

# **DUNE's low energy physics searches**

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The Deep Underground Neutrino Experiment (DUNE) is a long-baseline neutrino experiment that will precisely measure neutrino oscillation parameters, observe astrophysical neutrinos, and search for processes beyond the Standard Model such as nucleon decays, heavy neutral leptons, and dark matter. DUNE will build four Liquid Argon Time Projection Chambers (LAr-TPC), as far detectors, with a total mass of ~ 70 kt LAr located at Sanford Underground Research Facility (SURF), 1.5 km underground. A near-site complex, hosting different detectors, will measure the neutrino flux from an accelerated particle beam (1.2 MW) produced at the Long Baseline Neutrino Facility (LBNF) at Fermilab, 1300 km away from SURF. Additionally, this few-MeV low energy regime is of particular interest for detecting the burst of neutrinos from a galactic core-collapse supernova and solar neutrinos. DUNE will measure oscillation parameters from the solar neutrino flux analysis with much higher precision than previous experiments and discover the yet-unobserved hep flux component.

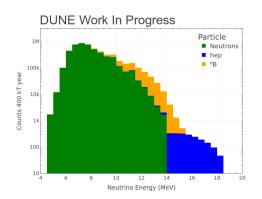
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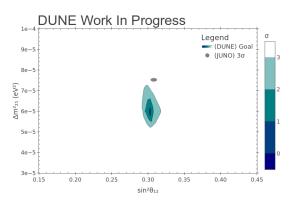
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## 1. Low Energy Interactions

the dominant neutrino interaction channel in LAr is the charged-current (CC) absorption of an electron neutrino ( $\nu_e$ ) on  $^{40}$ Ar, with a Q-value of 1.5 MeV and cross-section ranging from  $1 \times 10^{-43}$  cm<sup>2</sup> to  $1 \times 10^{-38}$  cm<sup>2</sup>. The observable topology is a short electron track surrounded by deexcitation gammas ( $\gamma$ ), which appear as TPC blips, up to 1 m apart from the de-excitation of the resulting  $^{40}$ K\* nucleus. Less frequent, but equally important, is the elastic scattering (ES) of neutrinos on electrons, characterized by tracks and the absence of gammas, that conserve some directional information of the incoming particle. This channel is fundamental to performing pointing studies and can provide a flavor-independent flux measurement. According to current simulation standards, DUNE's reconstruction threshold corresponds to  $\sim 1$  MeV (but could potentially be lower), with an energy resolution between 10 and 20% in the few MeV range, allowing the experiment to access a full range of low energy searches. In the following, we focus on two prominent examples.

#### 1.1 Solar Neutrinos





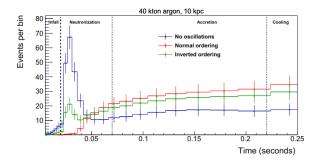
(a) Reconstructed solar neutrino energy spectrum.

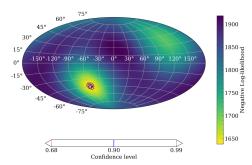
**(b)** DUNE's sensitivity contours to  $\Delta m_{21}^2$  and  $\sin^2 \theta_{12}$ .

Our Sun produces a continuous flux of  $v_e$  as a by-product of fusion reactions in its core. DUNE has the potential to characterize and study the oscillation of the two most energetic solar neutrino chains (see Figure 1a) with reconstructed energies centered at 10 MeV and an expected interaction rate of  $\sim 10\,\mathrm{mHz}$ . The measurement of  $^8\mathbf{B}\ \nu_\mathrm{e}$  flux is motivated by the existing tension (recently reduced by the latest Super–Kamiokande results) between the preferred  $\Delta m_{21}^2$  value measured by reactor- and solar-neutrino experiments [1]. A background model has been implemented accounting for intrinsic LAr isotopes and contaminants, detector components, and even external radiation from SURF's cavern. Within this realistic model, neutron captures from the cavern are expected to be the measurement's main background due to a similar topology to CC solar  $v_e$  events and a deep penetration length avoiding rejection by fiducialization. The highest and irreducible capture energy (released as a  $\gamma$  cascade from the nucleus) corresponds to  $\sim 9 \,\mathrm{MeV} \,(\mathrm{n} + {}^{36}\mathrm{Ar} \to \gamma^{\mathrm{s}})$  which could be reconstructed up to 14 MeV fake neutrino signals. Assuming DUNE's exposure of 400 kt yr, current studies estimate a Signal-to-Background of 1.19 with > 400k events above 10 MeV (~ 2k events above 14 MeV in the **hep** region) and predicts sensitivities to  $\Delta m_{21}^2$  with  $> 3\sigma$  separation to projected results from the JUNO experiment (see Figure 1b). Studies including the reconstruction of the drift-coordinate by matching the charge and light signals are ongoing.

### 1.2 Supernova Neutrino Detection and Pointing

A supernova releases an intense source of neutrinos of all flavors (99% of the of the released energy), which are ideal astrophysical messengers due to their prompt and direct travel upon release. DUNE can probe the inner evolution of the core-collapse mechanism by studying the neutrino's arrival time and energy spectra [2]. Within DUNE, three qualitative stages of the collapse can be distinguished (see Figure 2a). Preliminary studies show good sensitivity to the entire Milky Way, and possibly beyond, with thousands of interacting neutrinos (model–dependent). By measuring the spectral profiles and non–thermal features, DUNE will be sensitive to a broad range of supernova and neutrino physics phenomena, including neutrino mass ordering and collective effects.





(a) Time spectrum of a neutrino burst from a core-collapse supernova up to 10 kpc.

**(b)** Reconstructed pointing example from a simulated core–collapse supernova burst

By reconstructing the ES neutrino events during a supernova burst, its direction can be inferred. To minimize the head-tail ambiguities in track reconstruction, a new algorithm called "brems-flipping" has been developed [3]. So far, pointing resolution has been estimated under different reconstruction scenarios. **Perfect:** (ideal classification of ES events) resolution is 3.4 degrees (6.6 degrees) with 68% coverage for supernovae at a distance of 10 kpc and an effective fiducial volume of 40 kt (10 kt). **Optimistic** (incorrect classification of 4%  $\nu_e$  CC events as ES): resolution is 4.3 degrees (8.7 degrees). Accurate and fast determination is critical for multi-messenger astronomy, and efforts are ongoing to make pointing with DUNE reliable and efficient. Due to the presence of mainly CC events, machine–learning techniques will need to be implemented to reject non–ES tracks from contributing to the algorithm.

#### References

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- [2] B. Abi, R. Acciarri, M. Acero, G. Adamov, D. Adams, M. Adinolfi, Z. Ahmad, J. Ahmed, T. Alion, S. Monsalve, *et al.*, "Supernova neutrino burst detection with the deep underground neutrino experiment," *The European Physical Journal C*, vol. 81, no. 5, p. 423, 2021.
- [3] A. Abud, B. Abi, R. Acciarri, M. Acero, M. Adames, G. Adamov, M. Adamowski, D. Adams, M. Adinolfi, C. Adriano, et al., "Supernova pointing capabilities of dune," arXiv preprint arXiv:2407.10339, 2024.