

PoS

BSM scenarios with missing energy at future lepton colliders

Tania Robens^{*a*,*}

^a Rudjer Boskovic Institute, Bijenicka cesta 54, 10000 Zagreb, Croatia

E-mail: trobens@irb.hr

I will briefly discuss the signatures and discovery prospects of several new physics models containing dark matter candidates at future lepton colliders. In particular, I will discuss the two models that, among other signatures, lead to electroweak gauge bosons and missing energy: the Inert Doublet Model, as well as the THDMa, a two Higgs doublet model with an additional pseudoscalar that serves as a portal to the dark sector. Results are mainly based on a Snowmass Whitepaper [1].

RBI-ThPhys-2022-42

41st International Conference on High Energy physics - ICHEP20226-13 July, 2022Bologna, Italy

*Speaker

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Tania Robens

1. Introduction

I briefly discuss two models that enhance the scalar sector of the Standard Model by additional particle content, including dark matter candidates. These models lead to signatures with missing energy. I present current bounds on these models as well as perspectives and rates at future lepton colliders.

2. Inert Doublet Model

The Inert Doublet Model (IDM) [2–4] is a two Higgs Doublet Model (THDM) with a discrete exact \mathbb{Z}_2 symmetry containing a dark matter candidate. The model features 7 free parameters, which we chose in the so-called physical basis [5]: $v, M_h, M_H, M_A, M_{H^{\pm}}, \lambda_2, \lambda_{345}$, where the λ s correspond to potential parameters. As two parameters (the vacuum expecation value (vev) v and $M_h \sim 125 \text{ GeV}$) are fixed by experimental measurements, we end up with a total number of 5 free parameters. Here, we consider the case where H is the dark matter candidate, which implies $M_{A, H^{\pm}} \geq M_H$.

The model is subject to a large number of theoretical and experimental constraints [1, 5–9]. These lead to a large reduction of the allowed parameter space. As an example, the masses are usually quite degenerate, as can be seen from figure 1¹. We also consider the case when $M_H \leq$



Figure 1: Masses are requested to be quite degenerate after all constraints have been taken into account. *Left:* In the $(M_A, M_{H^{\pm}})$ plane (taken from [5]). *Right:* In the $(M_{H^{\pm}} - M_H, M_A - M_H)$ plane (taken from [8]).

 $M_h/2$, where constraints from $h \rightarrow$ invisible start to play an important role and an interesting interplay arises, between bounds from signal strength measurements, and bounds from dark matter relic density, see figure 2. In [5], it was found that this in general leads to a lower bound of $M_H \sim 50 \text{ GeV}$, with exceptions presented in [9]. The discovery potential of ILC and CLIC was investigated in [11–16] for several benchmark points proposed in [8], for varying center-of-mass energies from 250 GeV up to 3 TeV. We focus on AH and H^+H^- production with $A \rightarrow ZH$ and $H^{\pm} \rightarrow W^{\pm}H$, with leptonic decays of the electroweak gauge bosons. Event generation was done using WHizard 2.2.8 [17, 18], with an interface via SARAH [19] and SPheno 4.0.3 [20, 21] for model implementation. For CLIC results energy spectra [22] were also taken into account.

¹Note that BP11 from [8] is by now excluded from the newest direct detection constraints [10].



Figure 2: Interplay of signal strength and relic density constraints in the (M_H, λ_{345}) plane. Using XENON1T results, with golden points labelling those points that produce exact relic density (taken from [6]). Note that all points displayed here also pass the new LUX-Zeppelin bounds [10].

3. THDMa

The THDMa is a type II two-Higgs-doublet model that is extended by an additional pseudoscalar *a*. In the gauge-eigenbasis, the additional scalar serves as a portal to the dark sector, with a fermionic dark matter candidate, denoted by χ . More details can e.g. be found in [23–29].

The model contains the following particles in the scalar and dark matter sector: h, H, H^{\pm} , a, A, χ . It depends on 12 additional new physics parameters

 $v, m_h, m_H, m_a, m_A, m_{H^{\pm}}, m_{\chi}; \cos(\beta - \alpha), \tan\beta, \sin\theta; y_{\chi}, \lambda_3, \lambda_{P_1}, \lambda_{P_2},$

where v and either m_h or m_H are fixed by current measurements in the electroweak sector.

I here report on results of a scan that allows all of the above novel parameters float in specific predefined ranges [29]. Two examples for direct bounds in 2-dimensional planes are displayed in figure 4. Note that for this proceeding, on contrast to the results presented in [1, 29], we have now updated the value of $B_s \rightarrow \mu^+ \mu^-$ to the current PDG value [30], we have $B_s \rightarrow \mu^+ \mu^- = (3.01 \pm 0.35) \times 10^{-9}$. Following the logic explained in [29], this leads to $(B_s \rightarrow \mu^+ \mu^-)^{Spheno} \in [1.52; 3.34] \times 10^{-9}$. Note that the ΔM_s experimental value has also been updated [31] and now reads ΔM_s (ps⁻¹) = 17.765 \pm 0.004 \pm 0.004. However, this basically leads to similar bounds as the previous value [32], so we did not update the respective bounds.

If, for consistency, now taking again a 3σ allowed range for $B_s \rightarrow \mu^+ \mu^-$, the bounds from this branching ratio and ΔM_s basically overlapp. In turn, it means that now tan β values ≤ 1 are still allowed. The second plot displays the relic density as a function of the mass difference $m_a - 2 m_{\gamma}$.

I also present cross section values for production at e^+e^- colliders for points that pass all bounds considered in [29] at a 3 TeV collider in figure 5.



Figure 3: Discovery prospects at CLIC for the IDM in $\mu^+\mu^- + \not\!\!\!E$ (*left*) and $\mu^\pm e^\mp + \not\!\!\!E$ (*right*) final states, as a function of the respective production cross-sections (*top*) and mass sum of the produced particles (*bottom*). Taken from [11].

4. Conclusion

I briefly presented two scenarios for models with dark matter candidates and their prospective signatures and rates/ discovery ranges at future lepton colliders, with a focus on larger (O (TeV)) center-of-mass energies. For the IDM, a detailed study shows that many still viable parameter points should be accessible, depending on the specifics of the particular benchmark points. For the THDMa, regions in parameter space exist where $t\bar{t} + \vec{k}$ is the dominant production mode.

References

- [1] J. Kalinowski, T. Robens, and A. F. Zarnecki, New Physics with missing energy at future lepton colliders Snowmass White Paper, in 2022 Snowmass Summer Study, 2022, 2203.07913.
- [2] N. G. Deshpande and E. Ma, Phys. Rev. D18, 2574 (1978).
- [3] Q.-H. Cao, E. Ma, and G. Rajasekaran, Phys. Rev. D76, 095011 (2007), 0708.2939.
- [4] R. Barbieri, L. J. Hall, and V. S. Rychkov, Phys. Rev. D74, 015007 (2006), hep-ph/0603188.
- [5] A. Ilnicka, M. Krawczyk, and T. Robens, Phys. Rev. D 93, 055026 (2016), 1508.01671.
- [6] A. Ilnicka, T. Robens, and T. Stefaniak, Mod. Phys. Lett. A 33, 1830007 (2018), 1803.03594.





Figure 4: *Left:* Bounds on the $(m_{H^{\pm}}, \tan \beta)$ plane from B-physics observables, implemented via the SPheno [21]/ Sarah [33] interface, and compared to experimental bounds [34, 35]. The contour for low $(m_{H^{\pm}, \tan\beta})$ values stems from [36, 37]. *Right:* Dark matter relic density as a function of $m_a - 2 m_{\chi}$, with m_{χ} defining the color coding. The typical resonance-enhanced relic density annihilation is clearly visible. Right figure taken from [29].



Figure 5: Production cross sections for $t\bar{t}t\bar{t}$ (x-axis) and $t\bar{t} + \not\!\!\!E$ (y-axis) final state in a factorized approach, for an e^+e^- collider with a 3 TeVcenter-of-mass energy. *Left:* mediated via *HA*, *right:* mediated via *HA* and *Ha* intermediate states. Color coding refers to $m_H + m_A$ (*left*) and $M_H + 0.5 \times (m_A + m_a)$ (*right*). Figures taken from [29], with an update including results from [38].

- [7] D. Dercks and T. Robens, Eur. Phys. J. C 79, 924 (2019), 1812.07913.
- [8] J. Kalinowski, W. Kotlarski, T. Robens, D. Sokolowska, and A. F. Zarnecki, JHEP 12, 081 (2018), 1809.07712.
- [9] J. Kalinowski, T. Robens, D. Sokolowska, and A. F. Zarnecki, Symmetry 13, 991 (2021), 2012.14818.
- [10] LZ, J. Aalbers et al., (2022), 2207.03764.
- [11] J. Kalinowski, W. Kotlarski, T. Robens, D. Sokolowska, and A. F. Zarnecki, JHEP 07, 053 (2019), 1811.06952.
- [12] CLIC, J. de Blas et al., 3/2018 (2018), 1812.02093.
- [13] A. F. Zarnecki et al., PoS ALPS2019, 010 (2020), 1908.04659.

- [14] A. F. Zarnecki *et al.*, Searching Inert Scalars at Future e⁺e⁻ Colliders, in *International Workshop on Future Linear Colliders*, 2020, 2002.11716.
- [15] D. Sokolowska et al., PoS EPS-HEP2019, 570 (2020), 1911.06254.
- [16] CLICdp, J. Klamka and A. F. Zarnecki, Eur. Phys. J. C 82, 738 (2022), 2201.07146.
- [17] M. Moretti, T. Ohl, and J. Reuter, p. 1981 (2001), hep-ph/0102195.
- [18] W. Kilian, T. Ohl, and J. Reuter, Eur. Phys. J. C71, 1742 (2011), 0708.4233.
- [19] F. Staub, Adv. High Energy Phys. 2015, 840780 (2015), 1503.04200.
- [20] W. Porod, Comput. Phys. Commun. 153, 275 (2003), hep-ph/0301101.
- [21] W. Porod and F. Staub, Comput. Phys. Commun. 183, 2458 (2012), 1104.1573.
- [22] L. Linssen, A. Miyamoto, M. Stanitzki, and H. Weerts, (2012), 1202.5940.
- [23] S. Ipek, D. McKeen, and A. E. Nelson, Phys. Rev. D90, 055021 (2014), 1404.3716.
- [24] J. M. No, Phys. Rev. D93, 031701 (2016), 1509.01110.
- [25] D. Goncalves, P. A. N. Machado, and J. M. No, Phys. Rev. D95, 055027 (2017), 1611.04593.
- [26] M. Bauer, U. Haisch, and F. Kahlhoefer, JHEP 05, 138 (2017), 1701.07427.
- [27] P. Tunney, J. M. No, and M. Fairbairn, Phys. Rev. D96, 095020 (2017), 1705.09670.
- [28] LHC Dark Matter Working Group, T. Abe *et al.*, Phys. Dark Univ. **27**, 100351 (2020), 1810.09420.
- [29] T. Robens, Symmetry 13, 2341 (2021), 2106.02962.
- [30] Particle Data Group, R. L. Workman et al., PTEP 2022, 083C01 (2022).
- [31] HFLAV, Y. Amhis et al., (2022), 2206.07501.
- [32] HFLAV, Y. S. Amhis et al., Eur. Phys. J. C 81, 226 (2021), 1909.12524.
- [33] F. Staub, Comput. Phys. Commun. 185, 1773 (2014), 1309.7223.
- [34] CMS-PAS-BPH-20-003, LHCb-CONF-2020-002, ATLAS-CONF-2020-049.
- [35] HFLAV, Y. S. Amhis et al., Eur. Phys. J. C81, 226 (2021), 1909.12524.
- [36] M. Misiak, A. Rehman, and M. Steinhauser, JHEP 06, 175 (2020), 2002.01548.
- [37] M. Misiak, Private communication.
- [38] CERN Report No., , 2021 (unpublished), ATLAS-CONF-2021-029.