

The lattice extraction of the TMD soft function using the auxiliary field representation of the Wilson line

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The transverse momentum dependent (TMD) soft function can be obtained by formulating the Wilson line in terms of auxiliary 1-dimensional fermion fields on the lattice. In this formulation, the directional vector of the auxiliary field in Euclidean space has the form $\vec{n} = (in^0, \vec{0}_\perp, n^3)$, where the time component is purely imaginary. The components of these complex directional vectors in the Euclidean space can be mapped directly to the rapidities of the Minkowski space soft function. We present the results of the one-loop calculation of the Euclidean space analog to the soft function using these complex directional vectors. As a result, we show that the calculation is valid only when the directional vectors obey the relation: $|r| = |n^3/n^0| > 1$, and that this result corresponds to a computation in Minkowski space with space-like directed Wilson lines. Finally, we show that a lattice calculable object can be constructed that has the desired properties of the soft function.

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1. Introduction

Formulating the soft function on the lattice is a critical step in the investigation of TMD physics from lattice computations. The rapidity divergences present in computations of TMD objects, including the soft function, lead to unique challenges on the lattice, since the analog to rapidity in Euclidean space involves a Wick rotation. In the case of the soft function, we propose to perform the lattice computation of the soft function with Wilson lines whose directional vectors take a purely imaginary time component, i.e. $\vec{n} = (in^0, \vec{0}_\perp, n^3)$.

Using this formulation of the Wilson line would allow for a direct analytic continuation of the Euclidean soft function to its counterpart in Minkowski space in the space-like regime, since the soft function is time-independent. We implement the Wilson line on the lattice through the auxiliary field representation, whose equation of motion can be solved iteratively in Euclidean time. There are further subtleties and complications related to the lattice realization of an auxiliary field propagator that will be discussed later in this paper.

We will demonstrate that there is a direct connection between this complex Wilson line direction on the Euclidean lattice and the Minkowski space rapidity defined in Collins regularization scheme [1]. In other words, our lattice computation corresponds to the Minkowski space soft function defined with space-like directed Wilson lines.

When the Wilson line is pointing in a time-like direction, its auxiliary field representation corresponds to heavy quark effective theory (HQET). In this case, the directional vector of the Wilson line corresponds to the heavy quark velocity. It was proposed in [2] to calculate the soft function using HQET, modeling the soft function as the form factor of a heavy quark pair. In this formulation, the heavy quark velocity corresponds to the time-like directed Wilson lines in Minkowski space. Our work is motivated by this proposal, but differs in our use of Euclidean directional vectors that map to space-like directions in Minkowski space. Furthermore, we find that a Euclidean space computation of the soft function with directional vectors corresponding to time-like directions in Minkowski space gives a divergent integral.

This paper will proceed as follows. First, we will provide a brief review of TMD factorization in order to put our work into context. Then, we will go into more detail on the theoretical motivation, including a perturbative analysis of the Euclidean space soft function computed using the complex Wilson lines. After that, we will discuss the strategy for numerical implementation of this method.

2. Review of TMD factorization

A comprehensive review of TMD factorization may be found in [3], here we restrict ourselves to Drell-Yan scattering.

The Drell-Yan process involves the collision of two protons and produces a lepton pair in the final state: $pp \rightarrow \gamma^*/Z \rightarrow \ell^+\ell^-$. Here, measured final state variables are: Q , the invariant mass,

Y , the rapidity, and \vec{q}_\perp , the transverse momentum. The cross section may be factorized as:

$$\frac{d\sigma}{dQdYd^2\vec{q}_\perp} = \sum_{i,j} H_{ij}(Q, \mu) \int d^2\vec{b}_\perp e^{i\vec{b}_\perp \cdot \vec{q}_\perp} B_i \left(x_a, \vec{b}_\perp, \mu, \frac{\zeta_a}{\nu^2} \right) B_j \left(x_b, \vec{b}_\perp, \mu, \frac{\zeta_b}{\nu^2} \right) S_{ij}(b_\perp, \mu, \nu) \times \left[1 + \mathcal{O} \left(\frac{q_\perp^2}{Q^2}, \frac{\Lambda_{\text{QCD}}^2}{Q^2} \right) \right], \quad (1)$$

where the indices i, j represent the quark flavor or gluon. The soft function is quark flavor independent, so one can write $S_{ij} \equiv S_i$. Here, μ is the renormalization scale, and $\zeta_{a,b}$ is the Collins-Soper (CS) scale. The rapidity scale, ν cancels between the beam and soft function, and $\zeta_a \zeta_b = Q^4$. The variables $x_{a,b}$ are the Bjorken- x variables associated with the momentum fraction of the relevant parton inside of the proton. The transverse separation \vec{b}_\perp is the coordinate space conjugate to \vec{q}_\perp . Equation (1) is written in terms of the renormalized beam and soft functions, B_i and S_i , which can be obtained from the bare functions through:

$$B_i \left(x, \vec{b}_\perp, \mu, \zeta/\nu^2 \right) = \lim_{\epsilon \rightarrow 0} \lim_{\tau \rightarrow 0} Z_B^i(b_\perp, \mu, \nu, \epsilon, \tau, xP^+) \frac{B_i^{0(u)} \left(x, \vec{b}_\perp, \epsilon, \tau, xP^+ \right)}{S_i^{0(\text{subt})}(b_\perp, \epsilon, \tau)} \quad (2)$$

$$S_i(b_\perp, \mu, \nu) = \lim_{\epsilon \rightarrow 0} \lim_{\tau \rightarrow 0} Z_S^i(b_\perp, \mu, \nu, \epsilon, \tau) S_i^0(b_\perp, \epsilon, \tau), \quad (3)$$

where Z_B^i and Z_S^i are the UV renormalization factors, and ϵ is a parameter associated with dimensional regularization. The term in the denominator of Eq. (2), $S_i^{0(\text{subt})}$, is introduced to account for the overlap of infrared (IR) regions between the beam and soft function. As such, it takes care of possible double counting.

The beam functions capture long-distance physics associated with the incoming protons, while the soft function accounts for soft gluon radiation in the final state. It is not possible to directly measure the soft function experimentally, and so it must be calculated via non-perturbative methods.

A central aspect of TMD factorization are the rapidity divergences associated with partons moving in the opposite direction of their parent hadron with infinite rapidity. Because rapidity divergences are not associated with IR physics nor ultraviolet (UV) divergences, special care must be taken in regulating them. Above, τ represents a generic rapidity regulator, and ν the associated scale. There are a number of possible regulators to choose from, including the one introduced by Collins [1]. Being an unphysical divergence, the dependence on the rapidity scale necessarily cancels between the two beam functions and the soft function.

Choosing the Collins scheme in place of the generic rapidity regulator, τ , the beam and soft functions are defined as:

$$B_i^{0(u)} \left(x, \vec{b}_\perp, \epsilon, y_B, xP^+ \right) = \int \frac{db^-}{2\pi} e^{-ib^-(xP^+)} \langle P | \bar{\psi}_i^0 \left(b^-, \vec{b}_\perp \right) W_{n_B} \left(b^-, \vec{b}_\perp; -\infty, 0 \right) \times W_\perp \left(-\infty n_B; 0, b_\perp \right) W_{n_B} \left(0; 0, -\infty \right) \frac{\gamma^+}{2} \psi_i^0 \left(0 \right) | P \rangle, \quad (4)$$

$$S^0 \left(b_\perp, \epsilon, y_A - y_B \right) = \frac{1}{N_C} \langle 0 | W_{n_A} \left(b_\perp; 0, -\infty \right) W_{n_B} \left(b_\perp; -\infty, 0 \right) W_\perp \left(-\infty n_B; 0, b_\perp \right) W_{n_B} \left(0; 0, -\infty \right) W_{n_A} \left(0; -\infty, 0 \right) W_\perp \left(-\infty n_A; b_\perp, 0 \right) | 0 \rangle, \quad (5)$$

where

$$W_n(x; a, b) = P \exp \left\{ -ig_0 \int_a^b ds n^\mu A_\mu^{c0}(x + sn) t^c \right\}. \quad (6)$$

The beam and soft functions are naively defined with Wilson lines pointing in light-like directions, defined by directional vectors $n = (1, 0, 0, 1)$, and $\bar{n} = (1, 0, 0, -1)$. In the Collins scheme, however, the Wilson lines now point in space-like directions, and we make the replacements:

$$n \rightarrow n_A = n - e^{-2y_A} \bar{n}, \quad \bar{n} \rightarrow n_B = \bar{n} - e^{2y_B} n, \quad (7)$$

where y_A and y_B are the rapidities of the Wilson lines, and the light-cone direction is recovered in the limits $y_A \rightarrow \infty$ and $y_B \rightarrow -\infty$. The TMDPDF can then be constructed from a combination of the beam and soft function:

$$f_i(x, \vec{b}_\perp, \mu, \zeta) = \lim_{\epsilon \rightarrow 0} Z_{UV}^i(\mu, \epsilon, \zeta) \times \lim_{\substack{y_A \rightarrow +\infty \\ y_B \rightarrow -\infty}} B_i(x, \vec{b}_\perp, \epsilon, y_B, xP^+) \sqrt{\frac{S_i(b_\perp, \epsilon, y_A - y_n)}{S_i(b_\perp, \epsilon, y_A - y_B) S_i(b_\perp, \epsilon, y_n - y_B)}}. \quad (8)$$

where $\zeta = 2(xP^+)^2 e^{-2y_n}$ is the CS scale [1] for the n -collinear proton.

The TMDPDF, along with the beam and soft functions, follows a set of evolution equations in the renormalization and CS scale. The equation relevant to our discussion is:

$$\frac{\partial \ln f_q(x, \vec{b}_\perp, \mu, \zeta)}{\partial \ln \sqrt{\zeta}} = \gamma_\zeta^q(\mu, b_\perp). \quad (9)$$

where we have chosen $i = q$ for the quark TMDPDF. The CS kernel, $\gamma_\zeta^q(\mu, b_\perp)$, governs the evolution of the TMDPDF with respect to the change of the rapidity scale, ζ . It can be obtained directly from the beam or the soft function. For instance,

$$\gamma_\zeta^q(\mu, b_\perp) = \frac{d \ln S_q(b_\perp, \epsilon, y)}{dy} - \text{UV counterterms}, \quad (10)$$

and so, by putting this formalism on the lattice, it is possible to determine the Collins-Soper kernel as well as the soft function from one lattice calculation.

The soft function and the CS kernel are also necessary components in the calculation of TMDPDFs on the lattice. Currently, there is a proposal for the extraction of TMDPDFs using the quasi-PDF approach ([2, 4–7]). The matching relation connecting the quasi-TMDPDF, \tilde{f}_q , to the light-cone TMDPDF in the case of unpolarized quarks is:

$$\tilde{f}_q(x, \vec{b}_\perp, \mu, \tilde{\zeta}, x\tilde{P}^z) = C_q(x\tilde{P}^z, \mu) \exp \left[\frac{1}{2} \gamma_\zeta^q(\mu, b_\perp) \log \frac{\tilde{\zeta}}{\zeta} \right] f_q(x, \vec{b}_\perp, \mu, \zeta) \times \left\{ 1 + O \left(\frac{1}{(x\tilde{P}^z b_\perp)^2}, \frac{\Lambda_{\text{QCD}}^2}{(x\tilde{P}^z)^2} \right) \right\}, \quad (11)$$

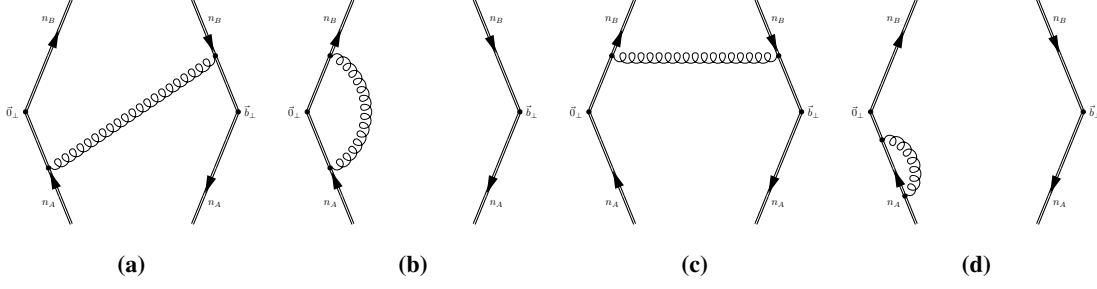


Figure 1: Diagrams contributing at one-loop to the soft function

where C_q is a perturbatively calculable matching kernel, and $\tilde{\zeta} = x^2 m_h^2 e^{2(y_{\bar{P}} + y_B - y_n)}$. The quasi-TMDPDF is constructed as [7]:

$$\tilde{f}_q(x, \vec{b}_\perp, \mu, \tilde{\zeta}, x\tilde{P}^z) = \tilde{f}_q^{\text{naive}}(x, \vec{b}_\perp, \mu, \tilde{\zeta}, x\tilde{P}^z) \sqrt{\frac{\tilde{S}_q^{\text{naive}}(b_\perp, \mu)}{S_q(b_\perp, \mu, 2y_n, 2y_B)}}, \quad (12)$$

where $\tilde{f}_q^{\text{naive}}$ and $\tilde{S}_q^{\text{naive}}$ are lattice calculable objects, and S_q is the Collins soft function. The lattice extracted Collins soft function and CS kernel can be used in obtaining f_q from Eqs. (11) and (12).

An extraction of the TMDPDF was performed by the Lattice Parton Collaboration (LPC) [8] using a different method [2] to obtain the soft function.

3. Theoretical considerations

We perform our perturbative analysis in the case of infinite and finite length Wilson lines. For infinite Wilson lines we demonstrate an exact correspondence between the Euclidean space and Minkowski space computation at one loop. Finally, we show that the finite length Wilson line computation approaches the infinite length result after constructing a ratio to cancel out the finite length effects.

3.1 Infinite Wilson lines

We first perform a one-loop computation of the soft function in Euclidean space with complex directions in order to make a connection with the Minkowski space result:

$$\tilde{n}_A = (in_A^0, \vec{0}_\perp, n_A^3), \quad \tilde{n}_B = (in_B^0, \vec{0}_\perp, -n_B^3) \quad (13)$$

For convenience we define the ratios: $r_a = n_A^3/n_A^0$, and $r_b = n_B^3/n_B^0$. We perform the computation in d space-time dimensions in order to regulate UV divergences. Before presenting the full result, we can gain important insights by first looking at diagram a in figure 1:

$$\begin{aligned} S_{1a}(b_T, \epsilon, r_a, r_b) &= ig^2 C_F(\tilde{n}_{A,\mu} \tilde{n}_{B,\nu}) \int_{-\infty}^0 ds \int_{-\infty}^0 dt A^\mu(b + s\tilde{n}_A) A^\nu(t\tilde{n}_B) \\ &= ig^2 C_F(\tilde{n}_A \cdot \tilde{n}_B) \int_{-\infty}^0 ds \int_{-\infty}^0 dt \int \frac{d^d k}{(2\pi)^d} e^{-ik(b+s\tilde{n}_A-t\tilde{n}_B)} \frac{-i}{k^2}, \end{aligned} \quad (14)$$

where on line 2 of Eq. (14), we wrote the gluon propagator in Euclidean space. From here, one usually proceeds with the computation in momentum space by first integrating over s and t to obtain the Wilson line propagator. However, these integrals are not well defined, since there is a finite real part in the exponential. Looking at the s integration:

$$\int_{-\infty}^0 ds e^{sn_A^0 k_4 - isn_A^3 k_3} = \frac{e^{sn_A^0 k_4 - isn_A^3 k_3}}{n_A^0 k_4 - in_A^3 k_3} \Bigg|_{s=-\infty}^0 \xrightarrow{k_4 < 0} \infty. \quad (15)$$

Without a clear way to handle this integration, we instead approach the computation by performing the momentum integral in Eq. (14) first. In order to do this, we apply the Schwinger parametrization, and complete the square in k :

$$\begin{aligned} S_{1a}(b_T, \epsilon, r_a, r_b) &= g^2 C_F (\tilde{n}_A \cdot \tilde{n}_B) \int_{-\infty}^0 ds \int_{-\infty}^0 dt \int_0^\infty du \int \frac{d^d k}{(2\pi)^d} e^{-uk^2} e^{-(b+s\tilde{n}_A-t\tilde{n}_B)^2/4u} \\ &= \frac{g^2 C_F}{4\pi^{2-\epsilon}} \frac{(b_T^2)^\epsilon}{2\epsilon_{\mathbb{R}}} \frac{\Gamma(1-\epsilon)}{2} \log \left(\frac{(r_a-1)(r_b-1)}{(r_a+1)(r_b+1)} \right) \frac{r_a r_b + 1}{(r_a + r_b)}. \end{aligned} \quad (16)$$

Thus, we obtain a finite result for diagram 1a. However, this result is only valid when $|r_a|, |r_b| > 1$. This can be seen from expanding the term in the exponential of line 1 in Eq. (16):

$$(b + s\tilde{n}_A - t\tilde{n}_B)^2 = b_\perp^2 + s^2(n_A^0)^2 (r_a^2 - 1) + t^2(n_B^0)^2 (r_b^2 - 1) + 2stn_A^0 n_B^0 (r_a r_b + 1) > 0. \quad (17)$$

Equation (17) must be positive in order for the integral in Eq. (16) to converge, which is only true when the absolute values of r_a and r_b are greater than one.

The full one-loop result in Euclidean space with complex directional vectors is:

$$S(b_T, \epsilon, r_a, r_b) = 1 + \frac{\alpha_s C_F}{2\pi} \left(\frac{1}{\epsilon} + \log(\pi b_\perp^2 \mu_0^2 e^{\gamma_E}) \right) \left\{ 2 + \log \left(\frac{(r_a-1)(r_b-1)}{(r_a+1)(r_b+1)} \right) \frac{r_a r_b + 1}{r_a + r_b} \right\} \quad (18)$$

Collins scheme uses space-like Wilson lines defined in Eq. (7), but in principle one can also use time-like Wilson lines to regulate the rapidity divergence. For completeness, we list both space-like and time-like cases here:

$$\text{Time-like: } n_A = \left(1 + e^{-2y_A}, \vec{0}_\perp, 1 - e^{-2y_A} \right), \quad n_B = \left(1 + e^{2y_B}, \vec{0}_\perp, -1 + e^{2y_B} \right) \quad (19)$$

$$\text{Space-like: } n_A = \left(1 - e^{-2y_A}, \vec{0}_\perp, 1 + e^{-2y_A} \right), \quad n_B = \left(1 - e^{2y_B}, \vec{0}_\perp, -1 - e^{2y_B} \right) \quad (20)$$

Because the complex directional vectors in Euclidean space are written in terms of the elements of their Minkowski space counterparts, we write r_a and r_b in terms of the components of the space-like or time-like directional vectors. For the time-like case, we get:

$$r_a = \frac{1 - e^{-2y_A}}{1 + e^{-2y_A}}, \quad r_b = \frac{1 - e^{2y_B}}{1 + e^{2y_B}}, \quad (21)$$

where we can see that $-1 < r_a, r_b < 1$. We immediately find that this fails to satisfy the condition set by Eq. (17). As for the space-like case, we find:

$$r_a = \frac{1 + e^{-2y_A}}{1 - e^{-2y_A}}, \quad r_b = \frac{1 + e^{2y_B}}{1 - e^{2y_B}}, \quad (22)$$

which satisfies Eq. (17). The third condition set by Eq. (17) demands that $n_A^0 n_B^0 (r_a r_b + 1) > 0$, which indicates that the Wilson lines must be both future pointing or both past pointing, corresponding to $e^+ e^-$ annihilation or DY type processes.

Substituting Eq. (22) into Eq. (18), we recover the Minkowski space result:

$$S(b_\perp, \epsilon, y_A, y_B) = 1 + \frac{\alpha_s C_F}{2\pi} \left(\frac{1}{\epsilon} + \ln \left(\pi b_\perp^2 \mu_0^2 e^{\gamma_E} \right) \right) \left\{ 2 - 2|y_A - y_B| \frac{1 + e^{2(y_B - y_A)}}{1 - e^{2(y_B - y_A)}} \right\} + \mathcal{O}(\alpha_s^2) \quad (23)$$

3.2 Finite Wilson lines

The calculation on the lattice is limited by its finite space-time volume, and Wilson lines therefore have a finite length L . To take this into account we also perform a one-loop calculation of finite length Wilson lines. One of the consequences of finite length Wilson lines is the presence of a linear divergence in L . We can remove the linear divergence by constructing the ratio ([2, 9]):

$$S(b_\perp, a, r_a, r_b) = \lim_{L \rightarrow \infty} \frac{S(b_\perp, a, r_a, r_b, L)}{\sqrt{S(b_\perp, a, r_a, -r_a, L) S(b_\perp, a, -r_b, r_b, L)}}, \quad (24)$$

where a is some generic regulator used for the UV divergence. We have found that Eq. (24) holds to one loop in perturbation theory, using a Polyakov regulator for the UV. Additionally, we have found that power corrections of order b_\perp^2/L^2 should also cancel in the ratio. Computing the ratio in Eq. 24 on the lattice at large enough L should produce a time-independent observable up to power suppressed corrections in b_\perp^4/L^4 , since L on the lattice will be proportional to Euclidean time.

4. Strategy of numerical implementation

It is well known ([10, 11]) that the Wilson line can be written in terms of a one-dimensional auxiliary fermion field that ‘lives’ along the path of the Wilson line:

$$P \exp \left\{ -ig \int_{s_i}^{s_f} ds n^\mu A_\mu(y(s)) \right\} = Z_\psi^{-1} \int \mathcal{D}\psi \mathcal{D}\bar{\psi} \psi \bar{\psi} \exp \left\{ i \int_{s_i}^{s_f} ds \bar{\psi} i n \cdot \partial \psi - g_0 \bar{\psi} n \cdot A \psi \right\}, \quad (25)$$

whose propagator solves the Green function:

$$i n \cdot D H_n(x - y) = i \delta^{(4)}(x - y), \quad (26)$$

with $D = \partial + ig_0 A$ the covariant derivative. The Euclidean space analog of Eq. (26) has a directional vector with a purely imaginary time component, and can be written as $\tilde{n} = (i n^0, \vec{n})$ in terms of the components of the Minkowski space vector. After performing a Wick rotation, the Euclidean space Green function is:

$$i \tilde{n} \cdot D_E H_{\tilde{n}}(x_E - y_E) = \delta^{(4)}(x_E - y_E). \quad (27)$$

As argued in [12, 13], meaningful solutions for Eq. (27) can only be obtained with a UV cutoff. Therefore, with the lattice spacing serving as our regulator, a discretized version of this propagator

can be constructed. The central idea then is to construct the soft function in terms of the lattice regulated versions of these auxiliary field propagators, $H_{\tilde{n}}$, which can then be shown to have a well-defined continuum limit.

Solving the auxiliary field propagator on the lattice proceeds in the exact same way as in HQET ([12–15]), but now we are no longer bound by the kinematical restrictions on heavy quarks. Therefore, we use the recursive relation given in Eq. (31) of [15] to construct the lattice auxiliary field propagators. Based on our perturbative analysis, we expect that computing Eq. (24) on the lattice will approach a time independent result at large enough Euclidean time.

5. Conclusions

Extracting the soft function through the auxiliary field method shows promise based on our perturbative analysis, given the time independence of the soft function, and the direct relationship between the complex directional vectors, \tilde{n}_A, \tilde{n}_B , in Euclidean space, and y_A, y_B in Minkowski space. An important point that we did not discuss here is that a direct analytic continuation of our lattice computation to Minkowski space may be complicated by the pole structure of the auxiliary field propagator. However, an analysis of this problem was already done in the context of moving-HQET in [13]. Finally, the construction of the ratio in Eq. (24) solves the issues related to the linear divergence in L and removes power corrections up to, at least, b_{\perp}^2/L^2 . We have started the numerical implementation of this formalism on the lattice. In the near future we will present first results with the aim of establishing and demonstrating the feasibility of the presented ideas. ¹

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¹After this proceeding was published, we were made aware of a lecture note [16] which implies that one can analytically continue the heavy-quark form factor [2] in the z -axis to obtain Euclidean Wilson lines along the $(n^0, 0, 0, in^3)$ direction. This would be equivalent to our proposal through $O(4)$ symmetry [16] if $n^3 < n^0$. Nevertheless, such an analytical continuation was not explicitly pointed out in Refs. [2, 16], nor was there mention of the failure of the standard timelike moving HQET in calculating the soft function, which was discovered in this work.

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