

Searching for top squarks from the string landscape at HL-LHC

Juhi Dutta^{*a*,*}

^aHomer L. Dodge Department of Physics and Astronomy, University of Oklahoma, Norman, OK 73019, USA

E-mail: juhi.dutta@ou.edu

Supersymmetric models with low electroweak fine-tuning are more prevalent on the string landscape than fine-tuned models. We assume a fertile patch of landscape vacua containing the minimal supersymmetric standard model (MSSM) as a low-energy EFT. Such models are characterized by light higgsinos in the mass range of a few hundred GeV whilst top squarks are in the 1-2.5 TeV range. Other sparticles are generally beyond current LHC reach. We evaluate prospects for top squark searches of the expected natural SUSY at HL-LHC.

The European Physical Society Conference on High Energy Physics (EPS-HEP2023) 21-25 August 2023 Hamburg, Germany

*Speaker

[©] Copyright owned by the author(s) under the terms of the Creative Commons Attribution-NonCommercial-NoDerivatives 4.0 International License (CC BY-NC-ND 4.0).

1. Introduction

Supersymmetry is one of the best-motivated candidates for a theory encompassing physics beyond the standard model (BSM). While early estimates of naturalness placed stringent upper limits on masses of supersymmetric particles, such as $m_{\tilde{t}_1} \leq 300-400$ GeV for $\Delta_{BG} < 10-30$ where Δ_{BG} is the Barbieri-Guidice measure of naturalness [1], defined as

$$\Delta_{BG} = max_i \left| \frac{p_i \partial m_Z^2}{m_Z^2 \partial dp_i} \right|.$$

where p_i are the fundamental free parameters of the theory and subsequent measures such as Δ_{HS} required third generation squarks to be less than 500 GeV[11, 12], current experimental searches from LHC place rather stringent limits ~ O(TeV) on the lightest stops. However, the naturalness measures considered earlier turned out to be large overestimates of the actual-finetuning[13–16]. Recently, more conservative estimates of electroweak fine-tuning Δ_{EW} [6] have come up to resolve the naturalness issue where Δ_{EW} is

$$\frac{m_Z^2}{2} = \frac{m_{H_d}^2 + \Sigma_d^d - (m_{H_u}^2 + \Sigma_u^u)\tan^2\beta}{\tan^2\beta - 1} - \mu^2$$
(1)

the ratio of the largest term on the right-hand side of the eq. 1 to $\frac{m_Z^2}{2}$, where m_Z is the mass of the Z boson, $m_{H_u}^2$ and $m_{H_d}^2$ refer to the Higgs soft breaking masses coupling to the up-type and down-type quarks respectively, $\tan \beta = \frac{v_u}{v_d}$ (v_q being the vaccuum expectation value of H_q , where q = (u, d)), and Σ_q^q , refer to the loop contributions from the particles and sparticles to the Higgs sector (with dominant one-loop contributions from the lightest top squarks) and μ is the higgsino mass parameter. Such a conservative measure of naturalness imposes relatively relaxed constraints on sparticle masses which are allowed up to several TeV at little cost to finetuning since their contributions to the weak scale are suppressed by loop factors.

Another possible resolution arises from the string landscape picture. In the string landscape, where order of 10⁵⁰⁰ vacua solutions arise from compactification from 10 to 4 spacetime dimensions, each vacuum solution corresponds to a different set of 4-d low energy effective field theory law of physics. The string landscape provides a natural setting for Weinberg's anthropic solution to the cosmological constant problem [7] in an eternally inflating multiverse. In the same spirit, one tries to address the origin of the SUSY breaking scale in the string landscape. Supersymmetric models with low electroweak fine-tuning are expected to be more prevalent on the string landscape than fine-tuned models [10]. In this work, we determine the properties of the stops from the landscape and prospects of observing them at the upcoming HL-LHC.

2. MSSM from the string landscape

Assuming a fertile patch of landscape vacua containing the minimal supersymmetric standard model (MSSM) as low energy effective field theory [14], the landscape statistically favours large soft terms via a power law [3],

$$f_{SUSY} = m_{soft}^{2n_F + n_D - 1}$$



Figure 1: The probability distributions for the lightest CP-even Higgs mass m_h .



Figure 2: The probability distribution of the lightest top squark $m_{\tilde{t}_1}$ (left) and mixing angle $\cos \theta$ in the stop sector (right).

where n_F and n_D refer to the number of F-term and D-term SUSY breaking terms and where f_{SUSY} is the expected statistical distribution of landscape soft terms. A statistical pull by the landscape to large soft terms is balanced by the requirement of a derived value of the weak scale in the pocket universe (*PU*), which is not too far from its measured value in our universe (*OU*) given by the ABDS window [4]

$$\frac{m_Z^{PU2}}{2} = \frac{m_{H_d}^2 + \Sigma_d^d - (m_{H_u}^2 + \Sigma_u^u) \tan^2 \beta}{\tan^2 \beta - 1} - \mu_{PU}^2$$

such that $m_{weak}^{PU} \sim (0.5 - 5) m_{weak}^{OU}$ in order to allow for complex nuclei (and hence atoms) in our universe. The string landscape approach to soft SUSY breaking within the MSSM statistically predicts a Higgs boson with mass 125 GeV and is characterized by light higgsinos in the 100-400 GeV range, lightest top squarks are in the 1-2.5 TeV range with large trilinear soft terms which helps to push $m_h \sim 125$ GeV and other squarks beyond the HL-LHC reach. To obtain a measure of *stringy naturalness*, we implement a linear scan[5] over the NUHM2 parameter space. Fig. 1 shows the probability distribution for the lightest CP-even Higgs mass, m_h from the string landscape with a n=1 draw to large soft terms. We observe that the distribution is peaked towards ~125 GeV while $m_{\tilde{t}_1}$ has a large number of events above the TeV scale within 1-2.5 TeV with a peak ~ 1.5 TeV. The reach of LHC Run 3 and HL-LHC would probe the peak probability region in the coming years, making the search for light top-squarks of supersymmetry a highly motivated priority. Fig. 2 (right) shows the variation of the cosine of mixing angle in the stop sector vs. the lightest stop mass. It is clear from the plot, $\cos \theta \sim 0.1$ over most of the stop mass range suggesting that the lightest stop is largely dominated by the right-handed top squark. Thus, the light top-squark decays comparably via $b\tilde{\chi}_1^{\pm}$ and $t\tilde{\chi}_1^0$ yielding mixed final states of $b\bar{b} + \not E_T$, $t\bar{b}/t\bar{b} + \not E_T$ and $t\bar{t} + \not E_T$.



Figure 3: The variation of the branching ratios of \tilde{t}_1 vs. the stop mass, $m_{\tilde{t}_1}$ into the different decay modes.



Figure 4: Stop pair production cross-section at NLO at $\sqrt{s} = 14$ TeV at the HL-LHC.

3. Collider study

In this section, we discuss the collider prospects of the lightest stops at the high luminosity LHC (HL-LHC). We choose a benchmark point shown in Table 1 consistent with current experimental constraints from flavour physics and LHC data and scan over the parameter A_t to vary the mass of the stop in the range 800-2200 GeV. The production cross-section of the lightest stops is shown in Fig. 4 for $\sqrt{s} = 14$ TeV. For the collider analyses, we consider the signal final states: $b\bar{b} + \not E_T$, 1 t + 1 b + $\not\!\!E_T$ and 2 t + $\not\!\!E_T$. The dominant SM backgrounds are $t\bar{t}, b\bar{b}Z, t\bar{t}Z, t\bar{t}W, b\bar{b}W$ and single top production channel. Using boosted jet techniques to reconstruct top jets with $p_T > 400$ GeV and R = 1.5 and relying on hard kinematic observables such as $\not E_T > 400$ GeV, $H_T > 1400$ GeV, $L_T (= H_T + \not\!\!\!E_T) > 1400$ GeV, $\min(m_T(t, \not\!\!\!E_T), m_T(b, \not\!\!\!E_T)) \ge 175$ GeV, $\Delta \Phi(b, \not\!\!\!\!E_T) \ge 40^\circ$, $\Delta \Phi(J, \not\!\!\!E_T) \geq 30^\circ$ to suppress the SM background for the $tb + \not\!\!\!E_T$, we observe that the key kinematic variable to discriminate between signal and backgrounds is M_{T_2} especially at the tails of the distribution as shown in Fig. 5 for $\sqrt{s} = 14$ TeV and integrated luminosity of 3000 fb⁻¹. Similar cuts (see [5]) are imposed on the $bb + \not\!\!\!E_T$ and $tt + \not\!\!\!E_T$ channels. In all cases, the top-squark pair production is revealed as an enhancement in the m_{T_2} distribution at high values of m_{T_2} . A combined reach of all channels at HL-LHC lead to a reach of ~1.7 TeV at 5 σ as seen in Fig. 5(right) and ~2 TeV at 2σ [5].

4. Summary and Conclusions

In this work, we have investigated the properties of top squarks from the string landscape where a power-law draw to large soft terms is expected. The derived value of the weak scale must lie within the ABDS window in order to allow for complex nuclei (and hence atoms) in each

Parameters	Benchmark point
m_0	5 TeV
$m_{1/2}$	1.2 TeV
A_0	-8 TeV
$\tan \beta$	10
μ	250 GeV
m_A	2 TeV
m _ĝ	2830 GeV
$m_{\tilde{t}_1}$	1714 GeV
$m_{\tilde{t}_2}$	3915 GeV
$m_{\tilde{\chi}_1^{\pm}}$	261.7 GeV
$m_{\tilde{\chi}_2^{\pm}}$	1020.6 GeV
$m_{\tilde{\chi}_1^0}$	248.1 GeV
$m_{\tilde{\chi}^0_2}$	259.2 GeV
$m_{\tilde{\chi}^0_2}$	541.0 GeV
$m_{\tilde{\chi}^0_A}$	1033.9 GeV
m_h	124.7 GeV
$\Omega^{std}_{\tilde{Y}_1}h^2$	0.016
$\sigma^{\tilde{SI}}_{SI}(ilde{\chi}^0_1,p)$ (pb)	2.2×10^{-9}
$\sigma^{SD}(\tilde{\chi}_1^0,p)$ (pb)	2.9×10^{-5}
$\Delta_{\rm EW}$	22

Table 1: Input parameters (TeV) and masses (GeV) for the stringy natural SUSY benchmark point from the NUHM2 model with $m_t = 173.2$ GeV using Isajet 7.88.



anthropically-allowed pocket universe. Under this stringy naturalness requirement, we find $m_{\tilde{t}_1} \sim 1-2.5$ TeV with large mixing which also facilitates to lift m_h to 125 GeV while minimizing the top squark contributions to the weak scale. Despite of the large mixing, the lighter top-squark is mainly a right-squark, and lead to mixed final states of $b\bar{b} + \not{E}_T$, $t\bar{t} + \not{E}_T$ and $tb + \not{E}_T$. Using boosted jet techniques to investigate the reach of stops at HL-LHC it is possible to reach $m_{\tilde{t}_1} \simeq 1.7$ at 5σ and $\simeq 2$ TeV at 2σ therefore covering most of the stringy natural parameter space at HL-LHC.

Acknowledgments

The author acknowledges support from the HEP Dodge Family Endowment Fellowship at the Homer L.Dodge Department of Physics & Astronomy at the University of Oklahoma. The author

Juhi Dutta

thanks her collaborators H.Baer, V.Barger, D.Sengupta and K.Zhang for the successful completion of the work Ref. [5], on which this talk is based on.

References

- R. Barbieri and G. F. Giudice, Nucl. Phys. B 306 (1988), 63-76 doi:10.1016/0550-3213(88)90171-X
- [2] H. Baer, V. Barger and D. Mickelson, Phys. Rev. D 88 (2013) no.9, 095013 doi:10.1103/PhysRevD.88.095013 [arXiv:1309.2984 [hep-ph]].
- [3] M. R. Douglas, doi:10.1142/9789814412551_0012 [arXiv:1204.6626 [hep-th]].
- [4] V. Agrawal, S. M. Barr, J. F. Donoghue and D. Seckel, Phys. Rev. D 57 (1998), 5480-5492 doi:10.1103/PhysRevD.57.5480 [arXiv:hep-ph/9707380 [hep-ph]].
- [5] H. Baer, V. Barger, J. Dutta, D. Sengupta and K. Zhang, doi:10.1103/PhysRevD.108.075027
 [arXiv:2307.08067 [hep-ph]].
- [6] H. Baer, V. Barger, P. Huang, A. Mustafayev and X. Tata, Phys. Rev. Lett. 109 (2012), 161802 doi:10.1103/PhysRevLett.109.161802 [arXiv:1207.3343 [hep-ph]].
- [7] S. Weinberg, Phys. Rev. Lett. 59 (1987), 2607 doi:10.1103/PhysRevLett.59.2607
- [8] H. Baer, V. Barger, D. Martinez and S. Salam, Phys. Rev. D 108 (2023) no.3, 035050 doi:10.1103/PhysRevD.108.035050 [arXiv:2305.16125 [hep-ph]].
- [9] C. Brust, A. Katz, S. Lawrence and R. Sundrum, JHEP 03 (2012), 103 doi:10.1007/JHEP03(2012)103 [arXiv:1110.6670 [hep-ph]].
- [10] H. Baer, V. Barger, S. Salam, D. Sengupta and K. Sinha, Eur. Phys. J. ST 229 (2020) no.21, 3085-3141 doi:10.1140/epjst/e2020-000020-x [arXiv:2002.03013 [hep-ph]].
- [11] C. Brust, A. Katz, S. Lawrence and R. Sundrum, JHEP 03 (2012), 103 doi:10.1007/JHEP03(2012)103 [arXiv:1110.6670 [hep-ph]].
- [12] R. Kitano and Y. Nomura, Phys. Rev. D 73 (2006), 095004 doi:10.1103/PhysRevD.73.095004
 [arXiv:hep-ph/0602096 [hep-ph]].
- [13] H. Baer, V. Barger and D. Mickelson, Phys. Rev. D 88 (2013) no.9, 095013 doi:10.1103/PhysRevD.88.095013 [arXiv:1309.2984 [hep-ph]].
- [14] H. Baer, V. Barger, D. Martinez and S. Salam, Phys. Rev. D 108 (2023) no.3, 035050 doi:10.1103/PhysRevD.108.035050 [arXiv:2305.16125 [hep-ph]].
- [15] A. Mustafayev and X. Tata, Indian J. Phys. 88 (2014), 991-1004 doi:10.1007/s12648-014-0504-8 [arXiv:1404.1386 [hep-ph]].
- [16] H. Baer, V. Barger, D. Mickelson and M. Padeffke-Kirkland, Phys. Rev. D 89 (2014) no.11, 115019 doi:10.1103/PhysRevD.89.115019 [arXiv:1404.2277 [hep-ph]].