

Constraints on neutrino millicharge and charge radius from neutrino-atom scattering

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We consider possible effects of neutrino electric charge (millicharge) and charge radius on the neutrino-atom interaction processes such as (i) atomic ionization by neutrino impact and (ii) coherent elastic neutrino-nucleus scattering. The bounds on the neutrino millicharge and charge radius that follow from, respectively, the GEMMA and COHERENT experiments are presented and discussed.

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1. Introduction

The development of our knowledge about neutrino masses and mixing provides a basis for exploring neutrino properties and interactions beyond the standard model (BSM). In this respect, the study of nonvanishing electromagnetic characteristics of massive neutrinos [1, 2] can help to constrain the existing BSM theories and/or to hint at new physics. The effects of neutrino electromagnetic properties can be searched in astrophysical environments, where neutrinos propagate in strong magnetic fields and dense matter, and in laboratory measurements of neutrinos from various sources. In the latter case, a very sensitive method is provided by the direct measurement of low-energy elastic neutrino scattering on atomic electrons and nuclei in a detector. In this contribution, we present our bounds on the neutrino millicharge and charge radii that have been derived from the data of the GEMMA [3] and COHERENT [4] scattering experiments, respectively, and included in the Particle Data Group's Review of Particle Physics [5].

2. Electromagnetic properties of massive neutrinos

There are at least three massive neutrino fields ν_i with respective masses m_i ($i = 1, 2, 3$), which are mixed with the three active flavor neutrinos ν_e, ν_μ, ν_τ . Therefore, the neutrino effective electromagnetic vertex, which in momentum-space representation depends only on the four-momentum $q = p_i - p_f$ transferred to the photon, can be presented as follows [1, 2]:

$$\Lambda_\mu(q) = (\gamma_\mu - q_\mu \not{q}/q^2) [f_Q(q^2) + f_A(q^2)q^2\gamma_5] - i\sigma_{\mu\nu}q^\nu [f_M(q^2) + if_E(q^2)\gamma_5]. \quad (2.1)$$

Here $\Lambda_\mu(q)$ is a 3×3 matrix in the space of massive neutrinos expressed in terms of the four Hermitian 3×3 matrices of form factors $f_Q = f_Q^\dagger, f_M = f_M^\dagger, f_E = f_E^\dagger$, and $f_A = f_A^\dagger$, where Q, M, E, A refer respectively to the real charge, magnetic, electric, and anapole neutrino form factors.

For the coupling with a real photon in vacuum ($q^2 = 0$) one has $f_Q^{fi}(0) = e_{fi}, f_M^{fi}(0) = \mu_{fi}, f_E^{fi}(0) = \varepsilon_{fi}$, and $f_A^{fi}(0) = a_{fi}$, where $e_{fi}, \mu_{fi}, \varepsilon_{fi}$ and a_{fi} are, respectively, the neutrino charge, magnetic moment, electric moment and anapole moment of diagonal ($f = i$) and transition ($f \neq i$) types. Even if the electric charge of a neutrino is zero, $f_Q(q^2)$ can still contain nontrivial information about neutrino electrostatic properties, namely the neutrino charge radius. The mean charge radius (in fact, it is the squared charge radius) of an electrically neutral neutrino is given by

$$\langle r_\nu^2 \rangle = \frac{1}{6} \left. \frac{df_Q(q^2)}{dq^2} \right|_{q^2=0}. \quad (2.2)$$

3. Elastic neutrino-electron scattering

Here we consider the process $\nu + e^- \rightarrow e^- + \nu$ where an ultrarelativistic neutrino with energy E_ν elastically scatters on an electron in a detector at energy transfer T . In the scattering experiments the observables are the kinetic energy T_e of the recoil electron and/or its solid angle Ω_e . From the energy-momentum conservation one gets

$$T_e = T, \quad \cos \theta_e = \left(1 + \frac{m_e}{E_\nu}\right) \sqrt{\frac{T}{T + 2m_e}}, \quad (3.1)$$

where θ_e is the angle of the recoil electron with respect to the neutrino beam and m_e is the electron mass. The cross section, which is differential with respect to the electron kinetic energy T_e , can be presented in the form of a sum of helicity-conserving (w, Q) and helicity-flipping (μ) components [6]:

$$\frac{d\sigma}{dT_e} = \frac{d\sigma_{(w,Q)}}{dT_e} + \frac{d\sigma_{(\mu)}}{dT_e}, \quad (3.2)$$

where $d\sigma_{(w,Q)}/dT_e$ is the electroweak cross section modified by the effect of the neutrino millicharge, charge radius and anapole moment, and $d\sigma_{(\mu)}/dT_e$ is the magnetic cross section due to the neutrino dipole magnetic and electric moments.

At small T_e values the contributions to the recoil-electron spectrum due to the weak, millicharge, and magnetic scattering channels exhibit qualitatively different T_e dependencies, namely

$$\mathcal{N}_{e^-}^{(w,Q)}(T_e) \propto \begin{cases} \text{const} & (e_\nu = 0), \\ \frac{2\pi\alpha^2}{m_e T_e^2} \left(\frac{e_\nu}{e_0}\right)^2 & (e_\nu \neq 0), \end{cases} \quad \text{and} \quad \mathcal{N}_{e^-}^{(\mu)}(T_e) \propto \frac{\pi\alpha^2}{m_e^2 T_e} \left(\frac{\mu_\nu}{\mu_B}\right)^2, \quad (3.3)$$

where α is the fine structure constant, e_ν and μ_ν are the neutrino (effective) millicharge and magnetic moment, and e_0 and μ_B are an elementary electric charge and a Bohr magneton, respectively. For the ratio \mathcal{R} of the millicharge and magnetic-moment contributions to the recoil-electron energy spectrum one thus has

$$\mathcal{R} = \frac{\mathcal{N}_{e^-}^{(Q)}(T_e)}{\mathcal{N}_{e^-}^{(\mu)}(T_e)} = \frac{2m_e}{T_e} \frac{(e_\nu/e_0)^2}{(\mu_\nu/\mu_B)^2}. \quad (3.4)$$

In case there are no observable deviations from the weak contribution to the electron spectrum it is possible to get the upper bound for the neutrino millicharge demanding that a possible effect due to e_ν does not exceed that due to the neutrino (anomalous) magnetic moment μ_ν . This implies that $\mathcal{R} < 1$ and from the relation (3.4), using the GEMMA data [3], namely the detector energy threshold ~ 2.8 keV and the μ_ν bound $\mu_\nu < 2.9 \times 10^{-11} \mu_B$, one obtains the following upper limit on the neutrino millicharge [7]:

$$|e_\nu| < 1.5 \times 10^{-12} e_0.$$

The e_ν range that expected to be probed in a few years with the GEMMA-II experiment (an effective threshold of 1.5 keV and the μ_ν sensitivity at the level of $1 \times 10^{-11} \mu_B$) is $|e_\nu| < 3.7 \times 10^{-13} e_0$.

4. Coherent elastic neutrino-nucleus scattering

Here we consider the process $\nu_\ell + a(Z, N) \rightarrow a(Z, N) + \nu_{\ell=e, \mu, \tau}$ where an ultrarelativistic neutrino with energy E_ν elastically scatters on an atomic nucleus, having Z protons and N neutrons, in a detector at energy-momentum transfer $q = (T, \vec{q})$. For a spin-zero nucleus and $T_a \ll E_\nu$, where $T_a = T$ is the nuclear recoil kinetic energy, the differential cross section due to the weak and charge-radius scattering channels is given by [6, 8]

$$\frac{d\sigma_{(w,r_\nu)}}{dT_a} \simeq \frac{G_F^2 M_a}{\pi} \left(1 - \frac{M_a T_a}{2E_\nu^2}\right) \left\{ [(g_V^p - \delta_{\ell\ell}) F_Z(|\vec{q}|^2) + g_V^n F_N(|\vec{q}|^2)]^2 + F_Z^2(|\vec{q}|^2) \sum_{\ell' \neq \ell} |\delta_{\ell\ell'}|^2 \right\}, \quad (4.1)$$

where M_a is the nuclear mass, $g_V^p = 1/2 - 2\sin^2\theta_W$ and $g_V^n = -1/2$ (the neglected radiative corrections are too small to affect the results). $F_{Z,N}(|\vec{q}|^2)$, such that $F_Z(0) = Z$ and $F_N(0) = N$, are the nuclear form factors, which are the Fourier transforms of the corresponding nucleon density distribution in the nucleus and describe the loss of coherence for $|\vec{q}|R \gtrsim 1$, where R is the nuclear radius. The effect of the neutrino charge radii is accounted for through

$$\delta_{\ell\ell'} = \frac{2}{3} m_W^2 \sin^2 \theta_W \langle r_{\nu_{\ell\ell'}}^2 \rangle, \quad \text{with} \quad \langle r_{\nu_{\ell\ell'}}^2 \rangle = \sum_{i,j} U_{\ell i}^* U_{\ell' j} \langle r_{\nu_{ij}}^2 \rangle,$$

where θ_W is the Weinberg angle, m_W is the W -boson mass, and U is the neutrino mixing matrix. The diagonal ($\ell = \ell'$) charge radii are already predicted in the standard model [9]:

$$\langle r_{\nu_e}^2 \rangle_{\text{SM}} = -0.83 \times 10^{-32} \text{ cm}^2, \quad \langle r_{\nu_\mu}^2 \rangle_{\text{SM}} = -0.48 \times 10^{-32} \text{ cm}^2, \quad \langle r_{\nu_\tau}^2 \rangle_{\text{SM}} = -0.30 \times 10^{-32} \text{ cm}^2. \quad (4.2)$$

However, the transition ($\ell \neq \ell'$) charge radii are essentially the BSM quantities.

The results of our fit of the time-dependent COHERENT data [4] are presented in Ref. [8]. In addition to the customary, diagonal charge radii, from the COHERENT data we have obtained for the first time limits on the neutrino transition charge radii [8]:

$$\left(|\langle r_{\nu_{e\mu}}^2 \rangle|, |\langle r_{\nu_{e\tau}}^2 \rangle|, |\langle r_{\nu_{\mu\tau}}^2 \rangle| \right) < (22, 38, 27) \times 10^{-32} \text{ cm}^2,$$

at 90% CL, marginalizing over reliable allowed intervals of the rms neutron radii $R_n(^{133}\text{Cs})$ and $R_n(^{127}\text{I})$. This is an interesting information on the BSM physics which can generate the neutrino transition charge radii [10].

References

- [1] C. Giunti and A. Studenikin, *Rev. Mod. Phys.* **87** (2015) 531.
- [2] C. Giunti, K. A. Kouzakov, Y.-F. Li, A. V. Lokhov, A. I. Studenikin and S. Zhou, *Ann. Phys. (Berlin)* **528** (2016) 198.
- [3] A. Beda *et al.* (GEMMA Collaboration), *Adv. High Energy Phys.* **2012** (2012) 350150.
- [4] D. Akimov *et al.* (COHERENT Collaboration), *Science* **357** (2017) 1123.
- [5] C. Patrignani *et al.* (Particle Data Group), *Chin. Phys. C* **40** (2016) 100001; M. Tanabashi *et al.* (Particle Data Group), *Phys. Rev. D* **98** (2018) 030001 and 2019 update.
- [6] K. A. Kouzakov and A. I. Studenikin, *Phys. Rev. D* **95** (2017) 055013.
- [7] A. Studenikin, *Europhys. Lett.* **107** (2014) 21001.
- [8] M. Cadeddu, C. Giunti, K. A. Kouzakov, Y.-F. Li, A. I. Studenikin and Y. Y. Zhang, *Phys. Rev. D* **98** (2018) 113010.
- [9] J. Bernabeu, L. G. Cabral-Rosetti, J. Papavassiliou and J. Vidal, *Phys. Rev. D* **62** (2000) 113012; J. Bernabeu, J. Papavassiliou and J. Vidal, *Phys. Rev. Lett.* **89** (2002) 101802; *Nucl. Phys.* **B680** (2004) 450.
- [10] H. Novales-Sanchez, A. Rosado, V. Santiago-Olan and J. Toscano, *Phys. Rev. D* **78** (2008) 073014.