

The IDM and THDMa - current constraints and future prospects

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We discuss two models that extend the Standard Model by additional particles, leading to additional scalar states, and also providing dark matter candidates. We briefly review the current constraints and comment on production cross sections at present and future collider facilities.

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1. Introduction

We discuss two new physics models that extend the Standard Model (SM) particle sector by additional scalars and also both provide a dark matter candidate. The models are confronted with current theoretical and experimental constraints. From the theoretical side, these include the minimization of the vacuum as well as the requirement of vacuum stability and positivity. We also require perturbative unitarity to hold, and perturbativity of the couplings at the electroweak scale.

Experimental bounds include the agreement with current measurements of the properties of the 125 GeV resonance discovered by the LHC experiments, as well as agreement with the null-results from searches for additional particles at current or past colliders. We also confront the models with bounds from electroweak precision observables via S, T, U [1–3], B-physics observables ($B \rightarrow X_s \gamma, B_s \rightarrow \mu^+ \mu^-, \Delta M_s$), as well as agreement with astrophysical observables (relic density and direct detection bounds). We use a combination of private and public tools in these analyses, where the latter include HiggsBounds [4], HiggsSignals [5], 2HDMC [6], SPheno [7], Sarah [8], micrOMEGAs [9, 10], and MadDM [11]. Experimental numbers are taken from [12, 13] for electroweak precision observables, [14] for $B_s \rightarrow \mu^+ \mu^-$, [15] for ΔM_s and [16] and [17] for relic density and direct detection, respectively. Bounds from $B \rightarrow X_s \gamma$ are implemented using a fit function from [18, 19]. Predictions for production cross sections shown here have been obtained using Madgraph5 [20].

2. Inert Doublet Model

The Inert Doublet Model (IDM) [21–23] is a two-Higgs-doublet Model (THDM) that obeys an exact Z_2 symmetry, under which the SM-like particles are even, while the novel scalars from the second (dark) doublet are odd. The exactness of the symmetry induces directly stability of the lightest new scalar state, providing a dark matter candidate. The model has been widely discussed in the literature, see e.g. [24–27]. The dark sector contains four new particles: H, A and H^\pm . After electroweak symmetry breaking, the model contains 5 free parameters which we chose to be $M_H, M_A, M_{H^\pm}, \lambda_2, \lambda_{345} \equiv \lambda_3 + \lambda_4 + \lambda_5$.

In [25] (see also [28]), benchmark points were suggested which cover dark scalar masses up to 1 TeV and are in agreement with the current constraints discussed above. We here focus on the investigation of these points at various current and future colliders, concentrating on pair-production processes

$$pp \rightarrow HA, HH^\pm, AH^\pm, H^+H^-; \ell^+ \ell^- \rightarrow HA, H^+H^-$$

for various collider options, where $\ell \in \{e, \mu\}$. For some of these, we also considered VBF-type production processes, where the above are augmented by 2 additional jets for pp and neutrinos for $\ell\ell$ machines. The different collider specifications considered here are given in table 1, where more details can be found in [27]. Figure 1 shows the cross-sections for various collider options and center-of-mass energies, as a function of the sum of the two pair-produced particles. The dashed lines signify the minimal cross sections needed for a generation of 1000 events assuming collider center-of-mass energies and luminosities as specified in table 1. In summary, we can state that already at the HL-LHC, all considered processes should be accessible (via the criterium above)

collider	cm energy [TeV]	$\int \mathcal{L}$	1000 events [fb]
HL-LHC	13/ 14	3 ab^{-1}	0.33
HE-LHC	27	15 ab^{-1}	0.07
FCC-hh	100	20 ab^{-1}	0.05
ee	3	5 ab^{-1}	0.2
$\mu\mu$	10	10 ab^{-1}	0.1
$\mu\mu$	30	90 ab^{-1}	0.01

Table 1: Parameter used for the different collider options discussed in the text. More details, including references, can be found in [27].

up to a mass sum of 2 TeV apart from AA pair-production. The latter process depends on the parameter $\bar{\lambda}_{345} \equiv \lambda_3 + \lambda_4 - \lambda_5 = \lambda_{345} - 2 \frac{M_H^2 - M_A^2}{v^2}$ which determines the hAA coupling, i.e. also mass-differences between A and H are important. Furthermore, λ_{345} is highly constrained by direct detection [24–27, 29]. For AA production, therefore, the mass scale, defined as the sum of produced particle masses, can be up to 600 GeV at the HL-LHC and 2 TeV at the FCC-hh. At lepton colliders, on the other hand, the VBF channel is needed in order to access this process, where mass scales are around 600 GeV/ 1.4 TeV/ 2 TeV for CLIC with a center-of-mass energy of 3 TeV/ a muon-collider with a center-of-mass energy of 10 TeV/ 30 TeV, respectively.

3. THDMa

The THDMa is a type II two-Higgs-doublet model that is extended by an additional pseudoscalar mixing with the "standard" pseudoscalar A of the THDM. In the gauge-eigenbasis, the additional scalar serves as a portal to the dark sector, with a fermionic dark matter candidate, denoted by χ . This model has e.g. been discussed in [30–36].

The model features the following particles in the scalar and dark matter sector: h, H, H^\pm, a, A, χ . It depends on in total 12 additional parameters

$$v, m_h, m_H, m_a, m_A, m_{H^\pm}, m_\chi; \cos(\beta - \alpha), \tan\beta, \sin\theta; y_\chi, \lambda_3, \lambda_{P_1}, \lambda_{P_2},$$

where v and either m_h or m_H are fixed by current measurements in the electroweak sector. We refer the reader to the above references for more details.

We here report on results of a scan that allows all of the above novel parameters float in specific predefined ranges [36]. In such a scenario, it is not always straightforward to display bounds from specific constraints in 2-dimensional planes. Two examples where this is possible are given in figure 2. In the first plot, we show bounds in the $(m_{H^\pm}, \tan\beta)$ plane from B-physics observables. The result is similar in a simple THDM, and shows that in general low masses $m_{H^\pm} \lesssim 800 \text{ GeV}$ as well as values $\tan\beta \lesssim 1$ are excluded. The second plot features the relic density as a function of the mass difference $m_a - 2m_\chi$. Here, a behaviour can be observed that is typical in many models with dark matter candidates: in the region where this mass difference remains small, relic density annihilates sufficiently to stay below the observed relic density bound. On the other hand, too large differences lead to values $\Omega h_c \gtrsim 0.12$ and therefore are forbidden from dark matter considerations.

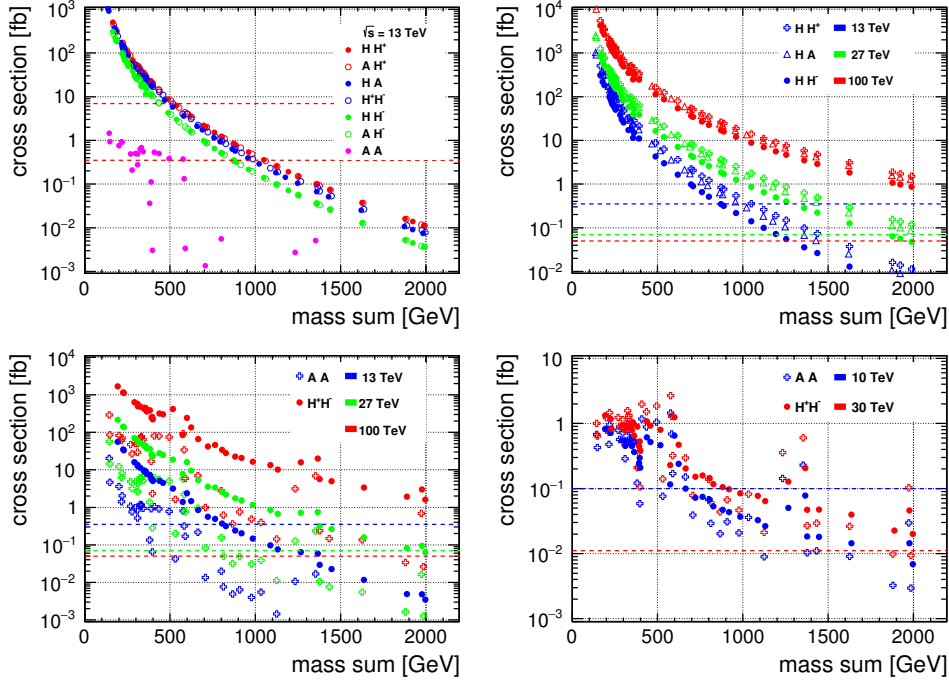


Figure 1: Predictions for production cross sections for various processes and collider options. *Top left:* Predictions for various pair-production cross sections for a pp collider at 13 TeV, as a function of the mass sum of the produced particles. *Top right:* Same for various center-of-mass energies. *Bottom left:* VBF-type production of AA and H^+H^- at various center-of-mass energies for pp colliders. *Bottom right:* Same for $\mu^+\mu^-$ colliders. Taken from [27]. The lines correspond to the cross-sections required to produce at least 1000 events using the respective design luminosity.

Finally, we can investigate which cross-section values would still be feasible for points that fulfill all constraints [36] at e^+e^- colliders. We are specifically interested in signatures that include missing energy and therefore do not exist in a THDM without a portal to the dark sector. Due to alignment, processes like $e^+e^- \rightarrow hA, h\bar{A}$ are suppressed, which makes $e^+e^- \rightarrow HA, H\bar{A}$ the most interesting channel featuring novel signatures. Due to the interplay of B-physics and electroweak constraints, such points typically have mass scales $\gtrsim 1$ TeV. We therefore consider production cross sections for an e^+e^- collider with a center-of-mass energy of 3 TeV. The corresponding production cross sections are displayed in figure 3, where we give predictions for $t\bar{t}t\bar{t}$ and $t\bar{t} + \cancel{E}$ final states using a factorized approach. We see that there is a non-negligible number of points where the second channel is dominant. A "best" point with a large rate for $t\bar{t} + \cancel{E}_\perp$ has been presented in [36].

4. Conclusion and Outlook

In this work, we presented two models that extend the particle content of the SM and also provide at least one dark matter candidate. We have presented production cross sections for various standard pair-production modes within these models; for the IDM, we have given an estimate of mass range reachability based on a simple counting criterium. A more dedicated investigation of

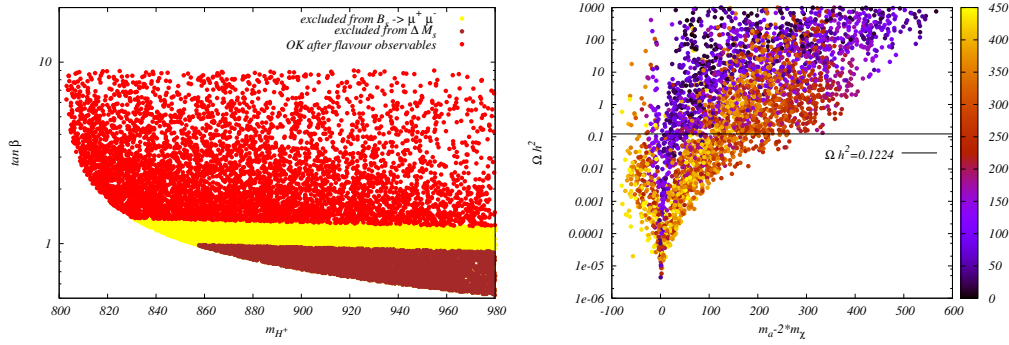


Figure 2: Left: Bounds on the $(m_{H^\pm}, \tan\beta)$ plane from B-physics observables, implemented via the SPheno [7]/ Sarah [8] interface, and compared to experimental bounds [14, 15]. The contour for low $(m_{H^\pm}, \tan\beta)$ values stems from [18, 19]. Right: Dark matter constraints in the THDMa model. Right: Dark matter relic density as a function of $m_a - 2m_\chi$, with m_χ defining the color coding. The typical resonance-enhanced relic density annihilation is clearly visible. Figures taken from [36].

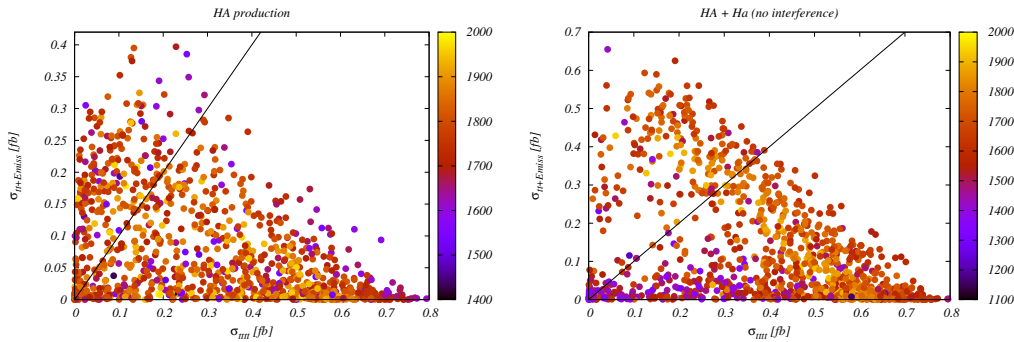


Figure 3: Production cross sections for $t\bar{t}t\bar{t}$ (x-axis) and $t\bar{t} + \cancel{E}$ (y-axis) final state in a factorized approach, for an e^+e^- collider with a 3 TeV center-of-mass energy. Left: mediated via HA, right: mediated via HA and Ha intermediate states. Color coding refers to $m_H + m_A$ (left) and $M_H + 0.5 \times (m_A + m_a)$ (right). Figures taken from [36].

the corresponding signatures, including background simulation and cut optimization, is in the line of future work.

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References

- [1] G. Altarelli and R. Barbieri, Phys. Lett. **B253**, 161 (1991).
- [2] M. E. Peskin and T. Takeuchi, Phys. Rev. Lett. **65**, 964 (1990).
- [3] M. E. Peskin and T. Takeuchi, Phys. Rev. **D46**, 381 (1992).

- [4] P. Bechtle *et al.*, Eur. Phys. J. **C80**, 1211 (2020), 2006.06007.
- [5] P. Bechtle *et al.*, Eur. Phys. J. C **81**, 145 (2021), 2012.09197.
- [6] D. Eriksson, J. Rathsmann, and O. Stål, Comput. Phys. Commun. **181**, 189 (2010), 0902.0851.
- [7] W. Porod and F. Staub, Comput. Phys. Commun. **183**, 2458 (2012), 1104.1573.
- [8] F. Staub, Comput. Phys. Commun. **185**, 1773 (2014), 1309.7223.
- [9] G. Bélanger, F. Boudjema, A. Goudelis, A. Pukhov, and B. Zaldivar, Comput. Phys. Commun. **231**, 173 (2018), 1801.03509.
- [10] G. Belanger, A. Mjallal, and A. Pukhov, Eur. Phys. J. C **81**, 239 (2021), 2003.08621.
- [11] F. Ambrogio *et al.*, Phys. Dark Univ. **24**, 100249 (2019), 1804.00044.
- [12] Gfitter Group, M. Baak *et al.*, Eur. Phys. J. **C74**, 3046 (2014), 1407.3792.
- [13] J. Haller *et al.*, Eur. Phys. J. **C78**, 675 (2018), 1803.01853.
- [14] CMS-PAS-BPH-20-003, LHCb-CONF-2020-002, ATLAS-CONF-2020-049.
- [15] HFLAV, Y. S. Amhis *et al.*, Eur. Phys. J. **C81**, 226 (2021), 1909.12524.
- [16] Planck, N. Aghanim *et al.*, Astron. Astrophys. **641**, A6 (2020), 1807.06209, [Erratum: Astron.Astrophys. 652, C4 (2021)].
- [17] XENON, E. Aprile *et al.*, Phys. Rev. Lett. **121**, 111302 (2018), 1805.12562.
- [18] M. Misiak, A. Rehman, and M. Steinhauser, JHEP **06**, 175 (2020), 2002.01548.
- [19] M. Misiak, Private communication.
- [20] J. Alwall, M. Herquet, F. Maltoni, O. Mattelaer, and T. Stelzer, JHEP **06**, 128 (2011), 1106.0522.
- [21] N. G. Deshpande and E. Ma, Phys. Rev. **D18**, 2574 (1978).
- [22] Q.-H. Cao, E. Ma, and G. Rajasekaran, Phys. Rev. **D76**, 095011 (2007), 0708.2939.
- [23] R. Barbieri, L. J. Hall, and V. S. Rychkov, Phys. Rev. **D74**, 015007 (2006), hep-ph/0603188.
- [24] A. Ilnicka, M. Krawczyk, and T. Robens, Phys. Rev. D **93**, 055026 (2016), 1508.01671.
- [25] J. Kalinowski, W. Kotlarski, T. Robens, D. Sokolowska, and A. F. Zarnecki, JHEP **12**, 081 (2018), 1809.07712.
- [26] D. Dercks and T. Robens, Eur. Phys. J. C **79**, 924 (2019), 1812.07913.
- [27] J. Kalinowski, T. Robens, D. Sokolowska, and A. F. Zarnecki, Symmetry **13**, 991 (2021), 2012.14818.
- [28] J. Kalinowski, W. Kotlarski, T. Robens, D. Sokolowska, and A. F. Zarnecki, JHEP **07**, 053 (2019), 1811.06952.
- [29] A. Ilnicka, T. Robens, and T. Stefaniak, Mod. Phys. Lett. A **33**, 1830007 (2018), 1803.03594.
- [30] S. Ipek, D. McKeen, and A. E. Nelson, Phys. Rev. **D90**, 055021 (2014), 1404.3716.
- [31] J. M. No, Phys. Rev. **D93**, 031701 (2016), 1509.01110.
- [32] D. Goncalves, P. A. N. Machado, and J. M. No, Phys. Rev. **D95**, 055027 (2017), 1611.04593.
- [33] M. Bauer, U. Haisch, and F. Kahlhoefer, JHEP **05**, 138 (2017), 1701.07427.
- [34] P. Tunney, J. M. No, and M. Fairbairn, Phys. Rev. **D96**, 095020 (2017), 1705.09670.
- [35] LHC Dark Matter Working Group, T. Abe *et al.*, Phys. Dark Univ. **27**, 100351 (2020), 1810.09420.
- [36] T. Robens, (2021), 2106.02962.