

# Static Potential At Non-zero Temperatures From Fine Lattices

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We report on a preliminary study of static quark anti-quark potential at non-zero temperature in 2 + 1 flavor QCD using  $96^3 \times N_{\tau}$  lattices with lattice spacing a = 0.028 fm, physical strange quark mass and light quark masses corresponding to pion mass of about 300 MeV. We use  $N_{\tau} =$ 32, 24, 20 and 16 that correspond to temperature range T = 220 - 441 MeV. The in order to obtain the potential we calculate the Wilson line correlator in Coulomb gauge with additional HYP smearing to reduce the noise at large quark anti-quark separations. We apply 0, 5 and 10 steps of HYP smearing to ensure that there is no physical effect from over-smearing. At the two highest temperatures we also consider a noise reduction technique that is based on an interpolation in the spatial separation between the static quark and anti-quark.

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#### 1. Introduction

There has been a considerable interest in studying quarkonium properties at non-zero temperature since the seminal paper of Matsui and Satz, that suggested that suppression of quarkonium production in heavy ion collisions can signal creation of a deconfined medium [1]. In-medium properties of heavy quarkonium are encoded in the spectral functions, which in turn can be related to Euclidean time meson correlation functions, see e.g. Ref. [2] for a review. Reconstruction of the quarkonium spectral functions from a discrete set of data points turned out to be a very challenging task, see e.g. Refs. [3, 4, 5]. At zero temperature quarkonium properties can be estimated quite well using potential models with static quark antiquark potential calculated in lattice QCD, see e.g. Ref. [6]. It has been also proposed to calculate the quarkonium spectral function at non-zero temperature using potential model with complex potential from lattice QCD [7, 8]. The complex potential at non-zero temperature is also important for real time modeling of quarkonium production in heavy ion collisions [9, 10]. The complex potential at non-zero temperature can be defined in terms of Wilson loops or correlators of Wilson lines in Coulomb gauge,  $W(r, \tau, T)$  through their spectral decomposition [11, 12]

$$W(r,\tau,T) = \int_{-\infty}^{\infty} d\omega \rho_r(\omega,T) e^{-\omega\tau}.$$

Here *r* is the spatial separation between the static quark and antiquark and acts as label index for the spectral function of static  $Q\bar{Q}$  pair,  $\rho_r(\omega, T)$ . If the spectral function has a peak at some value of  $\omega$ , the peak position gives the real part of the potential ReV(r,T), while its width gives the imaginary part of the potential, ImV(r,T). We still need to reconstruct the spectral function in order to determine the potential, but the structure of this spectral function is much simpler than the structure of quarkonium spectral function, and thus should be easier to reconstruct. One needs a large number of data points in the time direction to accomplish this task. Current state of the art calculations in 2+1 flavor QCD mostly use lattices with temporal extent  $N_{\tau} = 12$  and some  $N_{\tau} = 16$  lattices that are available only at high temperatures [13]. In this contribution we report preliminary calculations of the Wilson line correlators calculated on very fine lattices with lattices spacing  $a^{-1} = 7.04$  GeV and temporal extent up to  $N_{\tau} = 32$ .

## 2. Lattice setup

We perform calculations at non-zero temperature in 2+1 flavor QCD on  $96^3 \times N_{\tau}$  lattices with lattice spacing  $a^{-1} = 7.04$  GeV. The strange quark mass,  $m_s$  is fixed to its physical value, while for the light quark mass we use  $m_l = m_s/5$ , which corresponds to the pion mass of about 300 MeV in the continuum limit. We consider  $N_{\tau} = 32$ , 24, 20 and 16, which correspond to T =220, 294, 353 and 441 MeV, respectively. We also generated additional T = 0 (64<sup>4</sup>) lattices to serve as a reference. We use temporal Wilson line correlators in Coulomb gauge instead of Wilson loops to obtain a better signal for the potential. However, this approach turns out to be insufficient for our very fine lattices. Therefore, the temporal Wilson lines are constructed from HYP smeared [14] temporal gauge links for  $N_{\tau} = 32$  and 24, as well as for the zero temperature case. We use 5 and 10 steps of HYP smearing. For the two highest temperature we consider an alternative noise reduction approach, which will be discussed below.

#### **3. Effective Masses**

Before reconstructing the spectral function  $\rho_r(\omega, T)$  we need to understand the  $\tau$ -dependence of the Wilson line correlators at different temperatures. This can be done in terms of the effective mass

$$m_{\rm eff}(r,\tau,T) = \partial_{\tau} \ln W(r,\tau,T)$$
$$\simeq \frac{1}{a} \ln \frac{W(r,\tau,T)}{W(r,\tau+a,T)}.$$

At zero temperature the effective mass approaches a plateau that corresponds to the ground state energy, i.e. the static potential, V(r). Our results for the effective masses at zero temperature are shown in Fig. 1. Since the potential has a divergent part proportional to the inverse lattice spacing



Figure 1: The effective masses for Wilson line correlator for r/a = 8 for 0, 5 and 10 steps of HYP smearing.

and the coefficient of this divergence depends on the number of smearing steps, the effective masses have been shifted by a constant to match them in the plateau region. As one can see from the figure for  $\tau/a > 3$  more smearing steps leads to faster approach to the plateau. However, for the two smaller  $\tau$  values this is not the case, and for the smallest  $\tau$ , the effective mass is the largest for 10 steps of HYP smearing. Without HYP smearing we lose the signal for  $\tau/a \gtrsim 12$  and therefore, we do not show the corresponding results. At larger spatial separations the situation is even worse, and no hint of a plateau can be obtained from unsmeared results. The HYP smearing distorts the *r*dependence of the static potential. But these distortions are limited to small distances. For distances r/a > 5, which are relevant for our study, the distortions due to HYP smearing are negligible compared to other sources of errors.

Next, we examine the temperature dependence of the effective masses. In Fig. 2 we show the effective mass for T = 0, T = 220 MeV and T = 293 MeV at two different distances, r/a = 8 and r/a = 16 and 10 steps of HYP smearing. The results for 5 steps of HYP smearing are similar. We see that for small  $\tau$  the temperature effects are quite small and grow with increasing  $\tau$ . At non-zero temperature the effective masses do not show a plateau at large  $\tau$  but an approximately linear decrease followed by a very rapid strongly non-linear drop as  $\tau$  approaches 1/T. These features follow from the general properties of the spectral function  $\rho_r(\omega, T)$  [13]. The approximately linear decrease in the effective masses is due to the fact that the ground state acquires a thermal width.





Figure 2: The effective masses at different temperature as function of  $\tau$  for r/a = 8 (left) and r/a = 16 (right). The results for 10 steps of HYP smearing are shown.

The sharp drop at large  $\tau$  is caused by the low  $\omega$  tail of the broadened peak [13]. The thermal effects are obviously larger at larger separations.

For  $N_{\tau} = 20$  we tried another procedure for noise reduction. It is based on the idea that for any fixed  $\tau$  the Wilson line correlator is a smoothly decaying function of r. The noise problem shows up at large values of r, where the correlation functions is available for many different spatial separations around a particular r value, that differ by fraction of the lattice spacing. Since the correlation function should be smooth in r (lattice artifacts at large distances are negligible) we can reduce the fluctuations in the data set by performing smooth interpolation in r in a narrow region of r. We used second order polynomial interpolations in intervals of r/a that are smaller than 0.9 for  $r/a \ge 20$ . The statistical fluctuations are largely reduced as the result of these interpolations. Using the corresponding interpolations for each  $\tau$  we can calculate the effective mass using the smoothened data for prescribed r values. In Fig. 3 we show the result of such an analysis. We indeed see that the interpolation in r really helps to reduce the statistical noise in the correlator. The right panel of this figure shows that the procedure is very effective for large r values,  $r/a \ge 20$ .



**Figure 3:** The effective masses for T = 353 MeV ( $N_{\tau} = 20$ ) for different separation. In the left panel we show the effective mass obtained using interpolation in *r* and compared to the standard result on the effective mass. In the right panel we show results obtained form the interpolation procedure for r/a = 20, 25, 30, 35 and 40.

#### 3.1 Subtracted correlators and comparison with the previous results

In the previous section we have seen that at very small  $\tau$  the temperature dependence of the Wilson line correlators is very small. This is expected to be due to the fact that at small Euclidean time the correlation function mostly receives contributions from the high  $\omega$  part of the spectral function, and this part of the spectral function is largely temperature independent, see e.g. Ref. [2]. Therefore, for the spectral function we can write [15, 16]



**Figure 4:** The effective masses corresponding to the subtracted correlators at two distances: rT = 1/4 (left) and rT = 1/2 (right). The top panel shows the results for T = 220 MeV, while the bottom panels show the results for T = 300 MeV. We compare our results with the previous  $N_{\tau} = 12$  and  $N_{\tau} = 16$  results without smearing, i.e. bare results.

$$\rho_r(\omega, T) = \rho_r^{\text{peak}}(\omega, T) + \rho_r^{\text{high}}(\omega), \qquad (3.1)$$

with  $\rho^{peak}(\omega, T)$  corresponding to the ground state peak in the spectral function that is broadened at non-zero temperature and may have a large low  $\omega$  tail [13]. This equation implies that

$$W(r,\tau,T) = W^{\text{peak}}(r,\tau,T) + W^{\text{high}}(r,\tau).$$
(3.2)

The zero temperature spectral functions has a delta peak corresponding to the ground state,  $\rho_r(\omega, T = 0) = A\delta(\omega - V^{T=0}(r)) + \rho_r^{\text{high}}(\omega)$ . Therefore, by fitting the ground state contribution at zero temperature and then subtracting it from  $W(r, \tau, T = 0)$  we can estimate  $W^{\text{high}}(r, \tau)$ . If we subtract  $W^{\text{high}}(r, \tau)$  from the finite temperature Wilson line correlator we obtain the subtracted correlator that is mainly sensitive to the temperature dependent peak part of the spectral function,  $\rho_r^{\text{peak}}(\omega, T)$ . Therefore, it is important to analyze the effective masses for the subtracted correlator. The corresponding results are shown in Fig. 4. We also compare our results with the ones obtained on  $N_{\tau} = 12$  and  $N_{\tau} = 16$  lattice without HYP smearing [13]. The effective masses show a linear decrease in  $\tau$  when the Euclidean time separation is far away from 1/T. For  $\tau$  close to 1/T we see

a faster non-linear dependence. These features of the subtracted effective masses are also present in the unsmeared  $N_{\tau} = 12$  and  $N_{\tau} = 16$  data and we find a very good agreement with the corresponding results, see Fig. 4. For large time the new results have significantly smaller errors as the result of HYP smearing. The significant smearing dependence of the effective masses at small  $\tau$ is largely reduced due to the subtraction, i.e. HYP smearing was mostly affecting the high energy part of the spectral function. The subtracted effective masses at small  $\tau$  show some non-monotonic behavior at very small  $\tau$ . This is due to the distortions due to HYP smearing, which affect slightly differently the zero and finite temperature correlator. The unsmeared data, on the other hand, show the expected behavior.

## 4. Conclusion

To obtain the the complex static  $Q\bar{Q}$  potential at non-zero temperature one needs to calculate Wilson loops or Wilson line correlators on fine lattices with large temporal extent. This, however, is challenging because of the poor signal to noise ratio. Therefore, in this contribution we explored two possible avenues for noise reduction. The first one is to use HYP smearing on the temporal links. We observed that using 5 and 10 steps of HYP smearing on lattices with  $a^{-1} = 7.04$  GeV at T = 220 MeV and T = 330 MeV provides a good signal for all relevant  $Q\bar{Q}$  separations. The distortions due to HYP smearing are limited to very small  $\tau$ , namely  $\tau/a < 5$ . Therefore, HYP smearing is a viable strategy for noise reduction. We also explored interpolation in r to reduces the noise. This approach worked well at the two highest temperature, T = 353 and 441 MeV. The subtracted effective masses agree well with the previous calculations performed on  $N_{\tau} = 12$  and  $N_{\tau} = 16$  lattices but have much smaller errors for large  $\tau$  values. Using the presented data on the effective masses of the Wilson line correlators we can obtain the complex static  $Q\bar{Q}$  potential if a suitable Ansatz for the spectral function is introduced. The corresponding analysis is currently underway.

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