortest panel distance 2.5m) to 12 ns (largest panel distance 3.5m).

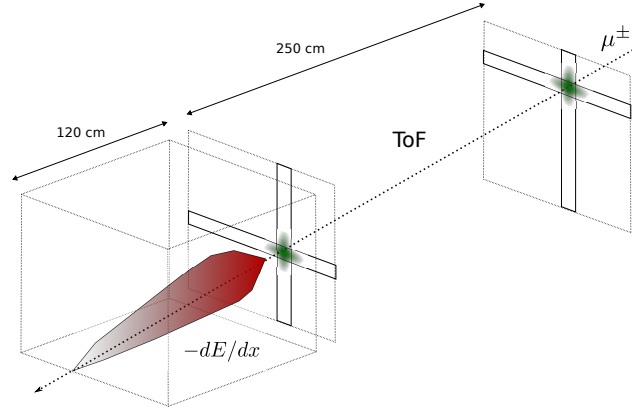


Figure 2: The general sketch of the MuTe detection scheme. The hodoscope estimates the event flux per trajectory depending on the fired pixels in the frontal and the rear panel. A ToF system measures the time taken for crossing particles subtracting the nano-scale time stamp recorded in both panels. The LAGO-WCD senses the Cherenkov photons generated in the water due to the charged particle interaction. The recorded photon yield is equivalent to energy loss.

Level T4 triggers a LAGO-WCD event when a PMT dynode pulse exceeds a 100 ADC threshold ( $\sim 100$  mV) in the LAGO-WCD discriminator. The digitized pulse is stored in a 12 bin vector with an inter-sample time step of 25 ns. Finally, when an event candidate raises the T3 and T4 flags, the T5 level occurs, containing all the particle information: trajectory, energy loss, and ToF.

#### 4. Hodoscope acceptance

The number of events detected by the hodoscope for a given direction  $r_{m,n}$  ( $m = i - k, n = j - l$  with  $\{i, j\}$  as the pixel coordinates for the frontal panel and  $\{k, l\}$  for the rear one) can be expressed as  $N(r_{m,n}, \Delta T) = I(r_{m,n}) \times \Delta T \times \mathcal{T}(r_{m,n})$  where  $I$  represents the flux of events given in  $\text{cm}^{-2}\text{sr}^{-1}\text{s}^{-1}$ ,  $\Delta T$  measures the period and the acceptance is  $\mathcal{T}(r_{m,n}) = S(r_{m,n}) \times \delta\Omega(r_{m,n})$ ;  $S$  is the detection surface and  $\delta\Omega$  the angular aperture. For two panels with  $N_x \times N_y$  pixels we can identify  $(2N_x - 1)(2N_y - 1)$  different discrete trajectories, where the angular aperture depends on the inter-panel distance  $D$ . The MuTe hodoscope can reconstruct 3481 trajectories with a solid angle in  $\theta_x = \theta_y = 0$  ( $r_{0,0}$ ) about  $8 \times 10^{-4}$  sr for an inter-panel distance of  $D = 134$  cm.

#### 5. Hodoscope and WDC calibrations

Several factors –material impurities during the scintillator production, WLS fiber assembling, SiPM coupling, signal conditioning, and transmission– can cause a variable detection rate in the scintillator bars. Thus, to reduce this channel response variability, the calibration starts with an equalization stage with a weighted gain to each bar. This is done by aiming the hodoscope to  $0^\circ$  zenith during 1 hour. The results, illustrated in Figure 3 display an average detection rate –after the equalization process– of  $836.3 \pm 96.3$  event/h, consistent with results reported in [19] for about

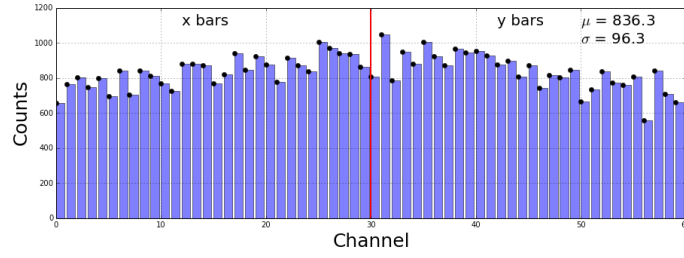


Figure 3: The results of the bar equalization process in an hour of data applying the triggers T1 and T2. The detection rate per bar is 836.3 event/h, and the variability is 96.3 event/h. The red line separates the  $x$  – bar and  $y$  – bar histograms.

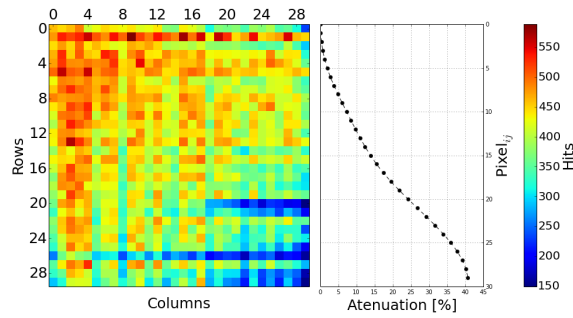


Figure 4: The number of hits detected by the frontal panel for 15 hours. The attenuation of the detection rate per pixel increases diagonally from the upper-left corner to the lower-right one. The most active pixel reaches 33.3 event/h whereas the least active only 13 events/h which means an attenuation about 40%.

1190 event/h. Furthermore, a bad optical coupling is exemplified in bar  $Y_{26}$  because of its low rate even after equalization.

The number of photons reaching the SiPM after a particle-scintillator interaction depends on the occurrence point. The higher the distance between the SiPM and the interaction point, the higher the attenuation in the photon yield. As a result of this effect, the scintillator panel is less sensitive (up to 40%) in the furthest corner from the SiPM placement as is shown in Figure 4.

On the other hand, the LAGO-WCD hardware calibration process consists of finding the optimum operation point of the PMT [21]. Such a point, located at the middle of a region called “plateau”, depends on the bias voltage  $V_b$  and the discrimination threshold.

The bias voltage was varied from 740 V to 1450 V in steps of 102 V, estimating the event rate for a 10 minutes time lapse. The optimum working point was selected using the minimization of the rate function derivative as follows

$$V_b^* = \arg \min \frac{d\Phi}{dV} \quad (5.1)$$

In figure 5 the solid lines represent the LAGO-WCD detection rate for a threshold of 110mV (black), 160mV (blue) and 210mV (red), while the dotted ones show their derivatives with the optimum operation voltages (shown by stars) at 1000V(110mV), 1096V(160mV), and 1109V(210mV).

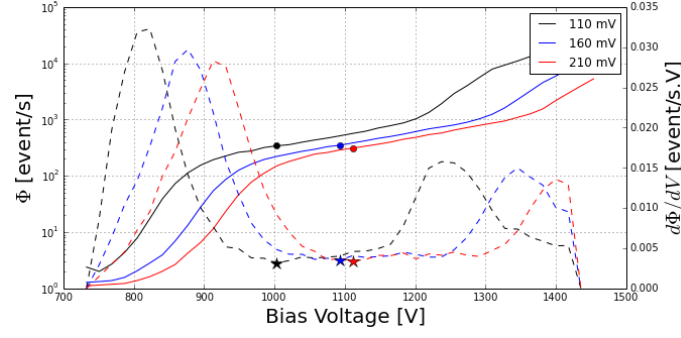


Figure 5: Detected event rate ranging from 740 V to 1450 V for a threshold of 110 mV (black), 160 mV (blue) and 210 mV (red). The optimum point for the LAGO-WCD functioning for each case is located at the star points on the plateau regions.

An important point is the estimation—from the charge histogram—of the energy loss of particles crossing the LAGO-WCD: a muon losses about 2 MeV/cm in water [22, 23].

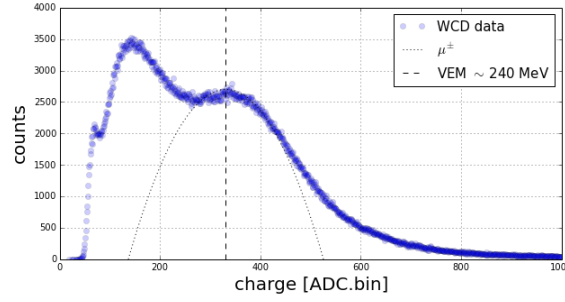


Figure 6: LAGO-WCD charge histogram. The highest hump represents the electromagnetic component of the detected events, with a mean value of 140 ADC.bin. The muonic component is fitted by a quadratic distribution with 1 VEM (331 ADC.bin) mean.

Figure 6 displays an one-hour charge histogram. It has two big humps; the first ( $\sim 140$  ADC.bin) due to the EM-component of EAS (i.e., electrons, positrons, and gammas) and the second one caused by the muonic component [22]. The deposited energy calibration is made taking into account the Vertical Muon Equivalent (VEM) point, which represents the deposited charge of the Cherenkov photons due to vertical muons ( $\sim 0^\circ$  zenith).

Subsequently, the energy loss per charge unit (ADC.bin) is defined as,

$$E_{loss} = (\text{ADC.bin}) \frac{240(\text{MeV})}{\text{VEM}_q(\text{ADC.bin})} \quad (5.2)$$

where (ADC.bin) is a unit of charge (i.e. equivalent to 0.72 MeV) and  $\text{VEM}_q$  is the charge at the muon hump mean (331.4 ADC.bin).

## 6. First measurements

After the calibration process, the MuTe hodoscope acquired data for 15 hours. Figure 7 displays the number of hits and the flux depending on the angular coordinates, inside the laboratory under a concrete shielding of about 30 cm. As expected, the number of detected events for orthogonal directions ( $r_{0,0}$ ) is greater than for others, reaching 68 hits in 15 hours. The flux has an annular modulation due to a dome shaped building placed above the MuTe. The maximum flux is about  $1.4 \times 10^{-4} \text{ cm}^{-2} \text{ sr}^{-1} \text{ s}^{-1}$ .

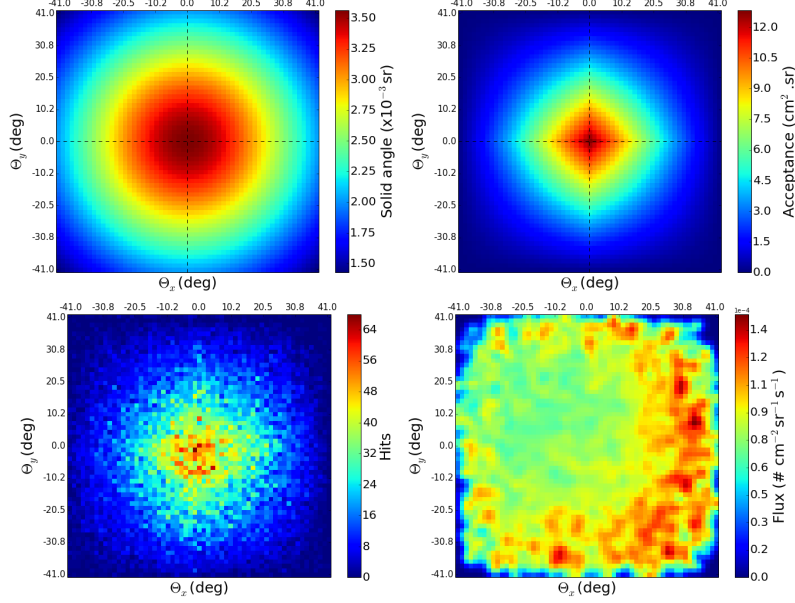


Figure 7: Angular properties of the hodoscope with  $N_x = N_y = 30$ ,  $d = 4 \text{ cm}$  y  $D = 134 \text{ cm}$ . Solid angle histogram as a function of the traversed trajectory (top-left). Hodoscope acceptance histogram (top-right). Number of events recorded by the hodoscope during 15 hours (bottom-left). Detected flux taking into account an ideal acceptance (bottom-right).

Figure 8 shows the charge histogram for all the events detected during 1 hour by the LAGO-WCD (blue) and for the events that triggered T5 (red) for a period of 10 hours. The T5 events represent only about 0.2% of the total number of events recorded by the whole LAGO-WCD. The low T5 flux is due to the diminished CR flux 1% at a zenith angle of  $70^\circ$  compared with  $0^\circ$  zenith. To test this background rejection capability of the LAGO-WCD, MuTe was pointed to  $90^\circ$  zenith with  $D = 250 \text{ cm}$ .

On the other hand, the muonic component of the T5 events, located around the VEM with a standard deviation of  $2\sigma$ , represents about the 40% of the T5 data. The background is composed of two regions: below 144 MeV is made of by the EM particles component and above 400 MeV dominated by the multiple-particle events.

## 7. Conclusions

This work reports the calibration and first measurements of a hybrid muon detector capable

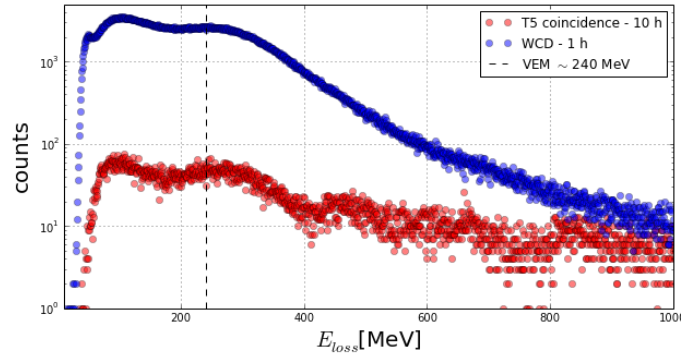


Figure 8: Deposited energy histogram for the T4 (blue) and T5 events (red) detected in the LAGO-WCD.

of rejecting the background signals in muography generated by the soft-component of EAS, low-momentum and upward-coming muons.

Muon generated events deposited in a energy range of ( $144\text{MeV} < dE/dx < 400\text{MeV}$ ), represent only about the 40% of the WCD-hodoscope acquired events. The other 60% of data is composed of ( $e^\pm$ ) events below 144MeV and multiparticle events above 400MeV. Subsequently, low-momentum muons ( $< 1 \text{ GeV}/c$ ), which are scattered by the volcano surface, are measured with a velocity of  $> 0.3 \text{ m/ns}$ .

Several reports show that background noise in muography –due to electrons, positron, and low-momentum muons– could rise as high as 60% of the acquired data [15, 16].

The Colombia MuTe combines a conventional muography method with particle identification techniques to discriminate noise background from data. This approach opens a window towards a new kind of hybrid muon telescopes with a better understanding of the background-origin and the phenomena behind the muography technique.

## Acknowledgments

The authors acknowledge the financial support of Departamento Administrativo de Ciencia, Tecnología e Innovación of Colombia (ColCiencias) under contract FP44842-082-2015 and to the Programa de Cooperación Nivel II (PCB-II) MINCYT-CONICET-COLCIENCIAS 2015, under project CO/15/02. We are particularly thankful to the Latin American Giant Observatory Collaboration and to Pierre Auger Observatory for their permanent support and inspirations.

## References

- [1] E.P. George *Cosmic rays measure overburden of tunnel*, *Commonwealth Engineer* **1** (1955) 455.
- [2] L.W. Alvarez et al., *Search for Hidden Chambers in the Pyramids*, *Science* **167** (1970) 832.
- [3] G. Blanpied et al., *Material discrimination using scattering and stopping of cosmic ray muons and electrons: Differentiating heavier from lighter metals as well as low-atomic weight materials*, *Nuclear Instruments and Methods in Physics Research Section A: Accelerators, Spectrometers, Detectors and Associated Equipment* **784** (2015) 352.



- [4] K. Morishima et al., *Discovery of a big void in khufu's pyramid by observation of cosmic-ray muons*, *Nature* **552** (2017) 386.
- [5] H. Gomez et al., *Studies on muon tomography for archaeological internal structures scanning*, in *Journal of Physics: Conference Series*, **718** (2016) 052016.
- [6] H. Fujii et al., *Performance of a remotely located muon radiography system to identify the inner structure of a nuclear plant*, *Progress of Theoretical and Experimental Physics* **2013** (2013) .
- [7] G. Saracino et al., *Imaging of underground cavities with cosmic-ray muons from observations at mt. echia (naples)*, *Scientific Reports* **7** (2017) .
- [8] L.F Thompson et al., *The application of muon tomography to the imaging of railway tunnels*, *ArXiv e-prints* (2019) [1906.05814].
- [9] H.K.M. Tanaka and L. Oláh, *Overview of muographers*, *Philosophical Transactions of the Royal Society A:Mathematical, Physical and Engineering Sciences* **377** (2019) 20180143.
- [10] H. Asorey et al., *Muon Tomography sites for Colombia volcanoes*, *ArXiv e-prints* (2017) [1705.09884].
- [11] H. Asorey et al., *Astroparticle projects at the eastern colombia region: facilities and instrumentation*, *Scientia et technica* **23** (2018) 391.
- [12] H. Asorey et al., *minimute: A muon telescope prototype for studying volcanic structures with cosmic ray flux.*, *Scientia et technica* **23** (2018) 386.
- [13] I.D. Guerrero et al., *Design and construction of a muon detector prototype for study the galeras volcano internal structure*, in *Journal of Physics: Conference Series*, **1247** (2019) 012020.
- [14] J.S. Useche-Parra, J.S. and C.A. Avila-Bernal, *Estimation of cosmic-muon flux attenuation by monserrate hill in bogota*, *Journal of Instrumentation* **14** (2019) P02015.
- [15] R. Nishiyama et al., *Experimental study of source of background noise in muon radiography using emulsion film detectors*, *Geoscientific Instrumentation, Methods and Data Systems* **3** (2014) 29.
- [16] H. Gómez et al., *Forward scattering effects on muon imaging*, *Journal of Instrumentation* **12** (2017) P12018.
- [17] J. Marteau et al., *Implementation of sub-nanosecond time-to-digital convertor in field-programmable gate array: applications to time-of-flight analysis in muon radiography*, *Measurement Science and Technology* **25** (2014) 035101.
- [18] L. Cimmino et al., *The muraves telescope front-end electronics and data acquisition*, *Annals of Geophysics* **60** (2017) S0104.
- [19] N. Lesparre et al., *Design and operation of a field telescope for cosmic ray geophysical tomography*, *Geoscientific Instrumentation, Methods and Data Systems* **1** (2012) 33.
- [20] M. Sofo-Haro et al., *The data acquisition system of the latin american giant observatory (LAGO)*, *Nuclear Instruments and Methods in Physics Research Section A: Accelerators, Spectrometers, Detectors and Associated Equipment* **820** (2016) 34.
- [21] H. Asorey et al., *Hardware-level calibration of the chitaga water cherenkov detector in the guane array for space weather study*, *Scientia et technica* **23** (2018) 563.
- [22] H. Asorey, *Los detectores cherenkov del observatorio pierre auger y su aplicación al estudio de fondos de radiación*, ph.d thesis, Instituto Balseiro, 2012.
- [23] A. Vásquez-Ramírez, *Estimación de la respuesta generada por el detector mute al paso de partículas cargadas*, master's thesis, Universidad Industrial de Santander, 2018.